# Coulomb Functions for Nuclear Physics

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 $24 \ \mathrm{June} \ 2024$ 

#### AZURE2 R-matrix Summer School University of Edinburgh June 23-28, 2024

CRB supported in part by the U.S. D.O.E. through grants No. DE-FG02-88ER40387 and DE-NA0004065







# References for Coulomb Functions

- M. Abramowitz, Coulomb Wave Functions, in M. Abramowitz and I. Stegun, Handbook of Mathematical Functions, Dover Publications (New York), 1965.
- I.J. Thompson, Coulomb Functions, in F.W. Olver et al., NIST Handbook of Mathematical Functions, Cambridge University Press (New York), 2010.
- I.J. Thompson, Coulomb Functions, in F.W. Olver et al., NIST Digital Library of Mathematical Functions, http://dlmf.nist.gov/.

Speaking of NIST, here is a useful website for the fundamental physical constants: https://physics.nist.gov/cuu/Constants/index.html.

## The Differential Equation

 $\blacktriangleright$  In terms of physical parameters, a Coulomb function u in coordinate space satisfies

$$-\frac{\hbar^2}{2\mu}\frac{d^2u}{dr^2} + \frac{Z_1Z_2e^2}{r}u + \frac{\hbar^2}{2\mu}\frac{\ell(\ell+1)}{r^2}u = Eu,$$

where  $r \ge 0$  is the radial coordinate, E is the center-of-mass energy,  $\mu$  is the reduced mass.

▶ We also have

$$V_c = \frac{Z_1 Z_2 e^2}{r},$$

the repulsive Coulomb potential,

▶ and

$$V_{\text{eff}} = \frac{\hbar^2}{2\mu} \frac{\ell(\ell+1)}{r^2},$$

an effective repulsive potential corresponding to the centrifugal or angular momentum barrier.

### **Dimensionless** Parameters

▶ In terms of the dimensionless parameters  $\rho$  and  $\eta$ , we have  $u(\ell, \eta, \rho)$  and this equation becomes

$$u'' + \left[1 - \frac{2\eta}{\rho} - \frac{\ell(\ell+1)}{\rho^2}\right]u = 0,$$

where  $\rho = kr$ ,  $k = \sqrt{2\mu E/\hbar^2}$ ,  $\eta k = Z_1 Z_2 e^2 \mu/\hbar^2$ , and  $' \equiv d/d\rho$ .

- Note that  $\rho \propto \sqrt{E}$  and  $\eta \propto 1/\sqrt{E}$ .
- For a given pair of nuclei and  $\ell$ , one can consider Coulomb functions to be functions of the two variables (E, r) or  $(\eta, \rho)$ .

## The Functions

- ▶  $F_{\ell}(\eta, \rho)$ : the regular Coulomb function,  $F_{\ell}(\eta, \rho \to 0) \propto \rho^{\ell+1}$
- ►  $G_{\ell}(\eta, \rho)$ : the irregular Coulomb function,  $G_{\ell}(\eta, \rho \to 0) \propto \rho^{-\ell}$

$$H_{\ell}^{\pm}(\eta,\rho) = G_{\ell}(\eta,\rho) \pm iF_{\ell}(\eta,\rho)$$

▶ Wronskian:  $F'_{\ell}G_{\ell} - F_{\ell}G'_{\ell} = 1$ 

## Various Phases

▶ The Coulomb phase shift  $\sigma_{\ell}$  is also sometimes needed:

$$e^{2i\sigma_{\ell}} = \frac{\Gamma(1+\ell+i\eta)}{\Gamma(1+\ell-i\eta)} = \frac{(\ell+i\eta)\dots(1+i\eta)}{(\ell-i\eta)\dots(1-i\eta)}e^{2i\sigma_{0}}.$$

Note that  $\sigma_{\ell} - \sigma_0$  does not require the  $\Gamma$  function.

▶ The asymptotic phase  $\theta_{\ell}$  is defined to be

$$\theta_{\ell} = \rho - \eta \log(2\rho) - \frac{1}{2}\ell\pi + \sigma_{\ell}.$$

For  $\rho \to \infty$ , we have

$$H_{\ell}^{\pm} \sim \exp(\pm i\theta_{\ell}),$$

which is useful for calculating the S matrix.

Amplitude, Phase, Penetration, and Shift

▶ The Amplitude  $A_{\ell}$ , Phase  $\phi_{\ell}$ , Penetration  $P_{\ell}$ , and Shift  $S_{\ell}$  are defined according to:

$$A_{\ell} = (F_{\ell}^2 + G_{\ell}^2)^{1/2}$$
  

$$\phi_{\ell} = \tan^{-1} F_{\ell}/G_{\ell}$$
  

$$H_{\ell}^{\pm} = A_{\ell} \exp(\pm i\phi_{\ell})$$
  

$$P_{\ell} = \frac{\rho}{A_{\ell}^2}$$
  

$$S_{\ell} = \frac{\rho A_{\ell}'}{A_{\ell}} = \frac{\rho(A_{\ell}^2)'}{2A_{\ell}^2}$$

- ▶ Note that we also have  $S_{\ell} + iP_{\ell} = \rho \frac{H_{\ell}^{+\prime}}{H_{\ell}^{+}}$ , which can be shown using the Wronskian.
- ▶ In *R*-matrix calculations, one traditionally uses  $P_{\ell}$ ,  $S_{\ell}$ , and  $\phi_{\ell}$  (you only need three).

