

Fast numerical methods for cosmological inference from LSS

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Matching early to late in simulations...

Early physics:

- GR effects (horizon+rel. species+aniso-stress)
- multi-species (CDM+baryon+photons+neutrinos)
- photon-baryon coupling + recombination
- perturbative quantity: δ and θ

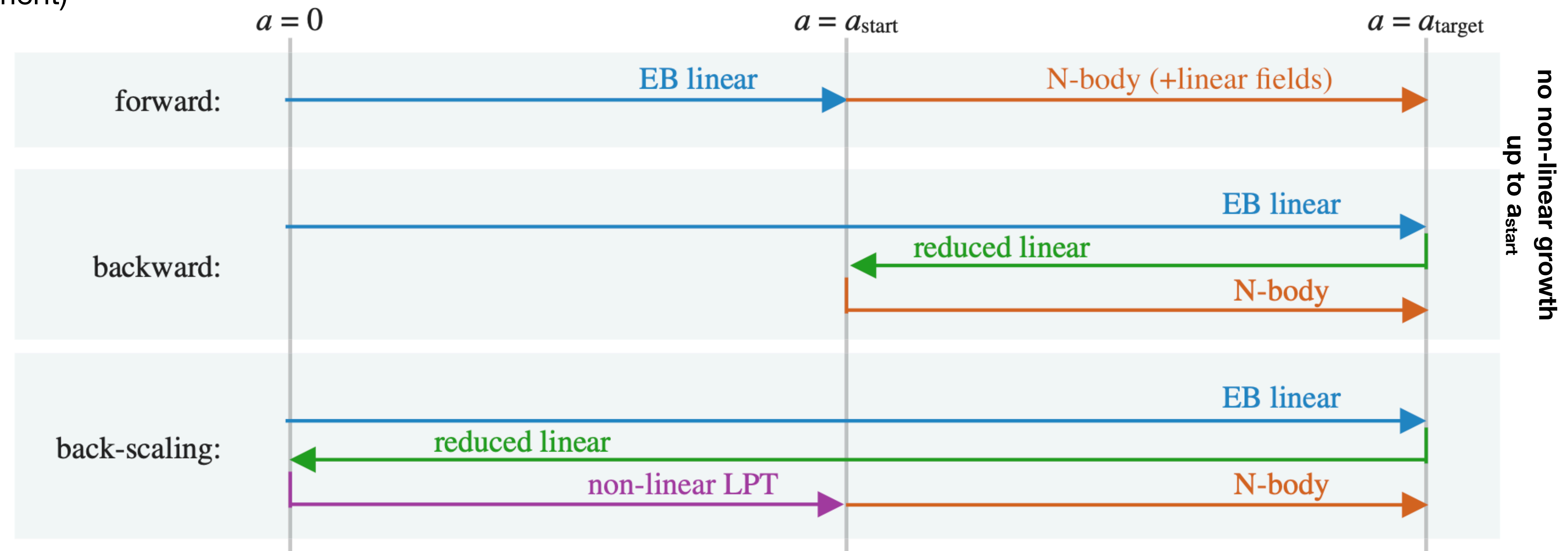
Late physics:

- Newtonian gravity + small corrections
- mostly interested in mass distribution, CDM+baryons
- non-linear growth
- perturbative quantity: ψ (displacement)

Eulerian

Lagrangian

choices in the literature:



Newtonian Gravitational Evolution: Vlasov-Poisson

Consider the (on-shell) phase-space density of non-interacting massive particles $f(x, v, t)$

Evolution given by non-relativistic limit of the relativistic Liouville equation = **Vlasov-Poisson**, assume config space $\mathcal{C} = \mathbb{T}^3$

$$d_t f = \partial_t f + \frac{v}{a^2} \cdot \nabla_x f + \frac{g}{a} \cdot \nabla_p f = 0 \qquad g := -\frac{3\Omega_m}{2} \nabla_x \nabla_x^{-2} (n - 1)$$

$$n = \int_{\mathbb{R}^3} d^3 p f \qquad \int_{\mathcal{C}} d^3 x n = 1$$

can be solved by **method of characteristics**:

Consider 1-parameter families of curves $\mathbf{X}(t)$, $\mathbf{V}(t)$ (the characteristics)

Study evolution of f along the characteristics

$$f_c : t \mapsto f(\mathbf{X}(t), \mathbf{V}(t), t) \quad \text{given} \quad \mathbf{X}(t = t_0) =: \mathbf{X}_0, \quad \mathbf{V}(t = t_0) =: \mathbf{V}_0$$

by chain rule

$$\frac{df_c}{dt} = \frac{\partial f_c}{\partial t} + \nabla_x f_c|_{x=\mathbf{X}(t)} \cdot \frac{d\mathbf{X}}{dt} + \nabla_v f_c|_{v=\mathbf{V}(t)} \cdot \frac{d\mathbf{V}}{dt}$$

comparing with VP

$$\frac{df_c}{dt} = 0 \quad \text{iff} \quad \begin{cases} \dot{\mathbf{X}}(t) = a^{-2} \mathbf{V}(t) \\ \dot{\mathbf{V}}(t) = -\nabla_x \phi|_{x=\mathbf{X}(t)} \end{cases}$$

has Hamiltonian

$$\mathcal{H} = \frac{1}{2} \frac{V^2}{a^2} + \phi$$

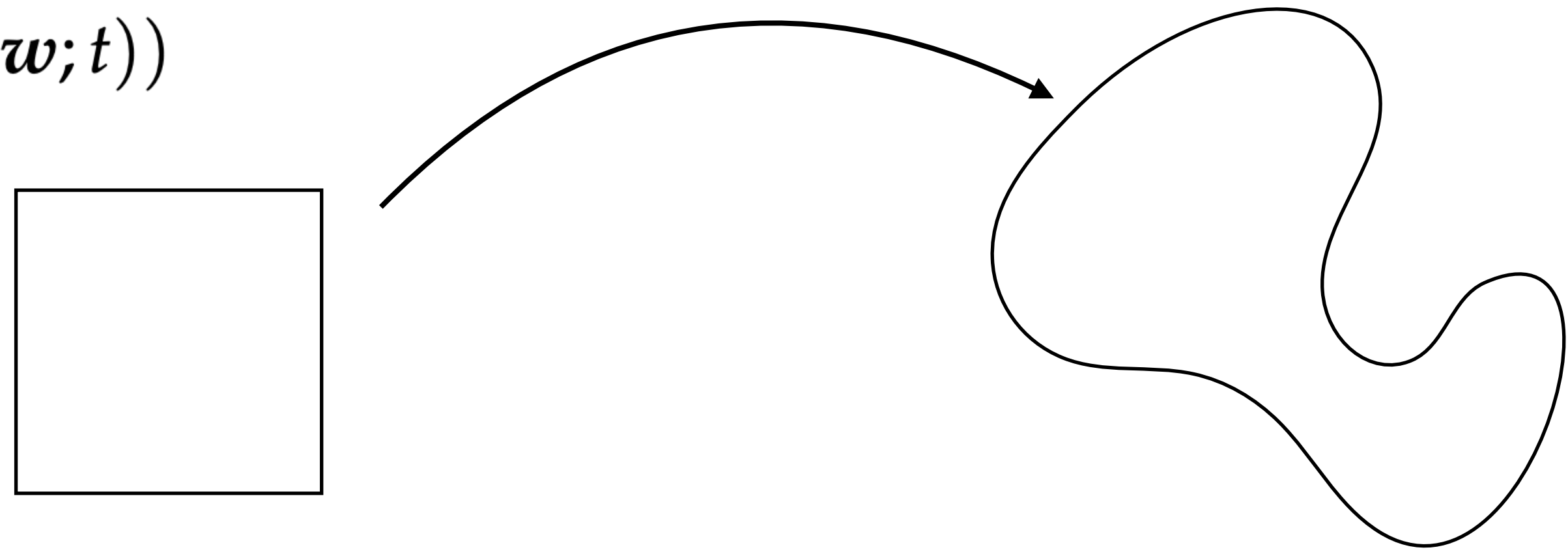
$$\nabla_x \phi = -\frac{g}{a}$$

The fluid and the N-body model

Lagrangian description, evolution of fluid element

Lagrangian map of the \mathbb{R}^{2d} phase space onto itself

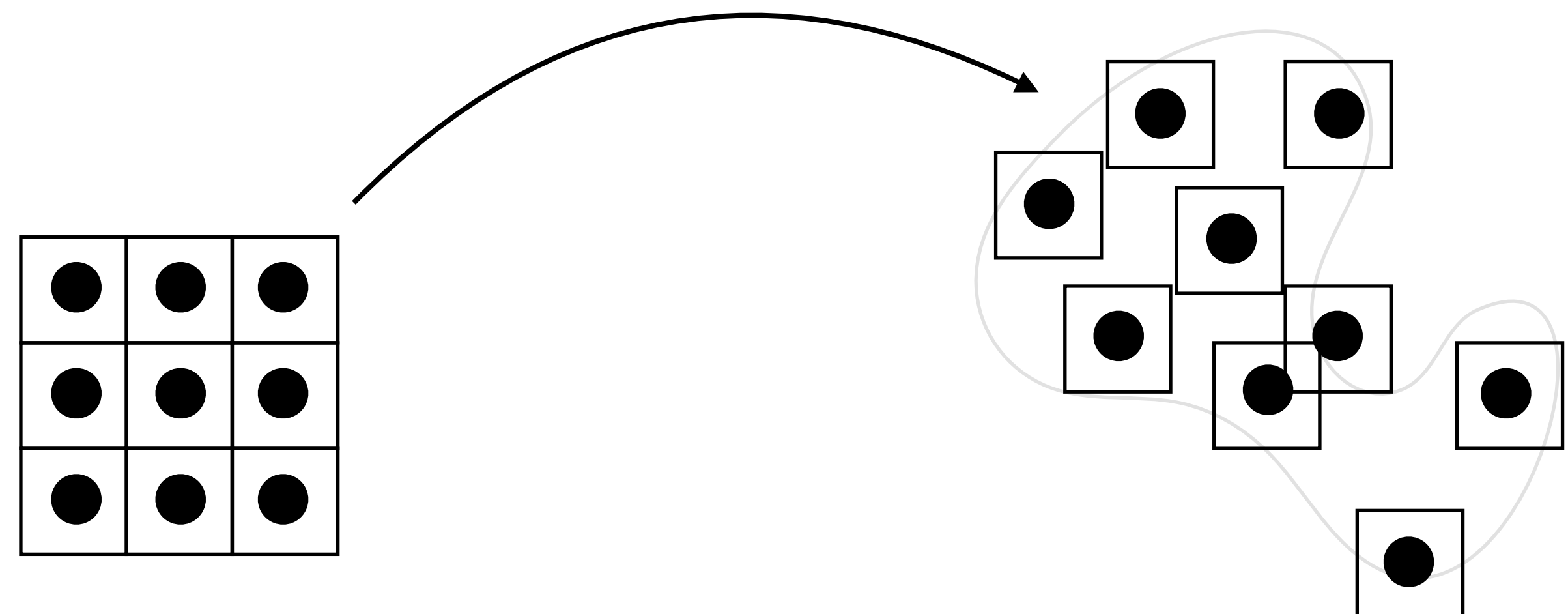
$$\mathbb{R}^{2d} \rightarrow \mathbb{R}^{2d} : (q, w) \mapsto (x(q, w; t), v(q, w; t))$$



The N-body approximation:

follow only N characteristics, use them to reconstruct the density field that sources the Poisson equation

$$\mathbb{R}^{N2d} \rightarrow \mathbb{R}^{N2d} : (Q_i, W_i) \mapsto (X_i(t), V_i(t)), \quad i = 1, \dots, N$$

