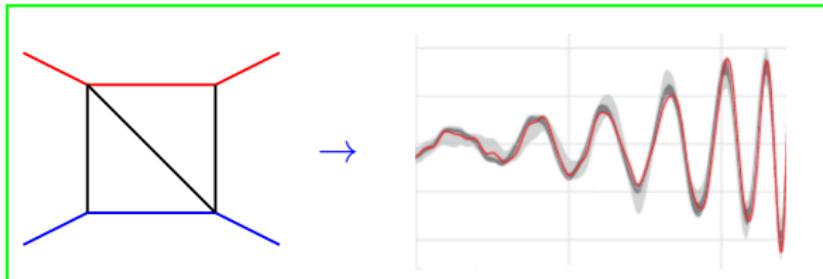


Numerical Unitarity for Binary Dynamics



Fernando Febres Cordero
Department of Physics, Florida State University

RADCOR 2023, Crieff, Scotland, 7/1/2023

Based on [\[arXiv:2205.07357\]](https://arxiv.org/abs/2205.07357)
with Manfred Kraus, Guanda Lin, Michael Ruf and Mao Zeng
and [*in progress*]

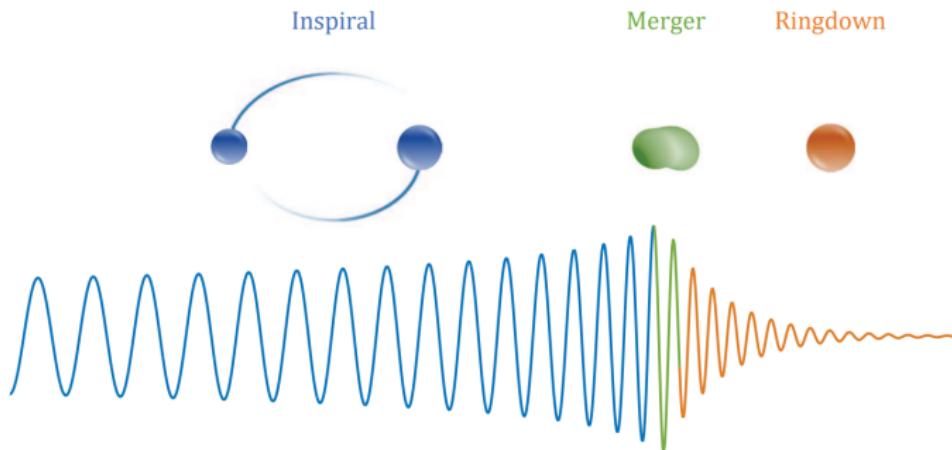
Outline

Precision in Gravitational Waveforms

Conservative Dynamics with Spinning Black Hole at $\mathcal{O}(G^3)$

Multi-Loop Numerical Unitarity for GW Astronomy

Binary Mergers and Gravitational Waves



- A precise theoretical description of the merger process is necessary to exploit the physics potential of measurements at gravitational wave observatories

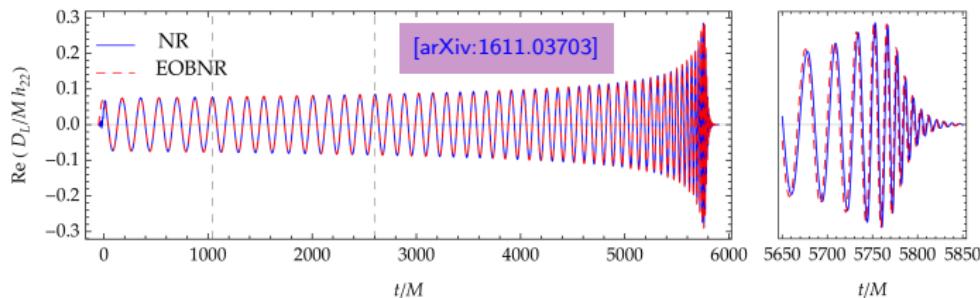
- The inspiral can be described through perturbative methods (in the *Post-Newtonian* and *Post-Minkowskian* frameworks), the merger with numerical relativity, and the ringdown with black hole perturbation theory

Waveforms with the Effective One Body (EOB) Model

- To estimate parameters associated to GW measurements, reliable semi-analytic models are required (NR expensive!).

