

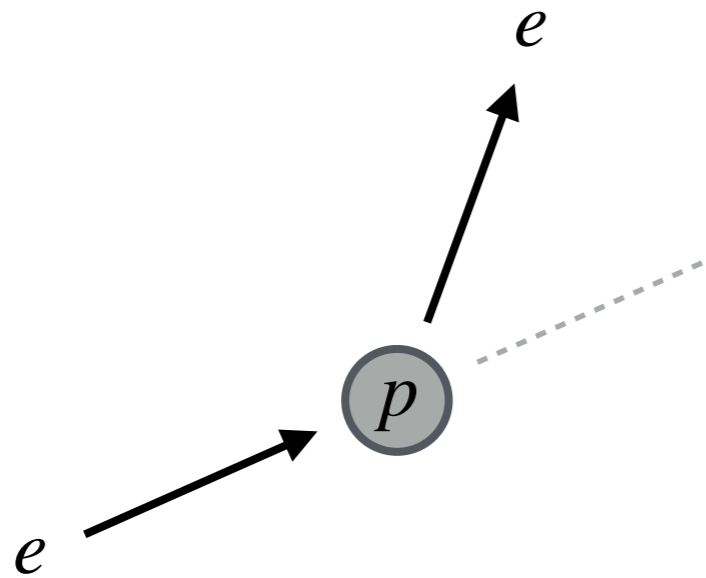
An Effective Field Theory For Extreme Mass Ratios

Clifford Cheung
Caltech

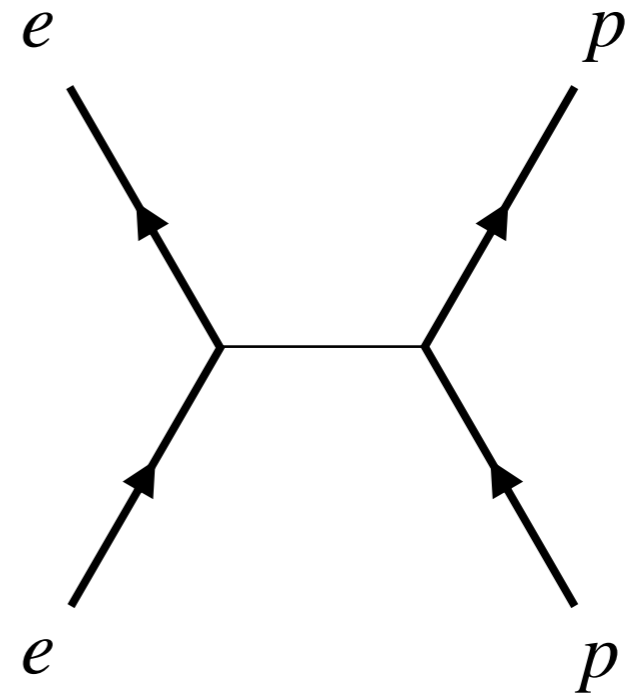
2308.14832 (CC, Parra-Martinez, Rothstein, Shah, Wilson-Gerow)
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Two ways to approach electron-proton dynamics:

i) probe in $1/r$ background

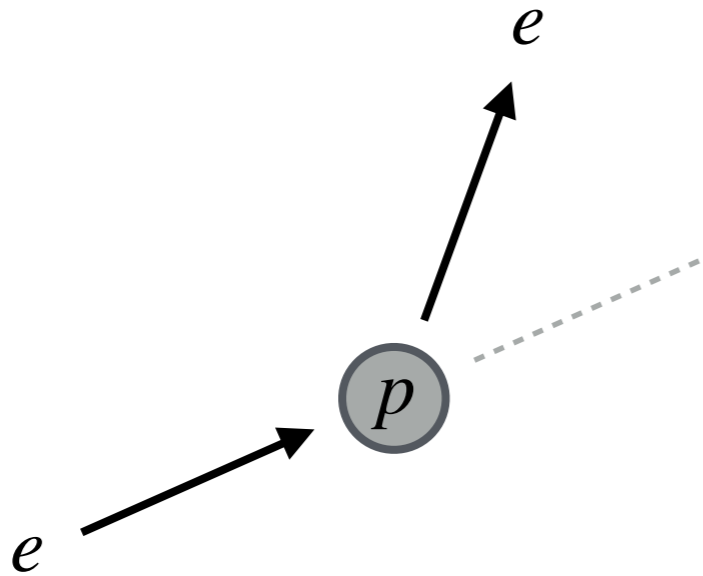


ii) perturbation theory

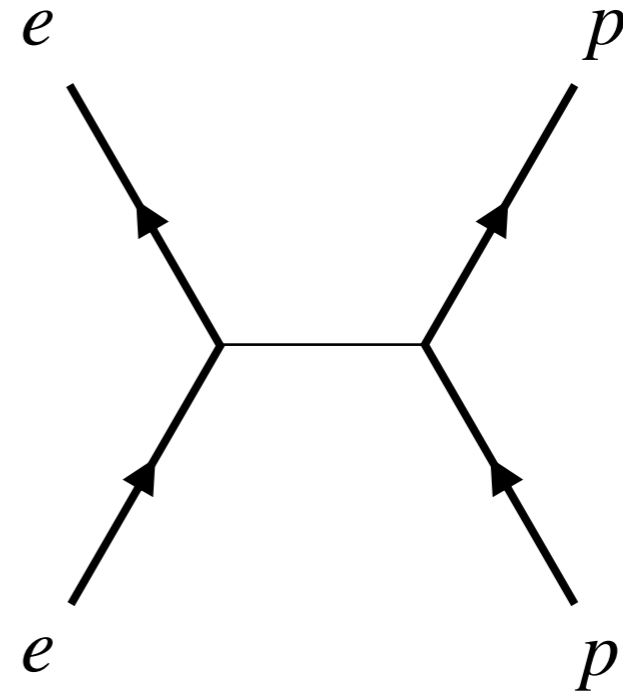


Two ways to approach electron-proton dynamics:

i) probe in $1/r$ background



ii) perturbation theory



When does background field theory end and perturbation theory on a trivial vacuum begin?

The breakdown of the background field theory is signaled by the importance of corrections that:

- enter as powers of $\lambda = m_L/m_H$
- encode heavy particle “wobble”
- are perturbatively calculable

As in HQET or NRQCD, we integrate out the heavy state, but we also keep all orders in α data!

Schwarzschild metric describes the $\mathcal{O}(\lambda^0)$ limit.

Quantum Tree Graphs and the Schwarzschild Solution

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(Received 7 July 1972)

Schwarzschild metric describes the $\mathcal{O}(\lambda^0)$ limit.

static matter source

$$g^{00} + 1 = \text{graviton 1pt function} = \begin{array}{c} \text{diagram 1} \\ G \end{array} + \begin{array}{c} \text{diagram 2} \\ G^2 \end{array} + \begin{array}{c} \text{diagram 3} \\ G^3 \end{array} + \dots$$

The diagram shows a series of Feynman diagrams representing the expansion of the graviton 1-point function. The first diagram is a single wavy line with a shaded circle at the top, labeled G . The second diagram is a wavy line with a shaded circle at the top and two shaded circles at the top, connected by a wavy line, labeled G^2 . The third diagram is a wavy line with a shaded circle at the top and three shaded circles at the top, connected by wavy lines, labeled G^3 . The series continues with an ellipsis.

Schwarzschild metric describes the $\mathcal{O}(\lambda^0)$ limit.

