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Gravitational Tensor-Monopole Moment of Hydrogen Atom To Order $\mathcal{O}(\alpha)$

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We calculate the gravitational tensor-monopole moment of the momentum-current density T^{ij} in the ground state of the hydrogen atom to order $\mathcal{O}(\alpha)$ in quantum electrodynamics (QED). The result is

 $\tau_H/\tau_0 - 1 = \frac{4\alpha}{3\pi} (\ln \alpha - 0.028)$

where $\tau_0 = \hbar^2/4m_e$ is the leading-order moment. The physics of the next-to-leading-order correction is similar to that of the famous Lamb shift for energy levels.

- Intriguing: a new atomic observable
- Is the physics really similar to the Lamb shift?
- Practically important: related to hadronic structure
- Experiments ongoing at JLab and planned in Electron-Ion Collider

Matrix elements of the energy-momentum tensor in the hydrogen atom

based on work with Yizhuang Liu, Jagiellonian University, Poland

Outline:

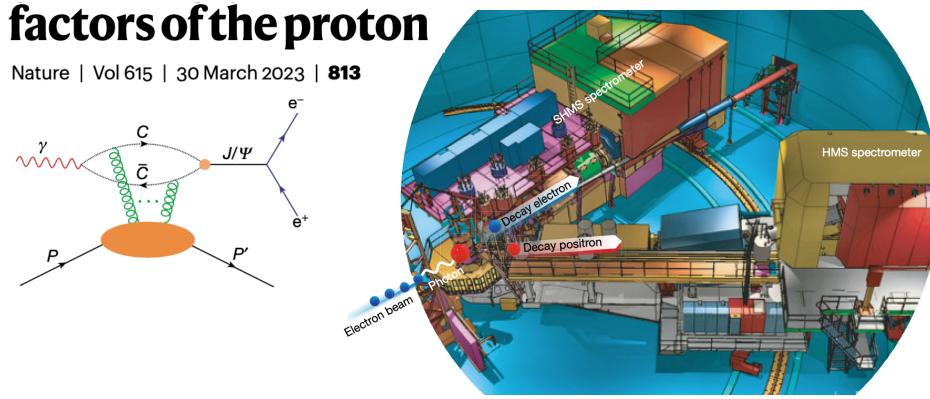
- Experimental relevance: hadronic physics (a new observable)
- Gravitational form-factors: the D-term
- D-term's sign vs stability of the system
- Logarithmic corrections: D-term vs Lamb shift

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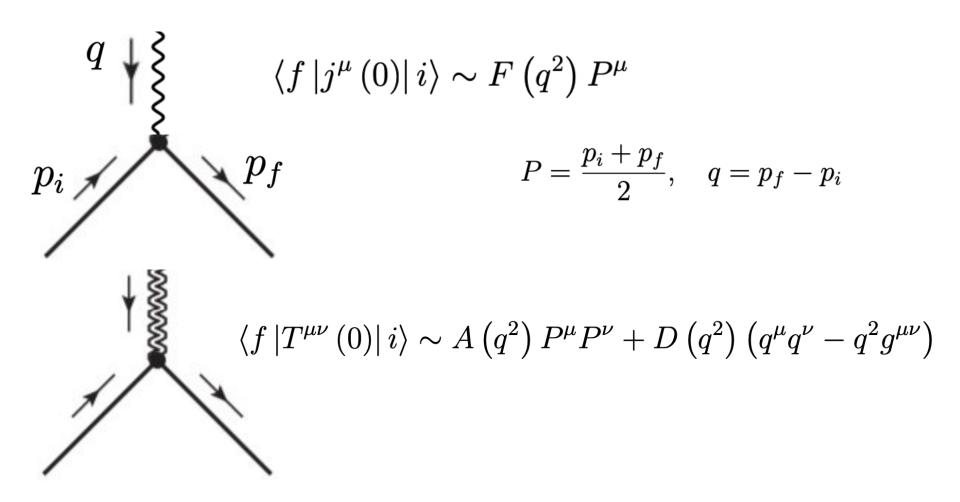
Experimental consequences