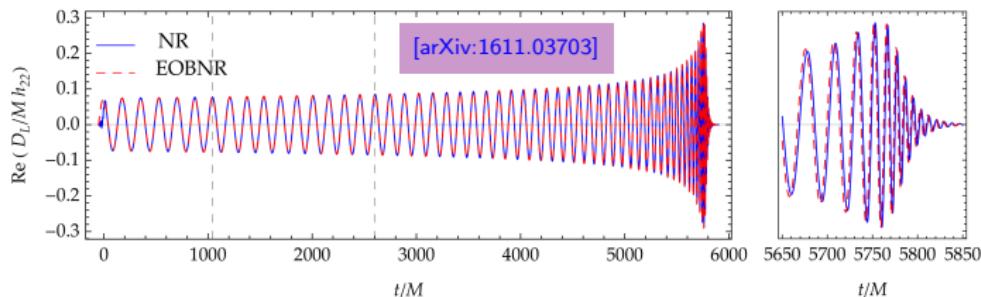
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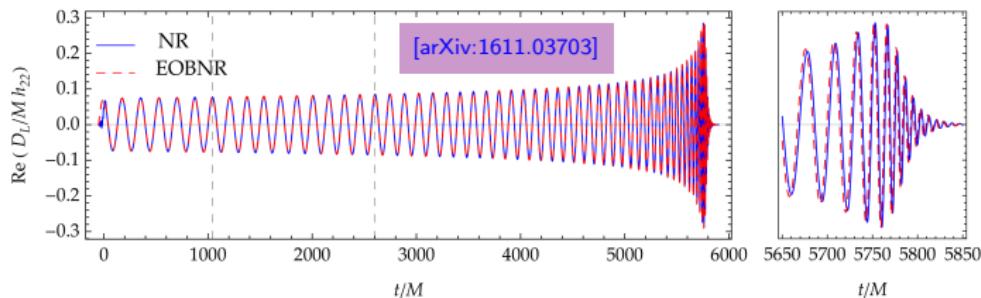
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As input (to fix corresponding parameters) the EOB model takes NR simulations, analytic results for the inspiral and for the ringdown phases of the merger.

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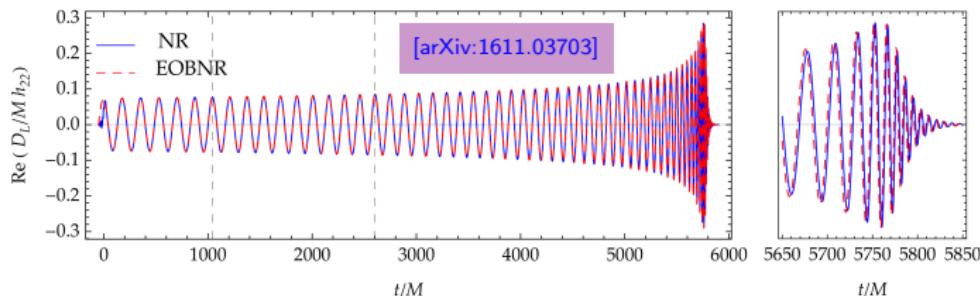
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→ scattering amplitudes can provide critical input for theoretical descriptions in the form of conservative potentials and scattering observables!

Frontier in Perturbative Corrections to Newton's Potential

Buonanno, Khalil, O'Connell, Roiban, Solon, Zeng [arXiv:2204.05194]

$$G(1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + \dots)$$

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- **New results** recently produced with QFT tools
- **State-of-the-art PN** results
- **Expected terms** given sensitivity of future GW observatories

Progress for Spinless Binary Merger

Slide by Manfred Kraus

$$H = \sum_{i=1}^2 \sqrt{\mathbf{p}^2 + m_i^2} + V^{(1)}(\mathbf{r}^2, \mathbf{p}^2) + V^{(2)}(\mathbf{r}^2, \mathbf{p}^2) \frac{\mathbf{L} \cdot \mathbf{S}}{\mathbf{r}^2} + V^{(3)}(\mathbf{r}^2, \mathbf{p}^2) \frac{(\mathbf{p} \cdot \mathbf{S})^2}{\mathbf{r}^2} + V^{(4)}(\mathbf{r}^2, \mathbf{p}^2) \frac{\mathbf{S}^2}{\mathbf{r}^2} + V^{(5)}(\mathbf{r}^2, \mathbf{p}^2) \frac{(\mathbf{r} \cdot \mathbf{S})^2}{\mathbf{r}^4} + \dots$$

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Newton	Einstein,Infeld Hoffmann	Ohta, Okamura, Kimura, Hida	Damour et al Blanchet, Faye	Damour et al	Mastrolia et al Blümlein et al	Blümlein et al

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$$\frac{v^2}{c^2} \sim \frac{GM}{r} \ll 1$$

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→ See Michael Ruf's talk next for **5PL** related calculation

Progress for Binary Merger with a Spinning Black Hole

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Spin Hamiltonian much less known!