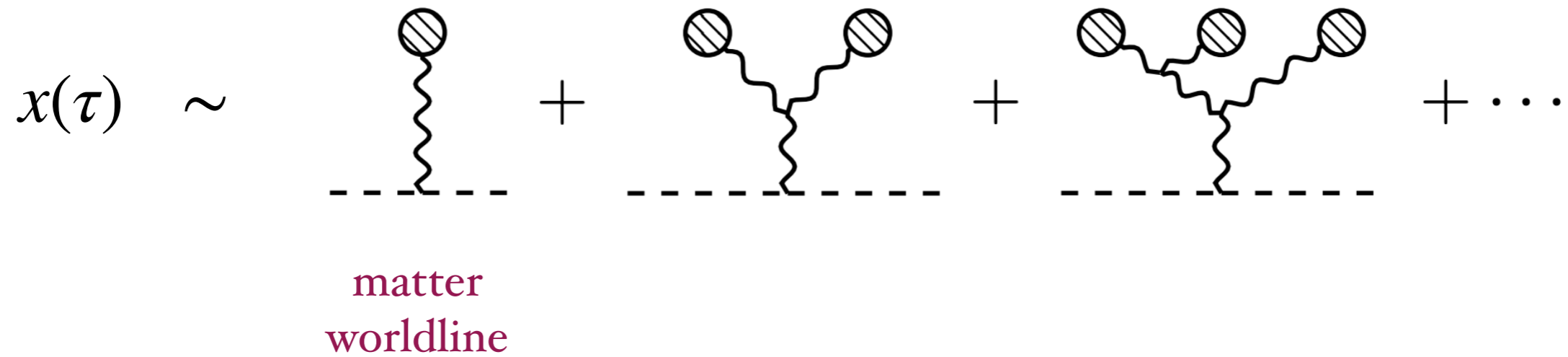
$$g^{00} + 1 = \begin{array}{c} \text{---} \\ | \\ \text{---} \\ G \end{array} + \begin{array}{c} \text{---} \quad \text{---} \\ | \quad | \\ \text{---} \\ G^2 \end{array} + \begin{array}{c} \text{---} \quad \text{---} \quad \text{---} \\ | \quad | \quad | \\ \text{---} \\ G^3 \end{array} + \dots$$

$$g^{00} = -1 - \frac{2GM}{r} - \frac{2G^2M^2}{r^2} + \dots = -\frac{r + GM}{r - GM}$$

all orders
in coupling!

Probe geodesics also encode the $\mathcal{O}(\lambda^0)$ limit.

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Probe geodesics also encode the $\mathcal{O}(\lambda^0)$ limit.

$$x(\tau) \sim \text{---} \begin{array}{c} \text{---} \\ | \\ \text{---} \end{array} \text{---} + \text{---} \begin{array}{c} \text{---} \\ | \\ \text{---} \\ | \\ \text{---} \\ | \\ \text{---} \end{array} \text{---} + \text{---} \begin{array}{c} \text{---} \\ | \\ \text{---} \\ | \\ \text{---} \\ | \\ \text{---} \\ | \\ \text{---} \end{array} \text{---} + \dots$$

Flat space perturbation theory builds the metric and probe geodesics. Let's do the opposite.

flat space loops

(hard)



(Duff)

metric + geodesic

(easy)

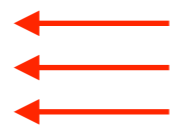
Probe geodesics also encode the $\mathcal{O}(\lambda^0)$ limit.

$$x(\tau) \sim \text{---} \begin{array}{c} \textcircled{\text{---}} \\ | \\ \text{---} \end{array} + \text{---} \begin{array}{c} \textcircled{\text{---}} \quad \textcircled{\text{---}} \\ | \quad | \\ \text{---} \end{array} + \text{---} \begin{array}{c} \textcircled{\text{---}} \quad \textcircled{\text{---}} \quad \textcircled{\text{---}} \\ | \quad | \quad | \\ \text{---} \end{array} + \dots$$

Flat space perturbation theory builds the metric and probe geodesics. Let's do the opposite.

flat space loops

(hard)



(this talk)

metric + geodesic

(easy)

Our goal: to maximally leverage all-orders-in- G classical results to flat space perturbation theory.

exact answer = background field method + everything else

(Schwarzschild metric,
geodesic trajectories)

(“recoil operator”)

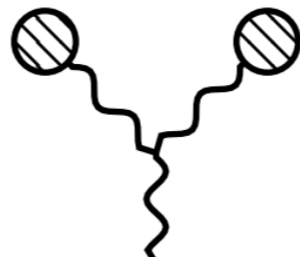
Obviously this is just a reshuffling of perturbation theory. But metrics + geodesics nicely encode loop integrand contributions automatically!

Classical solutions implement integral reduction.

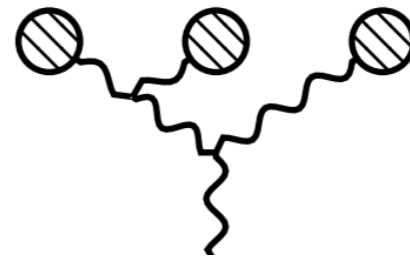
tensor
numerators



$$\frac{1}{r}$$



$$\frac{1}{r^2}$$



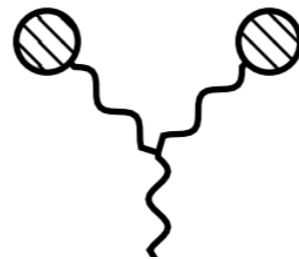
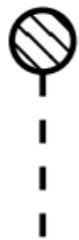
$$\frac{1}{r^3}$$

Classical solutions implement integral reduction.

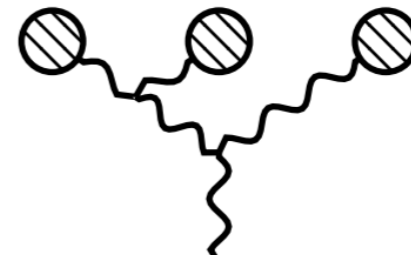
tensor
numerators



$$\frac{1}{r}$$



$$\frac{1}{r^2}$$



$$\frac{1}{r^3}$$



scalar
numerators

I. Electromagnetism

To warmup, take EM with a light + heavy particle,

$$S_{\text{EM}} = \int d^4x \left[-\frac{1}{4} F^2 \right] + \sum_{i=L,H} m_i \int d\tau \left[-\frac{1}{2} \dot{x}_i^2 - z_i \dot{x}_i^\mu A_\mu \right]$$

where we fix charge-to-mass-ratios $z_i = q_i/m_i$ to maximally parallel GR. The SF parameter is:

$$\lambda = \frac{m_L}{m_H}$$

Expand about a $\mathcal{O}(\lambda^0)$ background and trajectory.

$$\begin{array}{ccc} \mathcal{O}(\lambda^0) & & \mathcal{O}(\lambda^0) \\ x_i^\mu = \bar{x}_i^\mu + \delta x_i^\mu & & A_\mu = \bar{A}_\mu + \delta A_\mu \\ \nearrow & & \nearrow \\ \text{closed-form trajectory} & \text{deflection} & \sim 1/r \quad \text{photon} \end{array}$$

From full theory to background field theory:

$$S_{\text{EM}}[\delta A, \delta x_L, \delta x_H] \longrightarrow S_{\text{BF}}[\delta A, \delta x_L]$$

integrate out
heavy particle!

Now consider some explicit OSF backgrounds.