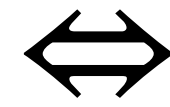


Non-Linear Evolution of Fluctuations

Cold Dark Matter lives on Lagrangian submanifold

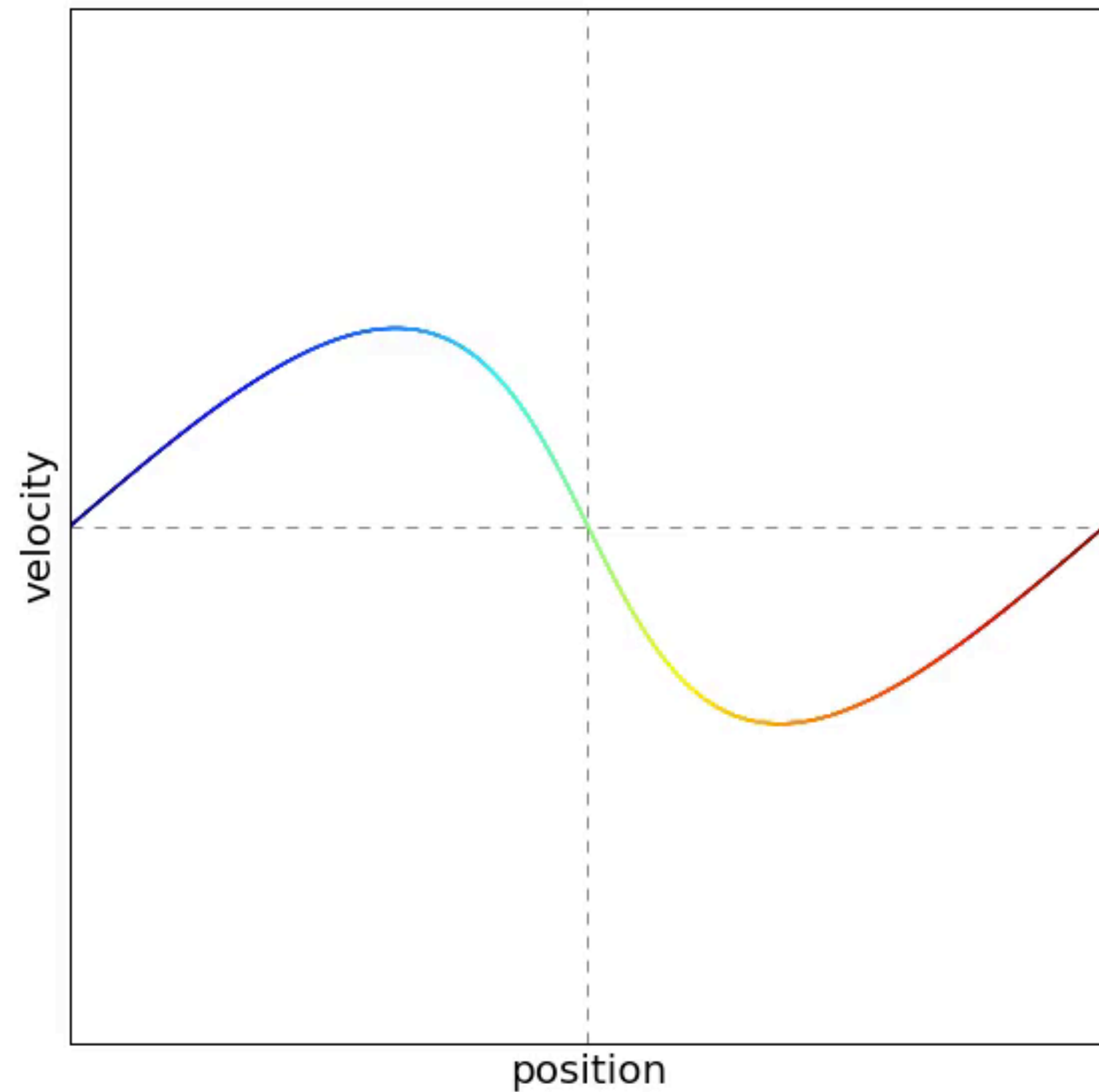
Solve Vlasov-Poisson on submanifold characteristics $(q, t) \mapsto (X_q(t), V_q(t))$,

$$\frac{\partial f}{\partial t} + \frac{v}{a^2} \cdot \nabla_x f - \nabla_x \phi \cdot \nabla_v f = 0$$