Determining the gluonic gravitational form

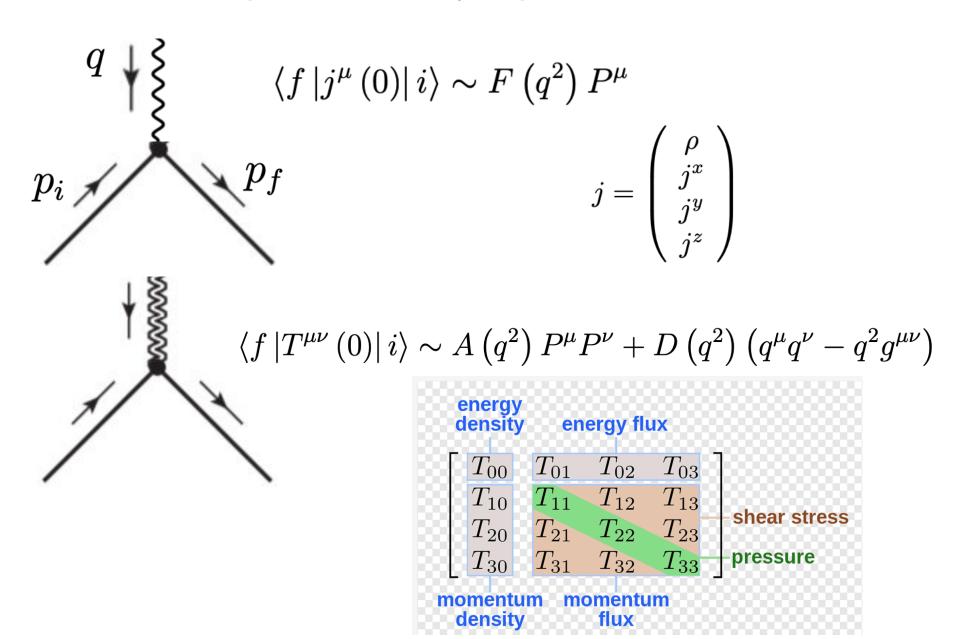


- Beyond earlier studies of the charge and spin distributions in the proton;
- New parameter: proton mass radius 0.52(3) fm.

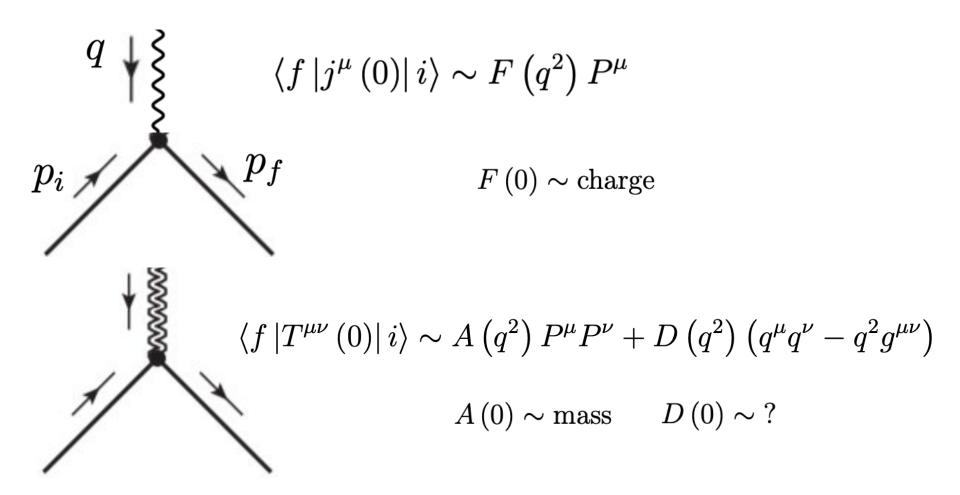
Photon vs. graviton couplings



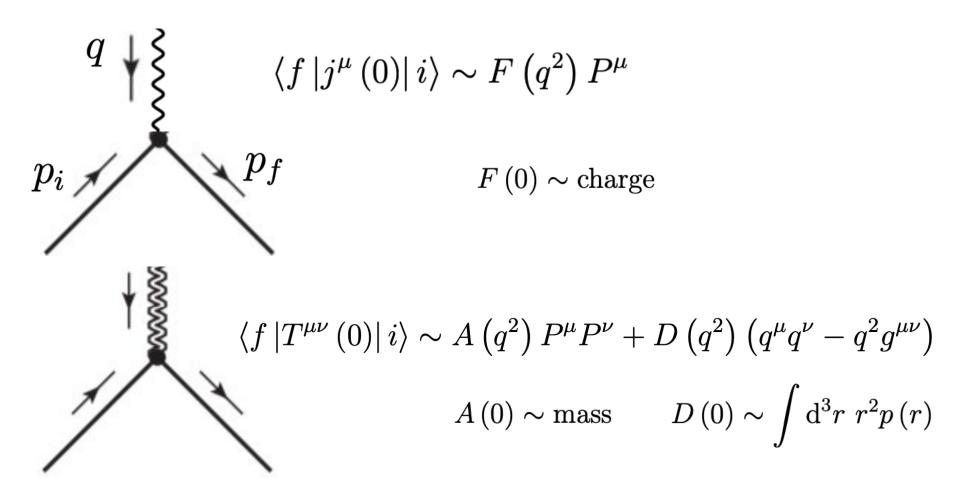
Photon vs. graviton couplings



Interpretation of form factors at $q^2=0$



Interpretation of form factors at $q^2=0$



Two integrals: moments of pressure distribution

$$0=\int p(r){
m d}^3r$$
 Max von Laue's stability condition $D\sim\int p(r)r^2{
m d}^3r$ assuming spherical symmetry