- heavy trajectory

$$\bar{x}_H^\mu(\tau) = u_H^\mu \tau$$

- EM field

$$\bar{A}_\mu(x) = \frac{z_H m_H u_{H\mu}}{4\pi r} \quad \text{s.t.} \quad r = \sqrt{(u_H x)^2 - x^2}$$

- light trajectory

$$\bar{x}_L^\mu(\tau) \quad \text{s.t.} \quad \ddot{\bar{x}}_L^\mu(\tau) - z_L \bar{F}^{\mu\nu}(\bar{x}_L) \dot{\bar{x}}_{L\nu} = 0$$

Perturbative contributions from EM field:

$$\bar{A}_\mu(x) = \frac{z_H m_H u_{H\mu}}{4\pi r} \quad \rightarrow \quad \tilde{A}_\mu(p) = \frac{z_H m_H u_{H\mu}}{p^2} \delta(u_H p)$$

What about the self-energy contributions?

$$\begin{aligned} \bar{A}_\mu(\bar{x}_H) &\sim \lim_{r \rightarrow 0} \frac{1}{r^{1-\epsilon}} \sim \lim_{r \rightarrow 0} \int d^{3-\epsilon} q \frac{e^{iqr}}{q^2} \\ &\rightarrow \int d^{3-2\epsilon} q \frac{1}{q^2} = 0 \quad \text{dimensional regulator} \end{aligned}$$

Perturbative contributions from EM field:

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$$\rightarrow \quad \bar{A}_\mu(\bar{x}_H) = \bar{F}_{\mu\nu}(\bar{x}_H) = 0$$

Terms in the action relevant to $\mathcal{O}(\lambda^1)$ are given by:

$$S_{\text{EM}} = m_H \int d\tau \left[-\frac{1}{2} \delta \dot{x}_H^2 - z_H \delta x_H^\mu \dot{\bar{x}}_H^\nu \delta F_{\mu\nu}(\bar{x}_H) + \dots \right]$$

$$+ \int d^4x \left[-\frac{1}{4} \delta F_{\mu\nu} \delta F^{\mu\nu} - \delta A_\mu \bar{J}_L^\mu \right]$$

with $\mathcal{O}(\lambda^1)$ current $\bar{J}_L^\mu(x) = z_L m_L \int d\tau \delta^4(x - \bar{x}_L) \dot{\bar{x}}_L^\mu$

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Gaussian in δx_H

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Integrating out δx_H yields the “recoil operator”.

$$m_H \int d\tau \left[-\frac{1}{2} \delta \dot{x}_H^2 - z_H \delta x_H^\mu \dot{\bar{x}}_H^\nu \delta F_{\mu\nu}(\bar{x}_H) \right]$$



$$S_{\text{recoil}} = -\frac{1}{2} z_H^2 m_H \int d\tau \dot{\bar{x}}_H^\alpha \delta F_{\alpha\mu}(\bar{x}_H) \frac{1}{\partial_\tau^2} \dot{\bar{x}}_H^\beta \delta F_{\beta}{}^\mu(\bar{x}_H)$$

$$\delta F_{\mu\nu} = F_{\mu\nu} - \bar{F}_{\mu\nu} = \partial_\mu \delta A_\nu - \partial_\nu \delta A_\mu$$

In summary, the EFT action for 1SF dynamics is

$$S_{\text{EFT}} = S_{\text{BF}} + S_{\text{recoil}}$$

$$S_{\text{BF}} = \int d^4x \left[-\frac{1}{4} \delta F_{\mu\nu} \delta F^{\mu\nu} - \delta A_\mu \bar{J}_L^\mu \right]$$

$$S_{\text{recoil}} = -\frac{z_H^2 m_H}{2} \int d\tau \dot{\bar{x}}_H^\alpha \delta F_{\alpha\mu}(\bar{x}_H) \frac{1}{\partial_\tau^2} \dot{\bar{x}}_H^\beta \delta F_{\beta}{}^\mu(\bar{x}_H)$$

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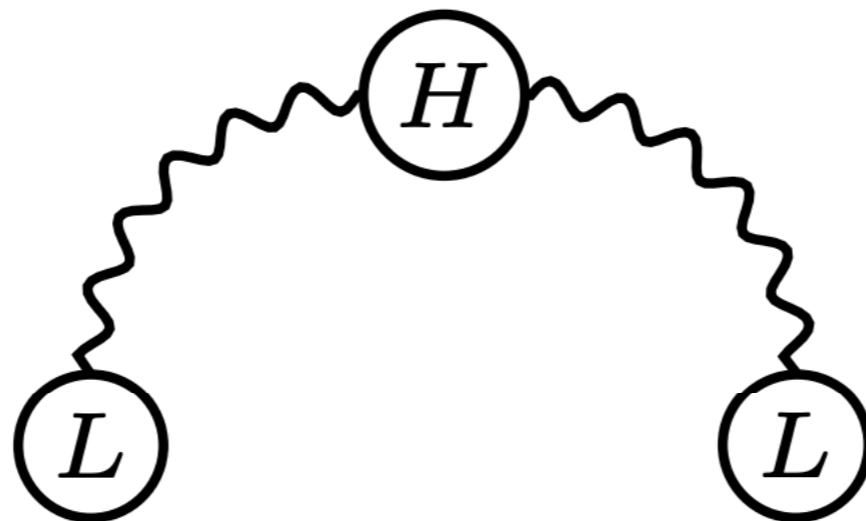
$$S_{\text{EFT}} = S_{\text{BF}} + S_{\text{recoil}}$$

$$S_{\text{BF}} = \int d^4x \left[\underbrace{\frac{1}{4} \delta F_{\mu\nu} \delta F^{\mu\nu}}_{\text{propagator}} - \delta A_{\mu} \underbrace{\tilde{J}^{\mu}}_{\text{source}} \right]$$

$$S_{\text{recoil}} = -\frac{z_H^2 m_H}{2} \int d\tau \dot{\bar{x}}_H^{\alpha} \delta F_{\alpha\mu}(\bar{x}_H) \underbrace{\frac{1}{\partial_{\tau}^2}}_{\text{recoil correction to propagator}} \dot{\bar{x}}_H^{\beta} \delta F_{\beta}{}^{\mu}(\bar{x}_H)$$

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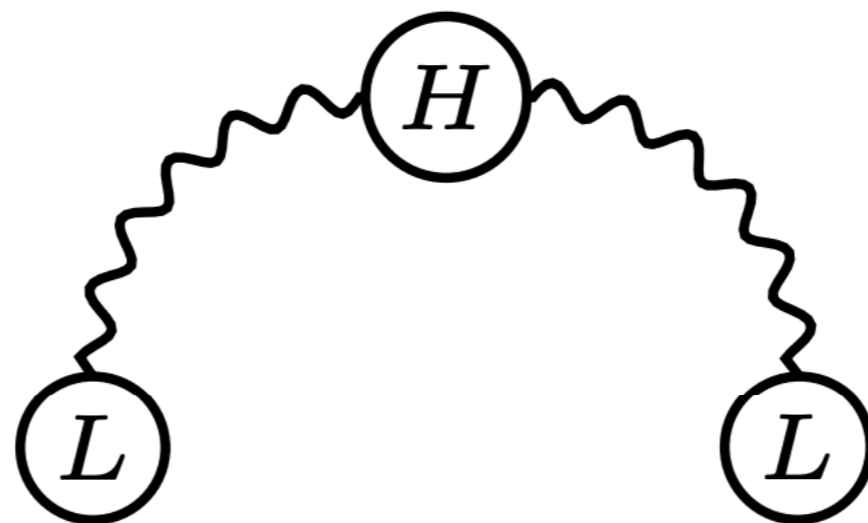
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1SF contribution to
conservative radial action