Post Newtonian

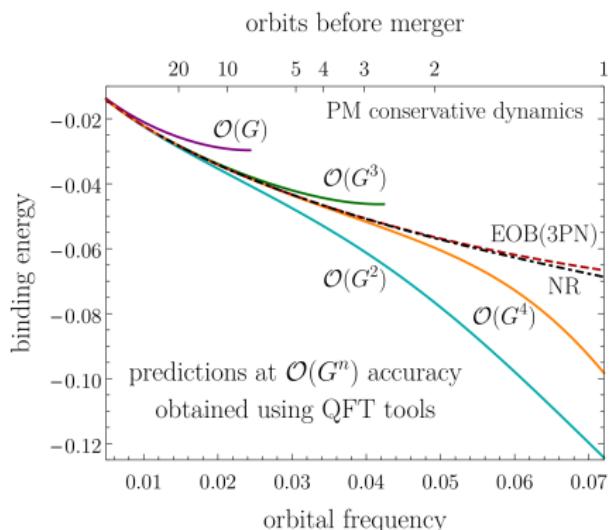
- \mathbf{S}^1 : 3PN $\mathcal{O}(G^4)$ [Levi et al arXiv:2208.14949, Mastrolia et al arXiv:2209.00611]
- \mathbf{S}^2 : 5PN $\mathcal{O}(G^4)$ [Kim,Levi,Yin arXiv:2112.01509]
- \mathbf{S}^3 : 4PN $\mathcal{O}(G^2)$ [Levi,Mougiakakos,Viera arXiv:1912.06276]
- \mathbf{S}^4 : 5PN $\mathcal{O}(G^2)$ [Levi,Teng arXiv:2008.12280]

Post Minkowskian

- \mathbf{S}^2 : 2PM [Bern et.al arXiv:2005.03071], [Kosmopoulos,Luna arXiv:2102.10137]
3PM [Febres Cordero, MK, Lin, Ruf, Zeng arXiv:2205.07357]
- \mathbf{S}^4 : 2PM [Chen,Chung,Huang arXiv:2111.13639]
- \mathbf{S}^5 : 2PM [Bern et al arXiv:2203.06202]
- \mathbf{S}^∞ : 2PM [Aoude,Haddad,Helset arXiv:2203.06197,arXiv:2205.02809]

Phenomenological Impact

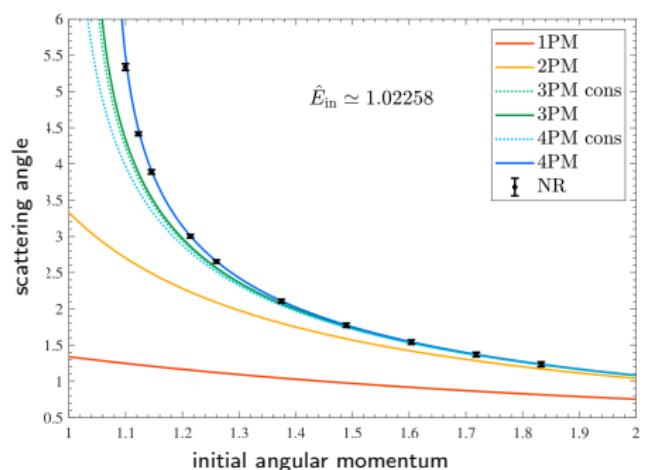
Antonelli, Buonanno, Steinhoff, van de Meent, Vines [arXiv:1901.07102]



- Exploration of impact of higher-order corrections
- Benchmark quantities: bounding energy and scattering angle
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- Both bound and unbound observables studied

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Damour, Rettegno [arXiv:2211.01399]



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QFT Techniques for Classical Observables