Max von Laue's stability condition

Energy-momentum conservation \rightarrow in terms of the EMT

$$\partial^{\mu}T_{\mu\nu} = 0$$

In a stationary state: no time dependence, $abla^i T_{ii} = 0$

$$\nabla^i T_{i\nu} = 0$$

Integral form (n: normal to an enclosing surface): $\int_{-\infty}^{\infty} T^{ij} n_j \mathrm{d}\sigma = 0$

Choose the surface as a cross-section of the system in the x-plane, closed at infinity:

Finally, integrate over x:

$$\int T^{xi} \mathrm{d}y \mathrm{d}z = 0.$$

$$\int T^{xi} \mathrm{d}^3 r = 0$$

Diagonal element like i=x: pressure

$$0 = \int p(r) \mathrm{d}^3 r$$

Zur Dynamik der Relativitätstheorie von M. Laue. 1911

The sign of the D-term and system's stability

On LHCb pentaquarks as a baryon- $\psi(2S)$ bound state – prediction of isospin $\frac{3}{2}$ pentaquarks with hidden charm *

Irina A. Perevalova, Maxim V. Polyakov, ^{2,3} and Peter Schweitzer^{4,5}

arXiv:1607.07008

A less trivial local criterion can be obtained by considering that at any chosen distance r the force exhibited by the system on an infinitesimal piece of area dAe_r^i must be directed outwards. If this was not the case, the system would collapse. Since this force is $F^i(\mathbf{r}) = T^{ij}(\mathbf{r}) dA e_r^j = \left[\frac{2}{3} s(r) + p(r)\right] dA e_r^i$ we obtain the criterion

$$\frac{2}{3}s(r) + p(r) > 0. (18)$$

We checked that the condition (18) is satisfied in all systems we are aware of where EMT densities were studied [9, 10, 25-31]. As this includes unstable systems, apparently also (18) is a necessary but not sufficient condition for stability. Due to its local character, it provides a stronger criterion than the von Laue condition (14) and will play an important role below. Interestingly, the criterion (18) allows one to draw a conclusion on the sign of the D-term. We see that

$$0 < 4\pi \int_0^\infty dr \ r^4 \left(\frac{2}{3} s(r) + p(r)\right) = -\frac{2d_1}{M_N} + \frac{4d_1}{5M_N} = -\frac{6d_1}{5M_N} \ . \tag{19}$$

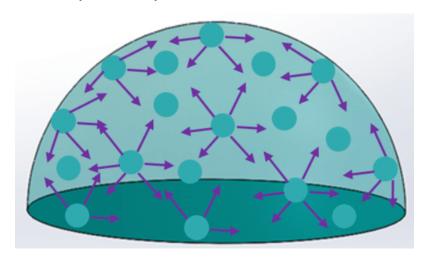
Thus, if a system satisfies the local stability criterion (18), then it must necessarily have a negative D-term (but a negative D-term does not imply that s(r) and p(r) satisfy (18), so the opposite is in general not true). Indeed, in all systems studied so far the D-terms were found to be negative [9, 10, 25 + 31].

Two types of stable systems:

$$0 = \int p(r) \mathrm{d}^3 r$$

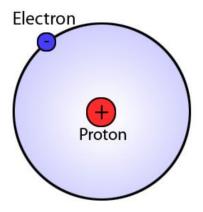
Max von Laue's stability condition

"Liquid droplet"