In summary, the EFT action for 1SF dynamics is

$$S_{\text{EFT}} = S_{\text{BF}} + S_{\text{recoil}}$$



all orders in α
from trajectory

1SF contribution to
conservative radial action

The EFT Feynman rules for EM are given by:

$$\text{propagator} \quad \text{~~~~~} = -\frac{i\eta_{\mu\nu}}{p^2}$$

$$\text{source} \quad \begin{array}{c} \text{~~~~~} \\ \textcircled{L} \end{array} = \lambda z_L m_H \int d\tau e^{-ip\bar{x}_L} \dot{\bar{x}}_L^\mu$$

$$\text{recoil} \quad \text{~~~~~} \textcircled{H} \text{~~~~~} = iz_H^2 m_H \frac{\delta(u_H p_1 + u_H p_2)}{(u_H p_1)(u_H p_2)} \mathcal{O}^{\alpha\mu_1}(u_H, p_1) \mathcal{O}_\alpha^{\mu_2}(u_H, p_2)$$

$$\mathcal{O}^{\alpha\mu}(u, p) = \eta^{\alpha\mu}(up) - u^\mu p^\alpha$$

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$$\text{propagator} = -\frac{i\eta_{\mu\nu}}{p^2}$$



source

$$= \lambda z_L m_H \int d\tau e^{-ip\bar{x}_L} \dot{\bar{x}}_L^\mu$$

probe trajectory
encodes integrand



$$\text{recoil} = iz_H^2 m_H \frac{\delta(u_H p_1 + u_H p_2)}{(u_H p_1)(u_H p_2)} \mathcal{O}^{\alpha\mu_1}(u_H, p_1) \mathcal{O}_\alpha^{\mu_2}(u_H, p_2)$$

recoil

$$\mathcal{O}^{\alpha\mu}(u, p) = \eta^{\alpha\mu}(up) - u^\mu p^\alpha$$

Starting from the EM light particle trajectory...

$$\bar{x}_L^\mu = \sum_{k=0}^{\infty} \bar{x}_k^\mu \quad \leftarrow \quad \begin{array}{l} \text{expansion in} \\ \text{fine structure} \\ \text{constant } \alpha \end{array}$$

Starting from the EM light particle trajectory...

$$\bar{x}_L^\mu = \sum_{k=0}^{\infty} \bar{x}_k^\mu$$

$$\hat{x}_0^\mu = \hat{b}^\mu + \frac{\hat{\tau}}{v} \check{u}_H^\mu + \frac{\hat{\tau} \sqrt{1-v^2}}{v} \check{u}_L^\mu$$

$$\hat{x}_1^\mu = \hat{\alpha} \left(a \sinh \hat{\tau} \check{u}_H^\mu - \frac{\hat{r}_0(\hat{\tau})}{v} \hat{b}^\mu \right)$$

$$\hat{x}_2^\mu = \hat{\alpha}^2 \left(\frac{a \sinh \hat{\tau}}{v \hat{r}_0(\hat{\tau})} \check{u}_H^\mu + \frac{\hat{\tau} - v^2 a \tan \hat{\tau}}{2v \sqrt{1-v^2}} \check{u}_L^\mu - \frac{2\hat{\tau} a \sinh \hat{\tau} + \hat{r}_0(\hat{\tau}) (\hat{\tau} v^2 a \tan \hat{\tau} + v^2 - 1)}{2v^2 \hat{r}_0(\hat{\tau})} \hat{b}^\mu \right)$$

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proper time
domain



$$\hat{x}_k^\mu = \bar{x}_k^\mu / b \quad \hat{\tau} = (\sigma^2 - 1)^{1/2} \tau / b \quad \hat{r}_0(\hat{\tau}) = \sqrt{1 + \hat{\tau}^2}$$

$$\check{u}_H^\mu = \frac{\sigma u_L - u_H}{\sigma^2 - 1} \quad \check{u}_L^\mu = \frac{\sigma u_H - u_L}{\sigma^2 - 1} \quad \hat{\alpha} = \alpha / b m_L (\sigma^2 - 1)^{1/2} \quad \sigma = \frac{1}{\sqrt{1-v^2}}$$

...then by applying simple identities such as

$$\text{asinh}(\hat{\tau}) = \frac{1}{\partial_{\hat{\tau}}} \left(\frac{1}{\hat{r}_0(\hat{\tau})} \right) \sim \frac{1}{\omega} \frac{1}{p^2}$$

yields EM integrands terms from trajectories:

$$\bar{x}_i^\mu = \frac{1}{\partial_{\hat{\tau}}^2} \left(\dots \frac{1}{\partial_{\hat{\tau}}^2} \left(\dots \frac{1}{\partial_{\hat{\tau}}^2} \left(\dots \frac{1}{\partial_{\hat{\tau}}^2} (\dots) \right) \right) \right)$$

...then by applying simple identities such as

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freedom here!

II. General Relativity

Now consider GR with a light + heavy particle,

$$S_{\text{GR}} = -\frac{1}{16\pi G} \int d^4x \sqrt{-g} R - \sum_{i=L,H} \frac{1}{2} m_i \int d\tau \dot{x}_i^\mu \dot{x}_i^\nu g_{\mu\nu}(x_i)$$

after gauge fixing the einbein to one. As before, mass ratio defines the SF expansion parameter:

$$\lambda = \frac{m_L}{m_H}$$

Expand about a $\mathcal{O}(\lambda^0)$ background and trajectory.

$$\begin{array}{ccc} \mathcal{O}(\lambda^0) & & \mathcal{O}(\lambda^0) \\ x_i^\mu = \bar{x}_i^\mu + \delta x_i^\mu & & g_{\mu\nu} = \bar{g}_{\mu\nu} + \delta g_{\mu\nu} \\ \nearrow \quad \nwarrow & & \nearrow \quad \nwarrow \\ \text{probe geodesic} \quad \text{deviation} & & \text{Schwarzschild} \quad \text{graviton} \end{array}$$

From full theory to background field theory:

$$S_{\text{GR}}[\delta g, \delta x_L, \delta x_H] \longrightarrow S_{\text{BF}}[\delta g, \delta x_L]$$

integrate out
heavy particle!

Now consider some explicit OSF backgrounds.

- heavy trajectory

$$\bar{x}_H^\mu(\tau) = u_H^\mu \tau$$

- metric

$$\bar{g}_{\mu\nu}(x) = f_+(r)^4 \eta_{\mu\nu} + \left[\frac{f_-(r)^2}{f_+(r)^2} - f_+(r)^4 \right] u_{H\mu} u_{H\nu}$$

$$f_{\pm}(r) = 1 \pm \frac{r_S}{4r}$$

- light trajectory

$$\bar{x}_L^\mu(\tau) \quad \text{s.t.} \quad \ddot{\bar{x}}_L^\mu(\tau) + \bar{\Gamma}^\mu_{\alpha\beta}(\bar{x}_L) \dot{\bar{x}}_L^\alpha \dot{\bar{x}}_L^\beta = 0$$

Perturbative terms from Schwarzschild metric:

$$\bar{g}_{\mu\nu}(x) = \eta_{\mu\nu} + \bar{\gamma}_{\mu\nu}(x)$$

$$\bar{\gamma}_{\mu\nu}(x) = \frac{r_S}{r}(\eta_{\mu\nu} - 2u_{H\mu}u_{H\nu}) + \frac{1}{8} \left(\frac{r_S}{r} \right)^2 (3\eta_{\mu\nu} + u_{H\mu}u_{H\nu}) + \dots$$

$$\begin{aligned} \tilde{\gamma}_{\mu\nu}(p) = & -\frac{4\pi r_S}{p^2}(\eta_{\mu\nu} - 2u_{H\mu}u_{H\nu})\delta(u_H p) \\ & -\frac{2\pi r_S^2}{\sqrt{-p^2}}(3\eta_{\mu\nu} + u_{H\mu}u_{H\nu})\delta(u_H p) + \dots \end{aligned}$$