- Techniques based on **scattering amplitudes** can produce information for classical observables
 - **Effective field theory** approaches
Goldberger, Rothstein; Cheung, Rothstein, Solon; ...
 - '**in-in'** formalism for computing (inclusive) observables
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 - Here: **EFT matching in dimensional regularization**, **modular arithmetics** and **numerical unitarity**

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Scalar-Vector Scattering

- Use $2 \rightarrow 2$ quantum scattering with massive spin s for conservative dynamics through $2s$ powers in spin Vaidya [arXiv:1410.5348]

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- Consider theory of **massive scalar and vector fields** minimally coupled to gravity

$$\mathcal{L} = \sqrt{-g} \left[-\frac{2R}{\kappa^2} + \frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - \frac{1}{2} m_\phi^2 \phi^2 - \frac{1}{4} g^{\mu\alpha} g^{\nu\beta} F_{\alpha\beta} F_{\mu\nu} + \frac{1}{2} m_A^2 g^{\mu\nu} A_\mu A_\nu \right]$$

$\kappa = \sqrt{32\pi G}$ and expand in the **weak field approximation**

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Our scattering process:



Scattering Amplitude Calculation

Parra-Martinez, Ruf, Zeng [arXiv:2005.04236]

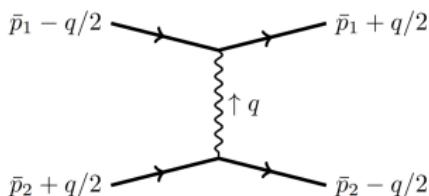
Kinematics:

$$V(p_1) + s(p_2) \rightarrow s(p_3) + V(p_4)$$

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$$\bar{p}_i \cdot q = 0$$

$$\boxed{\sigma = \frac{p_1 \cdot p_2}{m_A m_\phi}}$$



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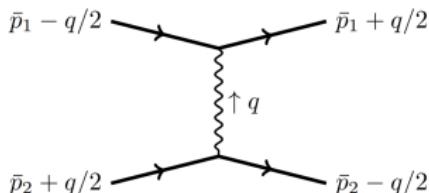
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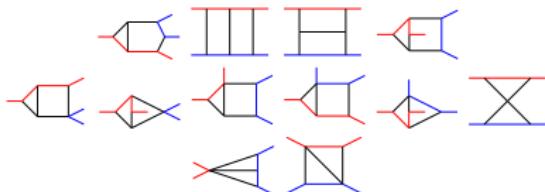
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Multi-loop numerical
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Bern, Cheung, Roiban, Shen, Solon, Zeng [arXiv:1908.01493]



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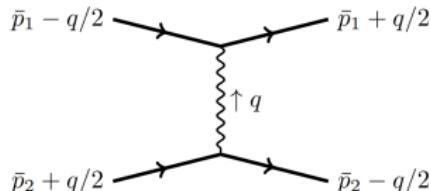
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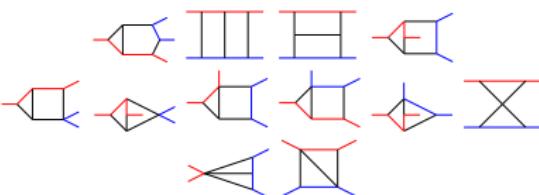
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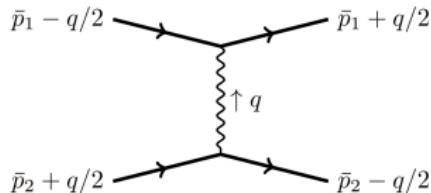
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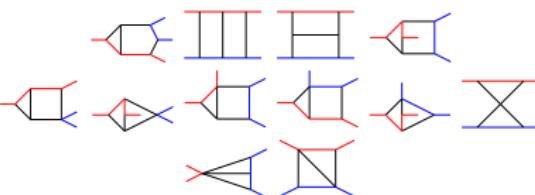
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Abreu, FFC, Ita, Kraus, Page, Pascual, Ruf, Sotnikov [arXiv:2009.11957]

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Smirnov, Chuharev [arXiv:1901.07808]