# Classical Turning Radius

- Consider the quantity  $1 \frac{2\eta}{\rho} \frac{\ell(\ell+1)}{\rho^2}$  in the differential equation.
- When it is > 0, the Coulomb functions are oscillatory.
- When it is < 0, the Coulomb functions are exponential.

• It = 0 for 
$$\rho_{\rm tr} = \eta + [\eta^2 + \ell(\ell+1)]^{1/2}$$
.

- Classical Turning Radius,  $r_{\rm tr}$ :  $kr_{\rm tr} = \rho_{\rm tr}$ .
- ▶ The location of  $r_{tr}$  relative to the nuclear surface strongly impacts the physics. In particular, if  $r_{tr}$  is well outside the nuclear surface, the reaction probability will be strongly reduced by Coulomb and/or angular momentum barriers.

## An Example Plot



Interestingly, both  $A_{\ell}$  and  $\phi_{\ell}$  are monotonic functions of r, for  $0 \leq r < \infty$ .

## Limiting Forms for Small and Large $\rho$

quantity	$\rho \rightarrow 0$	$\rho \rightarrow \infty$
$H_{\ell}^+$	$\left[\rho^{\ell}(2\ell+1)C_{\ell}(\eta)\right]^{-1}+\ldots+i\left[\rho^{\ell+1}C_{\ell}(\eta)+\ldots\right]$	$\exp(i\theta_{\ell})\left[1+\frac{\eta}{2\rho}+i\frac{\eta^2+\ell(\ell+1)}{2\rho}+\ldots\right]$
$A_\ell^2$	$[\rho^{\ell}(2\ell+1)C_{\ell}(\eta)]^{-2}+\ldots$	$1 + \frac{\eta}{\rho} + \frac{3\eta^2 + \ell(\ell+1)}{2\rho^2} + \dots$
$\phi_\ell$	$\rho^{2\ell+1}(2\ell+1)C_{\ell}^{2}(\eta) + \dots$	$\theta_{\ell} + \frac{\eta^2 + \ell(\ell+1)}{2\rho} + \dots$
$P_{\ell}$	$\rho^{2\ell+1}[(2\ell+1)C_{\ell}(\eta)]^2 + \dots$	$\rho - \eta - \frac{\eta^2 + \ell(\ell+1)}{2\rho} + \dots$
$S_\ell$	$-\ell$ +	$-\frac{\eta}{2\rho} - \frac{2\eta^2 + \dot{\ell}(\ell+1)}{2\rho^2} + \dots$

See C.R. Brune, G.M. Hale, and M.W. Paris, Monotonic properties of the shift and penetration factors, Phys. Rev. C 97, 024603 (2018), https://doi.org/10.1103/PhysRevC.97.024603

#### The Gamow factor is defined to be

$$C_{\ell}(\eta) = \frac{2^{\ell} e^{-\pi\eta/2} \left[ \Gamma(\ell+1+i\eta) \Gamma(\ell+1-i\eta) \right]^{1/2}}{\Gamma(2\ell+2)},$$

which for  $\ell = 0$  becomes

$$C_0 = \left[\frac{2\pi\eta}{\exp(2\pi\eta) - 1}\right]^{1/2}$$

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### Low-Energy Limit

For  $E \to 0$ , with radius fixed, we have

$$F_{\ell} \rightarrow \frac{C_0}{2\eta} \frac{x}{2} I_{2\ell+1}(x)$$
  

$$G_{\ell} \rightarrow \frac{x}{C_0} K_{2\ell+1}(x),$$

where  $x = (8\eta\rho)^{1/2}$  is independent of energy and  $\propto \sqrt{r}$ , and  $I_{2\ell+1}(x)$  and  $K_{2\ell+1}(x)$  are the regular and irregular *Modified* Bessel Functions, respectively.

See J. Humblet, Bessel functions expansions of Coulomb wave functions, J. Math. Phys. 26, 656-659, 1985 https://doi.org/10.1063/1.526602.

This implies

$$P_{\ell} \to \frac{\pi \exp(-2\pi\eta)}{4K_{2\ell+1}^2(x)}$$

## Negative Energies

► For negative energies, we need the exponentially-decaying Whittaker function:

$$W_{-\eta_b,\ell+1/2}(2\kappa_b r),$$

where  $\kappa_b = \sqrt{-2\mu E/\hbar^2}$  and  $\eta_b \kappa_b = Z_1 Z_2 e^2 \mu/\hbar^2$ .