Terms in the action relevant to $\mathcal{O}(\lambda^1)$ are given by:

$$S_{\text{GR}} = m_H \int d\tau \left[-\frac{1}{2} \delta \dot{x}_H^2 + \delta x_H^\rho \dot{\bar{x}}_H^\mu \dot{\bar{x}}_H^\nu \delta \Gamma_{\rho\mu\nu}(\bar{x}_H) + \dots \right]$$

$$+ \int d^4x \sqrt{-\bar{g}} \left[\mathcal{O}(\delta g^2) - \frac{1}{2} \delta g_{\mu\nu} \bar{T}_L^{\mu\nu} \right]$$

with $\mathcal{O}(\lambda^1)$ current $\bar{T}_L^{\mu\nu}(x) = m_L \int d\tau \frac{\delta^4(x - \bar{x}_L)}{\sqrt{-\bar{g}}} \dot{\bar{x}}_L^\mu \dot{\bar{x}}_L^\nu$

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Gaussian in δx_H

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$$+ \int d^4x \sqrt{-\bar{g}} \left[\mathcal{O}(\delta g^2) - \frac{1}{2} \delta g_{\mu\nu} \bar{T}^{\mu\nu} \right]$$

with $\mathcal{O}(\lambda^1)$ current $\bar{T}_L^{\mu\nu}(x) = m_L \int d\tau \frac{\delta^4(x - \bar{x}_L)}{\sqrt{-\bar{g}}} \dot{x}_L^\mu \dot{x}_L^\nu$

Integrating out δx_H yields the “recoil operator”.

$$m_H \int d\tau \left[-\frac{1}{2} \delta \dot{x}_H^2 + \delta x_H^\rho \dot{\bar{x}}_H^\mu \dot{\bar{x}}_H^\nu \delta \Gamma_{\rho\mu\nu}(\bar{x}_H) \right]$$



$$S_{\text{recoil}} = -\frac{1}{2} m_H \int d\tau \dot{\bar{x}}_H^\alpha \dot{\bar{x}}_H^\beta \delta \Gamma_{\alpha\beta}^\mu(\bar{x}_H) \frac{1}{\partial_\tau^2} \dot{\bar{x}}_H^\gamma \dot{\bar{x}}_H^\delta \delta \Gamma_{\mu\gamma\delta}(\bar{x}_H)$$

$$\delta \Gamma_{\rho\mu\nu} = \Gamma_{\rho\mu\nu} - \bar{\Gamma}_{\rho\mu\nu} = \frac{1}{2} (\partial_\mu \delta g_{\nu\rho} + \partial_\nu \delta g_{\mu\rho} - \partial_\rho \delta g_{\mu\nu})$$

In summary, the EFT action for 1SF dynamics is

$$S_{\text{EFT}} = S_{\text{BF}} + S_{\text{recoil}}$$

$$S_{\text{BF}} = \int d^4x \sqrt{-\bar{g}} \left[\mathcal{O}(\delta g^2) - \frac{1}{2} \delta g_{\mu\nu} \bar{T}^{\mu\nu} \right]$$

$$S_{\text{recoil}} = -\frac{1}{2} m_H \int d\tau \dot{\bar{x}}_H^\alpha \dot{\bar{x}}_H^\beta \delta\Gamma^\mu_{\alpha\beta}(\bar{x}_H) \frac{1}{\partial_\tau^2} \dot{\bar{x}}_H^\gamma \dot{\bar{x}}_H^\delta \delta\Gamma_{\mu\gamma\delta}(\bar{x}_H)$$

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recall that
 $\bar{g}_{\mu\nu}(\bar{x}_H) = \eta_{\mu\nu}$
 in dim reg

$$S_{\text{recoil}} = -\frac{1}{2} m_H \int d\tau \dot{\bar{x}}_H^\alpha \dot{\bar{x}}_H^\beta \delta\Gamma^\mu_{\alpha\beta}(\bar{x}_H) \frac{1}{\partial_\tau^2} \dot{\bar{x}}_H^\gamma \dot{\bar{x}}_H^\delta \delta\Gamma_{\mu\gamma\delta}(\bar{x}_H)$$

In summary, the EFT action for 1SF dynamics is

$$S_{\text{EFT}} = S_{\text{BF}} + S_{\text{recoil}}$$

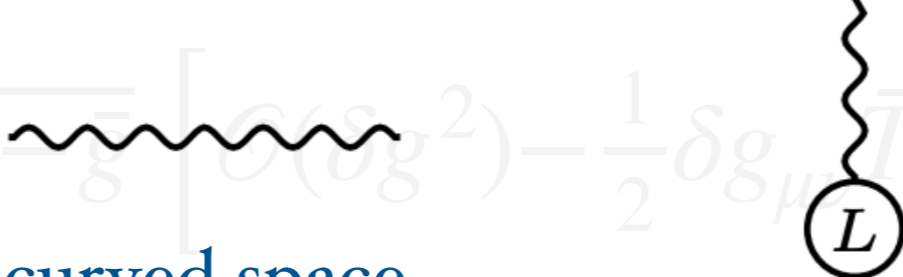
$$S_{\text{BF}} = \int d^4x \sqrt{-\bar{g}} \left[\mathcal{O}(\delta g^2) - \frac{1}{2} \delta g_{\mu\nu} \bar{T}^{\mu\nu} \right]$$

$$S_{\text{recoil}} = -\frac{1}{2} m_H \int d\tau \dot{\bar{x}}_H^\alpha \dot{\bar{x}}_H^\beta \delta\Gamma^\mu_{\alpha\beta}(\bar{x}_H) \frac{1}{\partial_\tau^2} \dot{\bar{x}}_H^\gamma \dot{\bar{x}}_H^\delta \delta\Gamma_{\mu\gamma\delta}(\bar{x}_H)$$

In summary, the EFT action for 1SF dynamics is


$$S_{\text{EFT}} = S_{\text{BF}} + S_{\text{recoil}}$$

$$S_{\text{BF}} = \int d^4x \sqrt{-\bar{g}} \left[\mathcal{L}(\partial g^2) - \frac{1}{2} \delta g_{\mu\nu} \overset{\text{source}}{\text{T}}^{\mu\nu} \right]$$