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The amplitudes
(an example):

$$\begin{aligned} M_4^{(2,2)} = & \frac{1}{\epsilon^2} \left[-\frac{(m_A^3 m_\phi + m_A m_\phi^3) \sigma (1 - 2\sigma^2)^2}{2(8\pi)^2 (\sigma^2 - 1)^2} - \frac{m_A^2 m_\phi^2 (1 - 2\sigma^2)(1 - 8\sigma^4)}{2(16\pi)^2 (\sigma^2 - 1)^2} \right] \\ & + \frac{1}{\epsilon} \left[-\frac{m_A^2 m_\phi^2 \sigma (-6 + \sigma^2)(1 + 2\sigma^2)}{2(4\pi)^2 (\sigma^2 - 1)^{3/2}} \operatorname{arccosh}(\sigma) + \frac{m_A m_\phi^3 \sigma (3 - 16\sigma^2 + 16\sigma^4)}{2(8\pi)^2 (\sigma^2 - 1)^2} \right. \\ & \left. + \frac{m_A^3 m_\phi (\sigma - 5\sigma^3 + 5\sigma^5)}{2(4\pi)^2 (\sigma^2 - 1)^2} + \frac{m_A^2 m_\phi^2 (77 + 189\sigma^2 - 360\sigma^4 + 136\sigma^6)}{3(16\pi)^2 (\sigma^2 - 1)^2} \right] + \mathcal{O}(\epsilon^0) \end{aligned}$$

...

FFC, Kraus, Lin, Ruf, Zeng [arXiv:2205.07357]

Effective Field Theory Amplitude

- We employ the **non-relativistic, non-local, 3D EFT**:

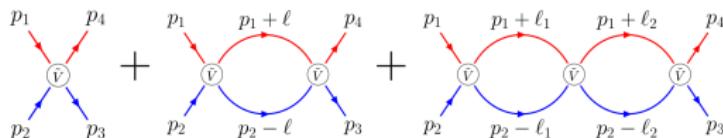
$$L_{\text{EFT}} = \int_{\mathbf{k}} \hat{\phi}^\dagger(-\mathbf{k}) \left(i\partial_t - \sqrt{\mathbf{k}^2 + m_\phi^2} \right) \hat{\phi}(\mathbf{k}) + \int_{\mathbf{k}} \hat{A}^{\dagger,i}(-\mathbf{k}) \left(i\partial_t - \sqrt{\mathbf{k}^2 + m_A^2} \right) \hat{A}^i(\mathbf{k}) \\ - \int_{\mathbf{k}, \mathbf{k}'} \tilde{V}_{ij}(\mathbf{k}, \mathbf{k}') \hat{A}^{\dagger,i}(\mathbf{k}') \hat{A}^j(\mathbf{k}) \hat{\phi}^\dagger(-\mathbf{k}') \hat{\phi}(-\mathbf{k})$$

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- As described by **Cheung, Rothstein, Solon [arXiv:1808.02489]** the EFT amplitude is given by **iterated bubbles**:

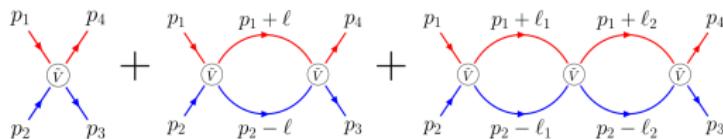


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- The potential \tilde{V}_{ij} is **decomposed** in terms of $\{\delta^{ij}, \bar{\mathbf{p}}^{[i}\mathbf{q}^{j]}, \dots\}$ $\sim \{\mathbb{1}, -i(\mathbf{q} \times \bar{\mathbf{p}}) \cdot \mathbf{S}\}$. From the **amplitudes' matching**, we can then fix the potential!

Example Result: Spin-Orbit Interaction

- The $(\mathbf{q} \times \bar{\mathbf{p}}) \cdot \mathbf{S}$ term in the **conservative potential** is a linear combination of the coefficients:

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$$\begin{aligned}
 c_i^{(2)}(\mathbf{k}^2) &= c_{i,\text{red}}^{(2)}(\mathbf{k}^2) + c_{i,\text{iter}}^{(2)}(\mathbf{k}^2) + \frac{c_i^{(1)}(\mathbf{k}^2)}{m_A^2(\gamma_1 + 1)} \\
 c_{1,\text{red}}^{(2)}(\mathbf{k}^2) &= -\frac{2\sigma m_\phi}{E\xi}, \quad c_{2,\text{red}}^{(2)}(\mathbf{k}^2) = \frac{m_\phi(4m_A + 3m_\phi)\sigma(5\sigma^2 - 3)}{4E\xi(\sigma^2 - 1)}, \\
 c_{3,\text{red}}^{(2)}(\mathbf{k}^2) &= \frac{m_\phi}{E\xi(\sigma^2 - 1)^2} \left[-2m_A^2\sigma(3 - 12\sigma^2 + 10\sigma^4) - \left(\frac{83}{6} + 27\sigma^2 - 52\sigma^4 + \frac{44}{3}\sigma^6\right)m_A m_\phi - m_\phi^2\sigma \left(\frac{7}{2} - 14\sigma^2 + 12\sigma^4\right) \right. \\
 &\quad \left. + \frac{(4m_A + 3m_\phi)E}{4}\sigma(2\sigma^2 - 1)(5\sigma^2 - 3) + 4m_A m_\phi\sigma(\sigma^2 - 6)(2\sigma^2 + 1)\sqrt{\sigma^2 - 1}\operatorname{arccosh}(\sigma) \right], \\
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 &\quad + c_1^{(1)} \left(c_1^{(2)} \left(\left(-\frac{3E^2\xi^2}{k^2} + 6\xi - 2\right) \frac{\partial c_1^{(1)}}{\partial \mathbf{k}^2} - \frac{4}{3}E^2\xi^2 \frac{\partial^2 c_1^{(1)}}{\partial (\mathbf{k}^2)^2} \right) \right. \\
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- Other terms in the potential **more involved**, but still relatively simple

Outline

Precision in Gravitational Waveforms

Conservative Dynamics with Spinning Black Hole at $\mathcal{O}(G^3)$

Multi-Loop Numerical Unitarity for GW Astronomy

Two-Loop Numerical Unitarity

Decompose \mathcal{A} in terms of *master integrals*:

$$\mathcal{A}^{(L)} = \sum_{\Gamma \in \Delta} \sum_{i \in M_{\Gamma}} c_{\Gamma,i} \mathcal{I}_{\Gamma,i}$$

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- **Master-Surface Basis**

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- Given diagram Γ make an *adaptive momentum parametrization*

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These integrand parametrizations have been **automated in Caravel**. Typically less than **$\mathcal{O}(5\%)$ (IBP-)reducible monomials** (with scattering-plane variables) remain!

Master/Surface Decompositions

[Ita '15; Abreu, FFC, Ita, Page, Zeng '17]

Consider the integration by parts (IBP) relation on Γ

$$0 = \int \prod_i d^D \ell_i \frac{\partial}{\partial \ell_j^\nu} \left[\frac{u_j^\nu}{\prod_{k \in P_\Gamma} \rho_k} \right]$$

making it *unitarity compatible* (controlling the propagator structure) [Gluza, Kadja, Kosower '10; Schabinger '11]

$$u_j^\nu \frac{\partial}{\partial \ell_j^\nu} \rho_k = f_k \rho_k$$

Write ansatz for u_j^ν expanded in external and loop momenta, and solve polynomial equations using algebraic geometry techniques

Build a full set of surface terms and fill the rest of the space with master integrands

Related [Boehm, Georgoudis, Larsen, Schulze, Zhang '16 - '19]
[Agarwal, von Manteuffel '19]

Surface Terms Factory

[Abreu, FFC, Ita, Jaquier, Page, Ruf, Sotnikov [arXiv:2002.12374]]

Solutions to u_j^ν are power-counting independent. When parametrizing a given numerator of a $\Gamma \in \Delta$ we need to consider the required power-counting for the theory at hand.

But we can *industrially* produce surface terms by considering polynomials $t_r(\ell_l)$, and then considering the vector $t_r(\ell_l)u_j^\nu$:

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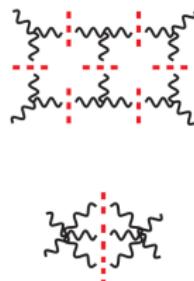
A four-graviton amplitude calculation in **Einstein gravity** structurally the same as a four-gluon amplitude calculation in **QCD**!