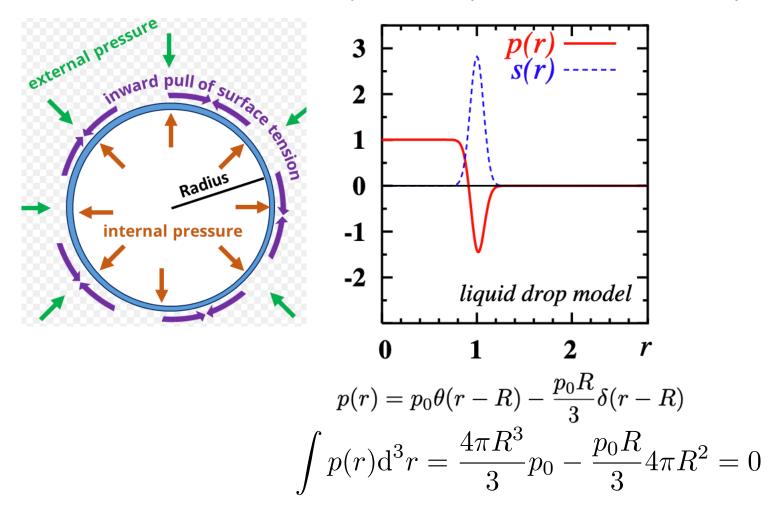
tension at large r: D < 0

"Atom"



tension at small r: D > 0

von Laue's stability example 1: liquid droplet



Forces inside hadrons: pressure, surface tension, mechanical radius, and all that

Maxim V. Polyakov^{1,2} and Peter Schweitzer³

1805.06596

von Laue's stability example 2: hydrogen atom

$$T^{ij}(\vec{r}) = mv^i v^j \delta^3 \left(\vec{r} - \vec{x}(t) \right) - E^i E^j + \frac{\delta^{ij}}{2} \vec{E}^2$$

Electron's motion

Electron's and proton's electric field

We want to show
$$\int d^3 \vec{r} \; T^{ii}(\vec{r}) = 2T + V = 0$$
 as in virial theorem

$$\vec{E} pprox rac{e}{4\pi} rac{\vec{r}}{r^3} - rac{e}{4\pi} rac{\vec{r} - \vec{x}(t)}{|\vec{r} - \vec{x}(t)|^3} \equiv \vec{E}_p + \vec{E}_e$$

$$\int d^3 \vec{r} \; \vec{E}_e \cdot \vec{E}_p = -\frac{\alpha}{|\vec{x}(t)|} = -\frac{\alpha}{R} = V$$

Example: D-term of a liquid drop

$$p(r) = p_0 \theta(r - R) - \frac{p_0 R}{3} \delta(r - R)$$

$$D = m4\pi^{2} \left(\int_{0}^{R} dr \ r^{4} p_{0} - \frac{p_{0}R}{3} R^{4} \right)$$
$$= -\frac{8\pi^{2}}{15} m p_{0}$$

It is negative because the negative pressure region is at the outer boundary.

Example: D-term of the hydrogen atom

Consider $\int d^3 \vec{r} \; r^2 T^{ii}$

In dimensional regularization, terms homogeneous in r vanish. Potential energy contributions give two integrals,

$$I = I_{1} + I_{2} \equiv \int d^{3}\vec{r} \left(-\frac{2}{|\vec{r}||\vec{r} - \vec{R}|} + \frac{2\vec{R} \cdot (\vec{r} - \vec{R})}{r|\vec{r} - \vec{R}|^{3}} \right)$$

$$I_{1}(D) = -\frac{2}{\pi} \int \frac{d\alpha_{1}d\alpha_{2}}{\sqrt{\alpha_{1}\alpha_{2}}} \int d^{D}\vec{r}e^{-(\alpha_{1} + \alpha_{2})r^{2} - \frac{\alpha_{1}\alpha_{2}}{\alpha_{1} + \alpha_{2}}R^{2}}$$

$$\to 4\pi R|_{D=3} ,$$

$$I_{2}(D) = -2R^{2} \frac{2}{\sqrt{\pi}} \frac{1}{\sqrt{\pi}} \int \frac{d\alpha_{1}d\alpha_{2}\sqrt{\alpha_{1}\alpha_{2}}}{\alpha_{1} + \alpha_{2}} \int d^{D}\vec{r}e^{-(\alpha_{1} + \alpha_{2})r^{2} - \frac{\alpha_{1}\alpha_{2}}{\alpha_{1} + \alpha_{2}}R^{2}}$$

$$\to -4\pi R$$

Together with the kinetic energy contribution, we get

$$\int d^3\vec{r} \ r^2 T^{ii}(\vec{r}) = mv^2 R^2 + \frac{e^2}{32\pi^2} \left(I_1 + I_2 \right) = \alpha R$$

This is positive, reflecting electron-proton attraction (rather than surface tension).