 curved space propagator

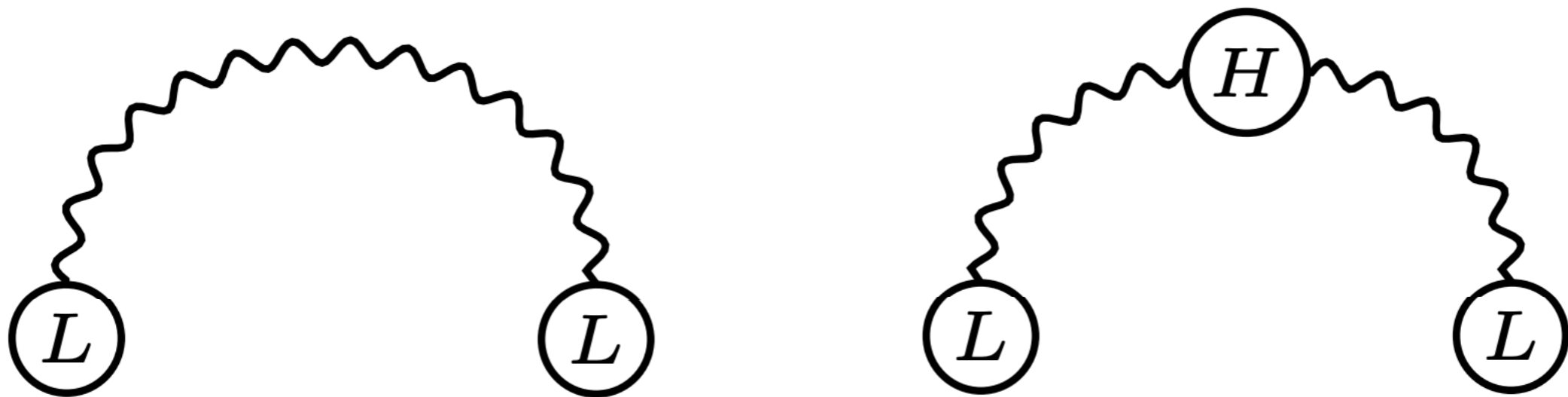
$$S_{\text{recoil}} = -\frac{1}{2} m_H \int d\tau \dot{\bar{x}}_H^\alpha \dot{\bar{x}}_H^\beta \delta \Gamma_{\alpha\beta}^\mu \overset{1}{\underset{\tau}{\text{H}}} \overset{\delta}{\text{H}} \delta \Gamma_{\mu\gamma\delta}(\bar{x}_H)$$



 recoil correction to propagator

In summary, the EFT action for 1SF dynamics is

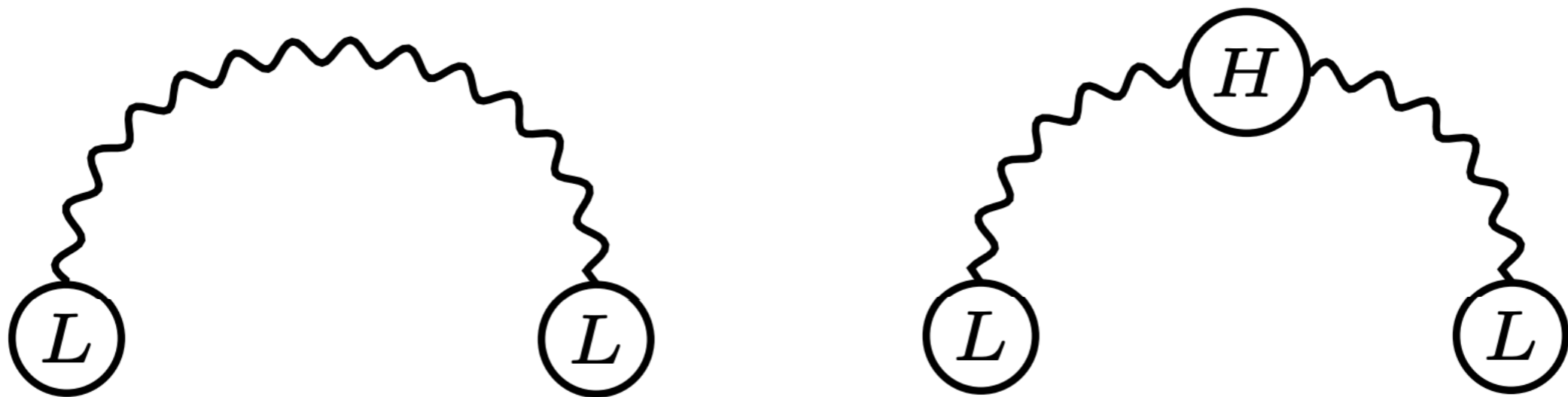
$$S_{\text{EFT}} = S_{\text{BF}} + S_{\text{recoil}}$$



1SF diagrams for
conservative radial action

In summary, the EFT action for 1SF dynamics is

$$S_{\text{EFT}} = S_{\text{BF}} + S_{\text{recoil}}$$

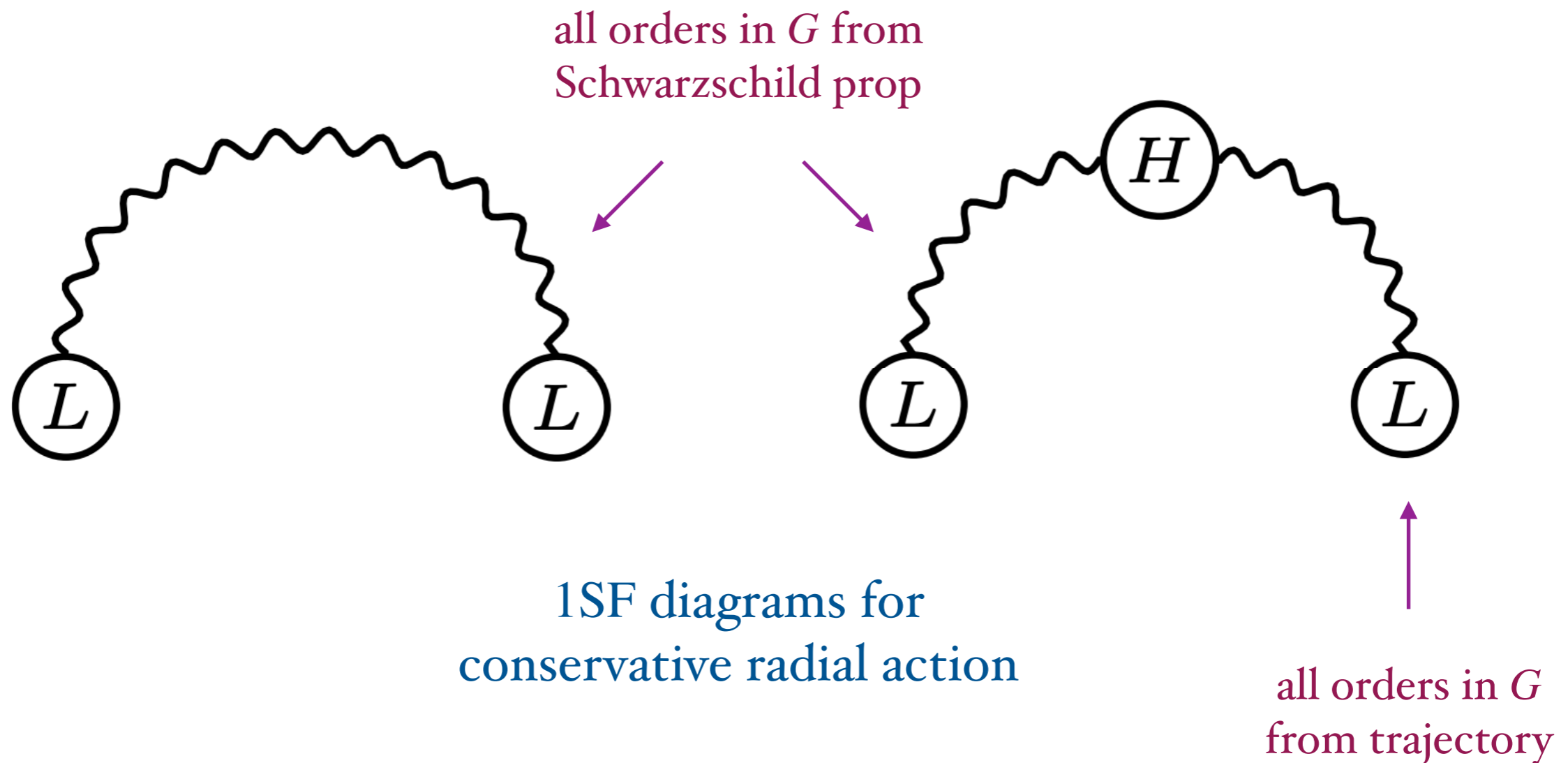


1SF diagrams for
conservative radial action

↑
all orders in G
from trajectory

In summary, the EFT action for 1SF dynamics is

$$S_{\text{EFT}} = S_{\text{BF}} + S_{\text{recoil}}$$



The EFT Feynman rules for GR are given by:

$$\begin{array}{c} \text{~~~~~} \\ \text{propagator} \end{array} = \frac{32\pi i G}{p^2} \left(\frac{\eta_{\mu\rho}\eta_{\nu\sigma} + \eta_{\mu\sigma}\eta_{\nu\rho}}{2} - \frac{\eta_{\mu\nu}\eta_{\rho\sigma}}{2} \right) + \text{PM corrections}$$

$$\begin{array}{c} \text{~~~~~} \\ \text{L} \\ \text{source} \end{array} = \lambda m_H \int d\tau e^{-ip\bar{x}_L} \dot{\bar{x}}_L^\mu \dot{\bar{x}}_L^\nu$$

$$\begin{array}{c} \text{~~~~~} \\ \text{H} \\ \text{recoil} \end{array} = \frac{i m_H \delta(u_H p_1 + u_H p_2)}{2 (u_H p_1)(u_H p_2)} \mathcal{O}^{\alpha\mu_1\nu_1}(u_H, p_1) \mathcal{O}_\alpha{}^{\mu_2\nu_2}(u_H, p_2)$$

$$\mathcal{O}^{\alpha\mu\nu}(u, p) = \frac{1}{2} \left((u^\mu \eta^{\nu\alpha} + u^\nu \eta^{\mu\alpha})(u p) - u^\mu u^\nu p^\alpha \right)$$

Loop integrands via GR light particle trajectory:

$$\bar{x}_L^\mu = \sum_{k=0}^{\infty} \bar{x}_k^\mu \quad \leftarrow \begin{array}{l} \text{expansion in} \\ \text{Schwarzschild} \\ \text{radius } r_S \end{array}$$

Loop integrands via GR light particle trajectory:

$$\bar{x}_L^\mu = \sum_{k=0}^{\infty} \bar{x}_k^\mu$$

$$\hat{x}_0^\mu = \hat{b}^\mu + \frac{\hat{\tau}}{v} \check{u}_H^\mu + \frac{\hat{\tau} \sqrt{1-v^2}}{v} \check{u}_L^\mu$$

$$\hat{x}_1^\mu = \frac{r_S}{bv^2} \left(-\frac{(v^2+1)\hat{r}_0(\hat{\tau})}{2} \hat{b}^\mu + v \operatorname{asinh} \hat{\tau} \check{u}_H^\mu + \frac{v(v^2+1) \operatorname{asinh} \hat{\tau}}{2\sqrt{1-v^2}} \check{u}_L^\mu \right)$$

$$\begin{aligned} \hat{x}_2^\mu = & \left(\frac{r_S}{bv^2} \right)^2 \left(\left[\frac{3v^3 \operatorname{atan} \hat{\tau}}{2} - \frac{v(v^2-1) \operatorname{asinh} \hat{\tau}}{2\hat{r}_0(\hat{\tau})} \right] \check{u}_H^\mu \right. \\ & + \left[\frac{\sqrt{1-v^2} v (v^2+1) \operatorname{asinh} \hat{\tau}}{4\hat{r}_0(\hat{\tau})} + \frac{2\hat{\tau} v (v^2+1)^2 + 3(v^2+4)v^3 \operatorname{atan} \hat{\tau}}{16\sqrt{1-v^2}} \right] \check{u}_L^\mu \\ & \left. + \left[\frac{\hat{\tau} (v^4-1) \operatorname{asinh} \hat{\tau}}{4\hat{r}_0(\hat{\tau})} - \frac{3\hat{\tau} (v^2+4)v^2 \operatorname{atan} \hat{\tau} + v^4 + 8v^2 - 2}{16} \right] \hat{b}^\mu \right) \end{aligned}$$

Loop integrands via GR light particle trajectory:

$$\bar{x}_L^\mu = \sum_{k=0}^{\infty} \bar{x}_k^\mu$$

rewrite as worldline propagators

$$\operatorname{asinh} \hat{\tau} = \frac{1}{\partial_{\hat{\tau}}} \partial_{\hat{\tau}} \operatorname{asinh} \hat{\tau} = \frac{1}{\partial_{\hat{\tau}}} \left(\frac{1}{(1 + \hat{\tau}^2)^{1/2}} \right)$$

$$\hat{x}_0^\mu = \hat{b}^\mu + \frac{\hat{\tau}}{v} \check{u}_H^\mu + \frac{\hat{\tau} \sqrt{1 - v^2}}{v} \check{u}_L^\mu$$

$$\hat{x}_1^\mu = \frac{r_S}{bv^2} \left(-\frac{(v^2 + 1) \hat{r}_0(\hat{\tau})}{2} \hat{b}^\mu + v \operatorname{asinh} \hat{\tau} \check{u}_H^\mu + \frac{v(v^2 + 1) \operatorname{asinh} \hat{\tau}}{2\sqrt{1 - v^2}} \check{u}_L^\mu \right)$$

$$\begin{aligned} \hat{x}_2^\mu = & \left(\frac{r_S}{bv^2} \right)^2 \left(\left[\frac{3v^3 \operatorname{atan} \hat{\tau}}{2} - \frac{v(v^2 - 1) \operatorname{asinh} \hat{\tau}}{2\hat{r}_0(\hat{\tau})} \right] \check{u}_H^\mu \right. \\ & + \left[\frac{\sqrt{1 - v^2} v (v^2 + 1) \operatorname{asinh} \hat{\tau}}{4\hat{r}_0(\hat{\tau})} + \frac{2\hat{\tau} v (v^2 + 1)^2 + 3(v^2 + 4) v^3 \operatorname{atan} \hat{\tau}}{16\sqrt{1 - v^2}} \right] \check{u}_L^\mu \\ & \left. + \left[\frac{\hat{\tau} (v^4 - 1) \operatorname{asinh} \hat{\tau}}{4\hat{r}_0(\hat{\tau})} - \frac{3\hat{\tau} (v^2 + 4) v^2 \operatorname{atan} \hat{\tau} + v^4 + 8v^2 - 2}{16} \right] \hat{b}^\mu \right) \end{aligned}$$

III. Perturbative Calculations

We use this EFT to compute old and new results for the conservative radial action in EM and GR.

1a) $\mathcal{O}(\alpha^3)$ in EM (electric charges)

1b) $\mathcal{O}(\alpha^3)$ in EM (dyons)

2a) $\mathcal{O}(G^3)$ in GR (spinless masses)

2b) $\mathcal{O}(G^3)$ in GR + ϕ + A_μ (spinless masses)

2c) $\mathcal{O}(G^3)$ in GR + EM (RN, DCSSG, EGB)

We use this EFT to compute old and new results for the conservative radial action in EM and GR.

1a) $\mathcal{O}(\alpha^3)$ in EM (electric charges) ← check

1b) $\mathcal{O}(\alpha^3)$ in EM (dyons)

2a) $\mathcal{O}(G^3)$ in GR (spinless masses) ← check

2b) $\mathcal{O}(G^3)$ in GR + ϕ + A_μ (spinless masses)

2c) $\mathcal{O}(G^3)$ in GR + EM (RN, DCSSG, EGB)

We use this EFT to compute old and new results for the conservative radial action in EM and GR.

1a) $\mathcal{O}(\alpha^3)$ in EM (electric charges)

1b) $\mathcal{O}(\alpha^3)$ in EM (dyons) ← new!