Unitarity Approach to Computing Integrand Coefficients

[Bern, Dixon, Dunbar, Kosower] [Britto, Cachazo, Feng]

- In on-shell configurations of ℓ_l , the integrand factorizes and produces a *cut equation*:

$$\sum_{\text{states}} \prod_{i \in T_\Gamma} \mathcal{A}_i^{\text{tree}}(\ell_l^\Gamma) = \sum_{\substack{\Gamma' \geq \Gamma \\ k \in Q_{\Gamma'}}} \frac{c_{\Gamma',k} m_{\Gamma',k}(\ell_l^\Gamma)}{\prod_{j \in (P_{\Gamma'}/P_\Gamma)} \rho_j(\ell_l^\Gamma)}$$



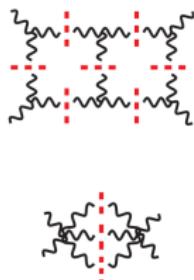
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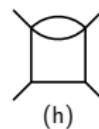
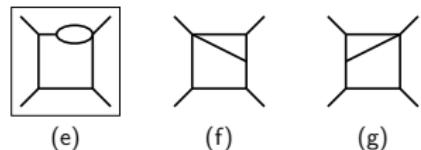
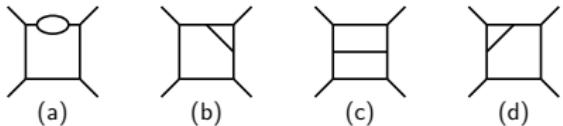


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Never construct analytic integrand, numerics for every phase-space point!

Solving Cut Equations

[Abreu, FFC, Ita, Jaquier, Page [arXiv:1703.05255]]

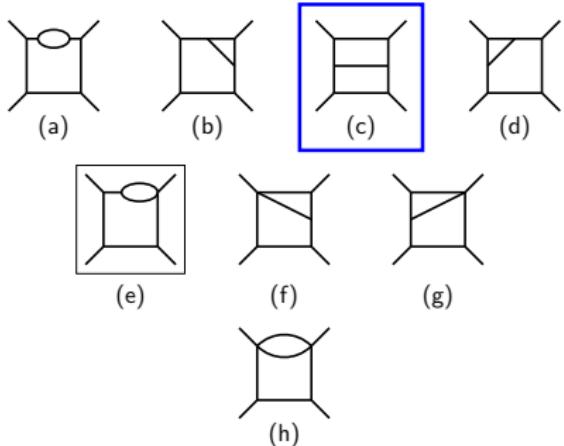


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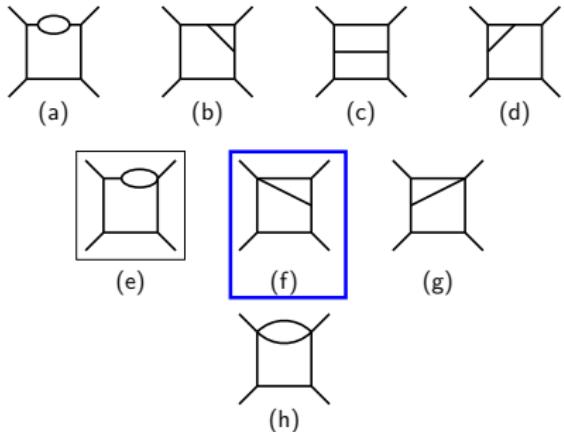
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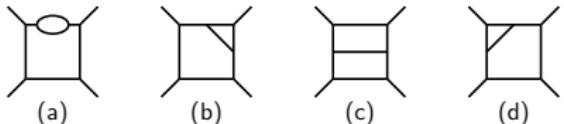


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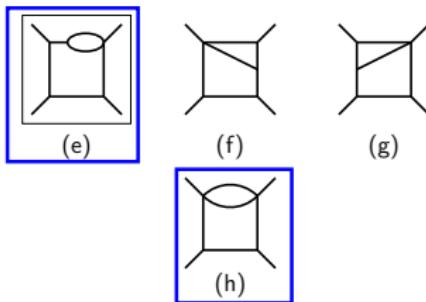
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And for the combined single-pole diagram an bubble-box:

$$N \left(\text{Diagram} \right) + \frac{1}{\rho_{he}} N \left(\text{Diagram}, \ell_l^h \right) = R \left(\text{Diagram}, \ell_l^h \right) - \frac{1}{\rho_{hf}} N \left(\text{Diagram}, \ell_l^h \right) - \frac{1}{\rho_{hg}} N \left(\text{Diagram}, \ell_l^h \right) - \frac{1}{(\rho_{he})^2} N \left(\text{Diagram}, \ell_l^h \right) - \frac{1}{\rho_{hf}\rho_{fb}} N \left(\text{Diagram}, \ell_l^h \right) - \frac{1}{\rho_{hf}\rho_{fc}} N \left(\text{Diagram}, \ell_l^h \right) - \frac{1}{\rho_{hg}\rho_{gd}} N \left(\text{Diagram}, \ell_l^h \right)$$