- ▶ It is proportional to the analytic continuation of  $H_{\ell}^+$  to negative energies.
- For E < 0, we have  $P_{\ell} = 0$  and

$$S_{\ell} = \frac{r}{W} \frac{dW}{dr}.$$

▶  $S_{\ell}(E)$  is continuous across E = 0, with a value of

$$S_{\ell}(0) = -\ell - \frac{xK_{2\ell}(x)}{K_{2\ell+1}(x)}.$$

In practice,  $\phi_{\ell}$  is not needed for E < 0.

### The Tail a Bound State

▶ Using the differential equation, it can be shown that

$$\frac{\hbar^2}{2\mu} \left[ \frac{W^2}{r} \frac{dS}{dE} \right]_{r=a} = \int_a^\infty W^2 \, dr,$$

see Eq. (A.29) in the appendix of Lane and Thomas.

▶ Normalization condition:

$$N + \sum_{c} C_c^2 \int_{a_c}^{\infty} W_c^2 dr_{\alpha} = 1$$
$$C_c = \frac{(2\mu_{\alpha}a_c)^{1/2}}{\hbar W_\ell(a_c)} N^{1/2} \gamma_c$$
$$N \left[ 1 + \sum_{c} \gamma_c^2 \left. \frac{dS_c}{dE} \right|_{E_R} \right] = 1$$

▶  $\frac{dS}{dE}$  can be computed by a continued fraction technique.

## Computer Codes

- I.J. Thompson and A.R. Barnett, COULCC: A continued-fraction algorithm for Coulomb functions of complex order with complex arguments, Computer Physics Communications 36, 363-372 (1985), https://doi.org/10.1016/0010-4655(85)90025-6, fortran90-ish, verson 36, code available from http://www.ianthompson.org/computation.htm
- N. Michel, Precise Coulomb wave functions for a wide range of complex l, η and z, Computer Physics Communications 176, 232-249 (2007), https://doi.org/10.1016/j.cpc.2006.10.004, c++, code available from CPC Program Library

## Computer Codes, continued

- GNU Scientific Library (GSL), https://www.gnu.org/software/gsl/, c
- A.R. Barnett, COULFG: Coulomb and Bessel functions and their derivatives, for real arguements, by Steed's method, Computer Physics Communications 27, 147-166 (1982), https://doi.org/10.1016/0010-4655(82)90070-4, fortran, c version is available from the LLNL github: https://github.com/LLNL/fudge
- ► I will refer to these as coulcc, cwfcomp, gsl, and coulfg, respectively.

# Comparisons

- For the comparisons, I have considered  ${}^{12}C + \alpha$ , a = 5.5 fm,  $\ell = 2$  and varied the energy from near zero to 10 MeV.
- ▶ The radius is typical of the channel radius that one would utilize for an *R*-matrix analysis of this system.
- ▶ This energy range spans from far below to above the Coulomb and angular momentum barriers.

## Caveats

- This study is only looking at a limited region of parameter space.
- ▶ Just because two codes agree, that does not mean they are correct.
- ▶ In the case of cwfcomp, I have used the default computational parameters (also the AZURE2 default):

```
precision = 1E - 10
sqrt_precision = 1E - 5.
```

Decreasing these values brings the cwfcomp results closer to coulcc, and increases the computational time.

### Shift Function



#### Penetration



#### Phase



## Computational Time



# Findings

- coulcc and cwfcomp agree reasonably well for all energies considered.
- ▶ gsl agrees with the other codes for very low energies and energies above 3 MeV.
- ▶ gsl shows significant disagreements with the other codes for 0.3 < E < 3 MeV.
- coulfg agrees will with coulcc and cwfcomp, except for very low energies.
- There are significant differences in computational speeds: gsl is the fastest, followed by coulfg, then by coulcc, and finally cwfcomp.

# What is going on with gsl?

- ▶ It turns out that gsl uses a WKB approximation when  $1.2 \le \rho < 2\eta$ . For  $\rho < 1.2$ , the power series are used. For  $\rho \ge 2\eta$ , continued fractions are used. Note that  $\rho = 2\eta$  corresponds to the classical turning radius for  $\ell = 0$ .
- ► The range  $1.2 \le \rho < 2\eta$  corresponds to 0.3 < E < 3 MeV for  $^{12}C + \alpha$  at 5.5 fm, which is a critical region of parameter space for this case. The WKB approximation is just not very accurate.
- Besides not being particularly accurate, the gsl Coulomb functions are not continuous functions of energy and radius. This may cause problems for parameter search algorithms in phenomenological *R*-matrix applications.

#### ▶ User beware!

# What about coulfg?

- ▶ It is fast and works very well, except for very low energies.
- ▶ The issue here is a loss of accuracy in the continued fraction method when  $P \ll S$ , which is a known issue. Other codes use different method in this regime.

# Conclusions

- ▶ There are significant differences in the accuracy and computational speed of four commonly used codes for computing Coulomb functions.
- I believe there is room for a new code that optimizes speed and accuracy for  $\rho$  and  $\eta$  real and positive.
- ▶ However, if computational speed is a truly limiting factor, other approaches, such as interpolation from pre-computed tables, should be considered.

# Thank you for your attention.