2a) $\mathcal{O}(G^3)$ in GR (spinless masses)

2b) $\mathcal{O}(G^3)$ in GR + ϕ + A_μ (spinless masses) ← new!

2c) $\mathcal{O}(G^3)$ in GR + EM (RN, DCSSG, EGB) ← new!

Example: $\mathcal{O}(\alpha^3)$ in EM (electric charges)

$$I_{\text{EM}}^{(i,j)} = \lambda^i m_L r_c \left(\frac{r_c}{b}\right)^{j-1} \mathcal{I}_{\text{EM}}^{(i,j)}(\sigma)$$

$$\begin{aligned} I_{\text{EM}}^{(1,2)} &= \frac{\lambda^2}{2} z_L^2 z_H^2 m_H^3 \int_{p_1, p_2} \frac{e^{-i(p_1+p_2)b} \delta(u_H(p_1+p_2)) \delta(u_L p_1) \delta(u_L p_2)}{p_1^2 p_2^2} \left(1 + \frac{\sigma^2(p_1 p_2)}{(u_H p_1)(u_H p_2)}\right) \\ &= \lambda m_L r_c \frac{r_c}{b} \frac{\pi}{2\sqrt{\sigma^2-1}}. \end{aligned}$$

$$\begin{aligned} I_{\text{EM}}^{(1,3)} &= -(m_H m_L)^2 (z_H z_L)^3 \int_{q, p_1, p_2, p_3} e^{iqb} \delta(u_H q) \delta(u_L q) \frac{\delta(q-p_1-p_2-p_3) \delta(u_H p_2) \delta(u_L p_1)}{p_1^2 p_2^2 p_3^2 (u_H p_1)^2 (u_L p_2)^2} \\ &\quad \times \left(-(p_1 p_3)(p_2 p_3) \sigma^3 - \frac{1}{2} q^2 (u_H p_1)(u_L p_2) \sigma^2 + (p_2 p_3)(u_H p_1)^2 \sigma + (p_1 p_3)(u_L p_2)^2 \sigma \right. \\ &\quad \left. + (u_L p_2)(u_H p_1)^3 + (u_H p_1)(u_H p_2)^3 \right) \\ &= -\lambda m_L r_c \left(\frac{r_c}{b}\right)^2 \frac{2(\sigma^4 - 3\sigma^2 + 3)}{3(\sigma^2 - 1)^{5/2}}, \end{aligned}$$

Example: $\mathcal{O}(\alpha^3)$ in EM (electric charges)

$$I_{\text{EM}}^{(i,j)} = \lambda^i m_L r_c \left(\frac{r_c}{b}\right)^{j-1} \mathcal{I}_{\text{EM}}^{(i,j)}(\sigma) \quad \longleftarrow \quad (i)\text{-SF and } (j)\text{-PM}$$

$$\begin{aligned} I_{\text{EM}}^{(1,2)} &= \frac{\lambda^2}{2} z_L^2 z_H^2 m_H^3 \int_{p_1, p_2} \frac{e^{-i(p_1+p_2)b} \delta(u_H(p_1+p_2)) \delta(u_L p_1) \delta(u_L p_2)}{p_1^2 p_2^2} \left(1 + \frac{\sigma^2(p_1 p_2)}{(u_H p_1)(u_H p_2)}\right) \\ &= \lambda m_L r_c \frac{r_c}{b} \frac{\pi}{2\sqrt{\sigma^2-1}}. \end{aligned}$$

$$\begin{aligned} I_{\text{EM}}^{(1,3)} &= -(m_H m_L)^2 (z_H z_L)^3 \int_{q, p_1, p_2, p_3} e^{iqb} \delta(u_H q) \delta(u_L q) \frac{\delta(q-p_1-p_2-p_3) \delta(u_H p_2) \delta(u_L p_1)}{p_1^2 p_2^2 p_3^2 (u_H p_1)^2 (u_L p_2)^2} \\ &\quad \times \left(-(p_1 p_3)(p_2 p_3) \sigma^3 - \frac{1}{2} q^2 (u_H p_1)(u_L p_2) \sigma^2 + (p_2 p_3)(u_H p_1)^2 \sigma + (p_1 p_3)(u_L p_2)^2 \sigma \right. \\ &\quad \left. + (u_L p_2)(u_H p_1)^3 + (u_H p_1)(u_H p_2)^3 \right) \\ &= -\lambda m_L r_c \left(\frac{r_c}{b}\right)^2 \frac{2(\sigma^4 - 3\sigma^2 + 3)}{3(\sigma^2 - 1)^{5/2}}, \end{aligned}$$

Example: $\mathcal{O}(G^3)$ in GR (spinless masses)

$$I_{\text{GR}}^{(i,j)} = \lambda^i m_L r_S \left(\frac{r_S}{b}\right)^{j-1} \mathcal{I}_{\text{GR}}^{(i,j)}(\sigma)$$

$$I_{\text{GR}}^{(1,2)} = \lambda m_L r_S \frac{r_S}{b} \frac{3\pi(5\sigma^2 - 1)}{4\sqrt{\sigma^2 - 1}}$$

$$I_{\text{GR}}^{(1,3)} = \lambda m_L r_S \left(\frac{r_S}{b}\right)^2 \left(\frac{\sigma(36\sigma^6 - 114\sigma^4 + 132\sigma^2 - 55)}{12(\sigma^2 - 1)^{5/2}} - \frac{(4\sigma^4 - 12\sigma^2 - 3) \operatorname{arccosh} \sigma}{2(\sigma^2 - 1)} \right)$$

In this EFT it is relatively easy to add additional fields to GR, including spectator scalars / vector matter, or gravitational axion / dilaton.

Example: $\mathcal{O}(G^3)$ in GR + ϕ + A_μ (spinless masses)

$$S_{\text{matt}} = \int d^4x \sqrt{-g} \left[\frac{1}{2} \nabla \phi^2 + \frac{1}{2} \xi R \phi^2 - \frac{1}{4} F^2 \right] - m_L \int d\tau \left[y_L \phi(x_L) + z_L A_\mu(x_L) \dot{x}_L^\mu \right]$$

$$I_{\text{scalar}}^{(1,2)} = -\lambda m_L r_S \left(\frac{r_\Phi}{b} \right) \left(\frac{\pi \sigma^2 - 1 + 4\xi}{8 \sqrt{\sigma^2 - 1}} \right)$$

$$I_{\text{scalar}}^{(1,3)} = -\lambda m_L r_S \left(\frac{r_S r_\Phi}{b^2} \right) \frac{\sigma (2\sigma^4 - \sigma^2 - 1 + \xi (6\sigma^2 - 3))}{6 (\sigma^2 - 1)^{3/2}}$$

$$I_{\text{vector}}^{(1,2)} = -\lambda m_L r_S \left(\frac{r_A}{b} \right) \left(\frac{\pi (3\sigma^2 - 1)}{8 \sqrt{\sigma^2 - 1}} \right)$$

$$I_{\text{vector}}^{(1,3)} = -\lambda m_L r_S \left(\frac{r_S r_A}{b^2} \right) \left(\frac{\sigma (8\sigma^4 - 28\sigma^2 + 23)}{12 (\sigma^2 - 1)^{3/2}} + \frac{(2\sigma^2 + 1) \operatorname{arccosh} \sigma}{(\sigma^2 - 1)} \right)$$

Conclusions

We derive an EFT for SF corrections to probe dynamics by integrating out a heavy particle.

Using Schwarzschild and geodesic trajectories, we mine any order in G data for loop integrands.

We perform several old and new calculations of the conservative radial action in EM and GR.

Thank You!