The CARAVEL Framework

A framework to *explore* multi-loop multi-leg scattering amplitudes in the SM and beyond



The CARAVEL Framework

A framework to *explore* multi-loop multi-leg scattering amplitudes in the SM and beyond

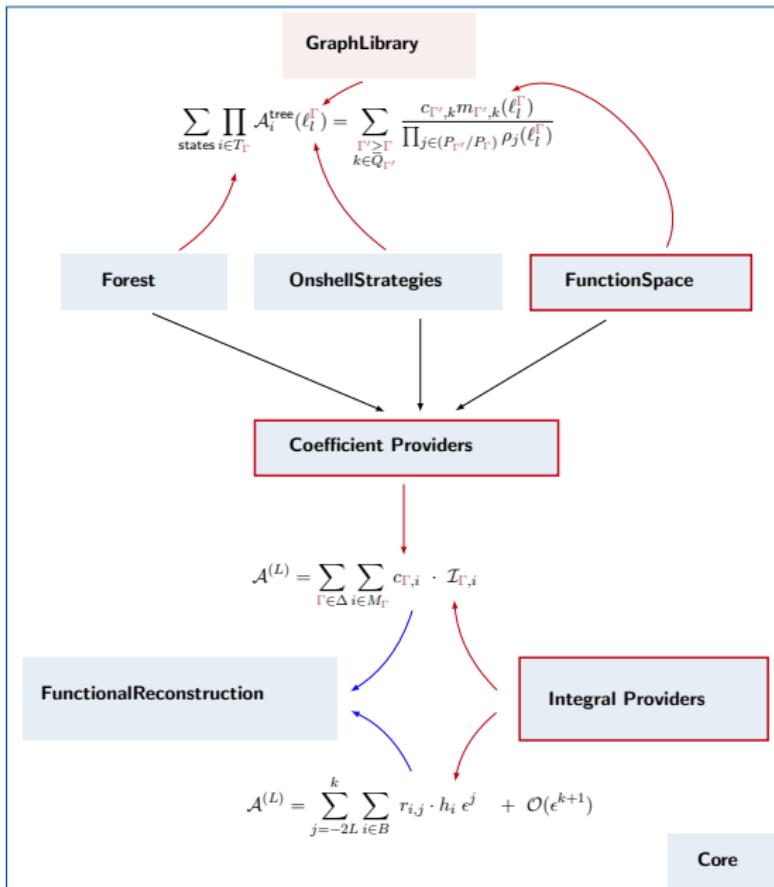
A modular C++17 library *implementing* the multi-loop numerical unitarity method

[Abreu, Dormans, FFC, Ita, Kraus, Page, Pascual, Ruf, Sotnikov, arXiv:2009.11957]

- Numerics in (high-precision) floating-point, rational and modular arithmetic
- Generic design for calculations in QFT, e.g. in the SM, gravity theories, and more
- Algebraic tools for semi-analytical calculations in C++
- Publicly available @ GitLab!



The CARAVEL Framework



Includes general tools for:

- *D*-dimensional kinematics
- graph isomorphism techniques
- tree-level and multi-loop cut calculations
- Generic scattering-plane integrand parametrizations
- Selected master-surface decompositions
- on-shell phase-space parametrizations
- Feynman integral handling
- Algebraic tools
- ...

Caravel @ GitLab:

[https://gitlab.com/
caravel-public/caravel](https://gitlab.com/caravel-public/caravel)

Outlook

- QFT techniques developed for collider phenomenology playing a key role in improving the description of compact binary mergers
- We presented the conservative potential for compact binary systems with a spinning black hole at $\mathcal{O}(G^3)$ and to all orders in velocity and including up-to S^2 terms
- The multi-loop numerical unitarity method is ready to tackle further calculations
- Future directions include radiation effects, higher spins, finite-size effects, and higher loops

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Thanks!