Logarithmic corrections: Lamb vs D-term

Vacuum fluctuations smear electron's position,

$$\overline{\delta^2} = \frac{4\alpha}{\pi m^2} \ln \frac{1}{\alpha}$$

Lamb

$$E = E^{(0)} \left(1 - \frac{16\alpha^3}{3\pi} \ln \frac{1}{\alpha} \right)$$
$$\frac{\Delta E}{E^{(0)}} \sim \frac{\overline{\delta^2}}{a_B^2}$$

D-term

$$D = D^{(0)} \left(1 - \frac{4\alpha}{3\pi} \ln \frac{1}{\alpha} \right)$$
$$\frac{\Delta D}{D^{(0)}} \sim \frac{\overline{\delta^2}}{\lambda_C^2}$$

Only S-states are affected

Universal log-correction: all states

Welton's interpretation of Lamb shift

Electron's position in the H-atom modified by vacuum fluctuations. This changes the potential experienced by the electron,

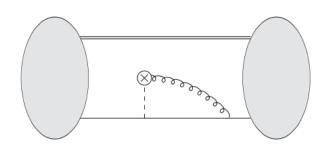
$$\left\langle U_{c}\left(oldsymbol{r}+oldsymbol{q}
ight)
ight
angle =U_{c}\left(r
ight)+\underbrace{\left\langle oldsymbol{q}
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angle }_{0}\cdotoldsymbol{
abla}U_{c}+rac{1}{2}\underbrace{\left\langle q^{i}q^{j}
ight
angle }_{rac{\delta_{ij}}{3}\left\langle q^{2}
ight
angle }
abla^{i}
abla^{j}U_{c}+\ldots,$$

$$\delta U = \langle U_c (\boldsymbol{r} + \boldsymbol{q}) \rangle - U_c (r) \simeq \frac{1}{6} \langle q^2 \rangle \nabla^2 U_c = \frac{1}{6} \langle q^2 \rangle \alpha 4\pi \delta^3 (\boldsymbol{r}).$$

Mean-squared displacement <q2>:

$$\begin{split} \left\langle q^2 \right\rangle &= 2 \int \left(\frac{e}{m\omega^2}\right)^2 \frac{V \mathrm{d}^3 k}{(2\pi)^3} E_{\boldsymbol{k}}^2 \\ &= \frac{2\alpha}{\pi m^2} \int \frac{\mathrm{d}k}{k}. \\ \left\langle q^2 \right\rangle &= \frac{2\alpha}{\pi m^2} \ln \frac{1}{\alpha^2} + \text{non-logarithmic terms.} \\ \left\langle \delta U \right\rangle_{2S} &= \frac{m}{3\pi} \alpha^5 \ln \frac{1}{\alpha} \end{split}$$

Log correction to the D-term



Ji & Liu 2022

$$D_{\rm NLO} = \frac{\alpha}{6\pi} \sum_M \frac{2\vec{v}_{0M} \cdot \vec{v}_{M0}}{D(E_M - E_0)} \bigg(\ln \frac{4(E_M - E_0)^2}{m_e^2} - \frac{1}{4} \bigg)$$
 dimension D = 3

Coefficient of the log: $\sum_{M} \frac{2 \vec{v}_{0M} \cdot \vec{v}_{M0}}{D(E_M - E_0)} \equiv \frac{1}{m_e}$

Summary

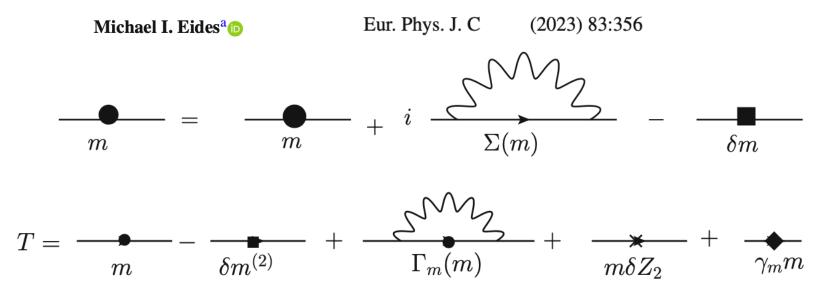
- There are interesting observables in atoms, related to the energy-momentum tensor, in addition to the usually studied electromagnetic current.
- Atomic examples help understand properties of the EMT
- Sign of the D-term can be positive for a stable system
- Logarithmic corrections to the D-term are universal, affecting not only S-states.

Michael Eides' recent work on energy-momentum tensor

Motivation:

A new insight into the EMT properties could arise from consideration of EMT in theories which allow perturbative [treatment]

One-loop electron mass and QED trace anomaly



Graviton-atom interaction:

Of recent interest because can be probed in scattering experiments, via Generalized Parton Distributions [X. Ji; A. V. Radyushkin]