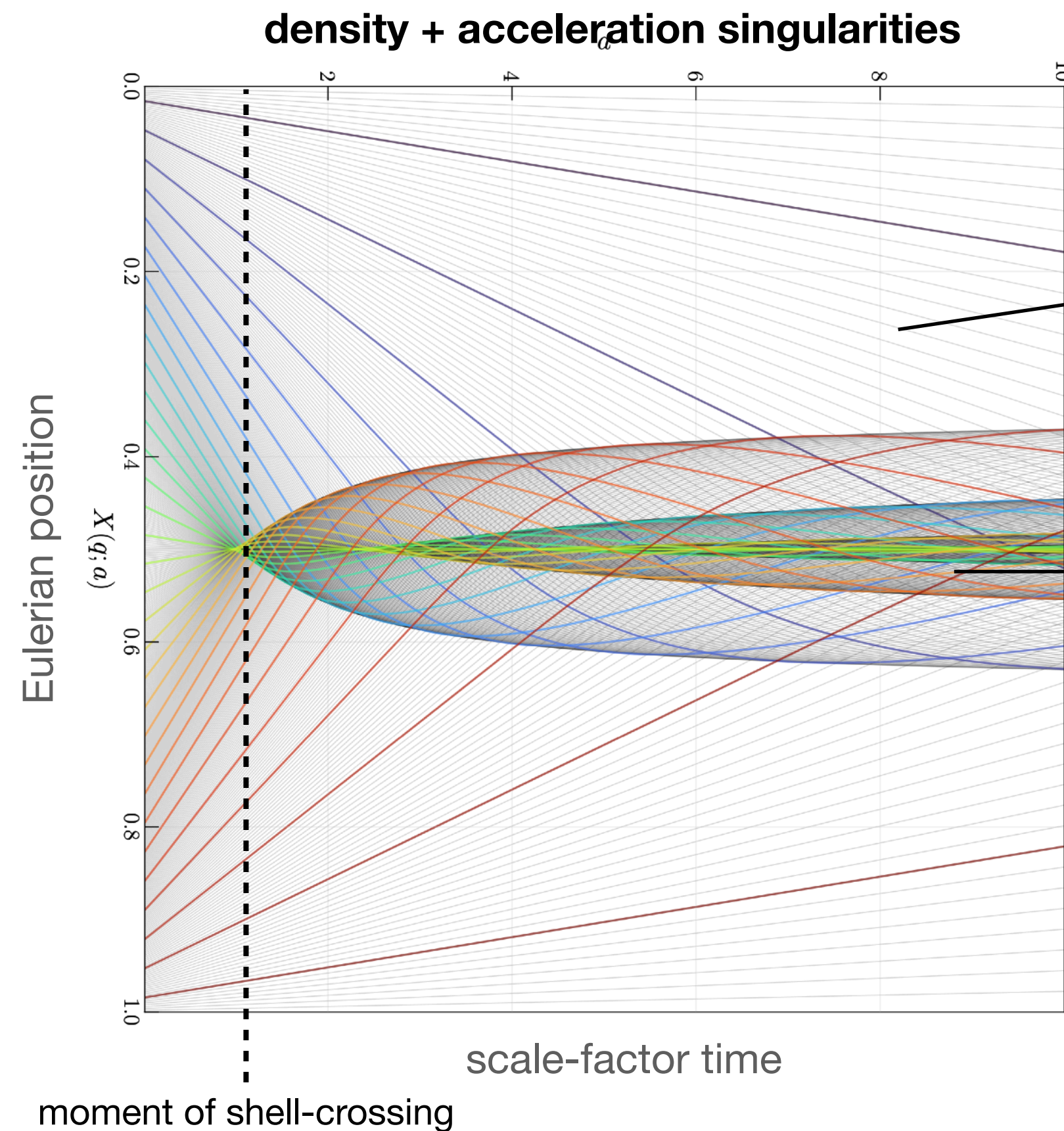


$$X_q'' + \mathcal{H}X_q' = -\nabla\phi(X_q)$$

entering the multi-stream region is non-analytic (only finitely many bounded derivatives)



1D singularities: Rampf, Frisch & OH (2021)



monokinetic, single-valued (analytic treatment possible)

multikinetic, multi-valued (simulations, EFTs [by integrating out])

Zeldovich (1970) solution (straight lines) is exact prior to shell-crossing and outside shell-crossed regions

Lagrangian Perturbation Theory

(for single fluid with cold initial data)

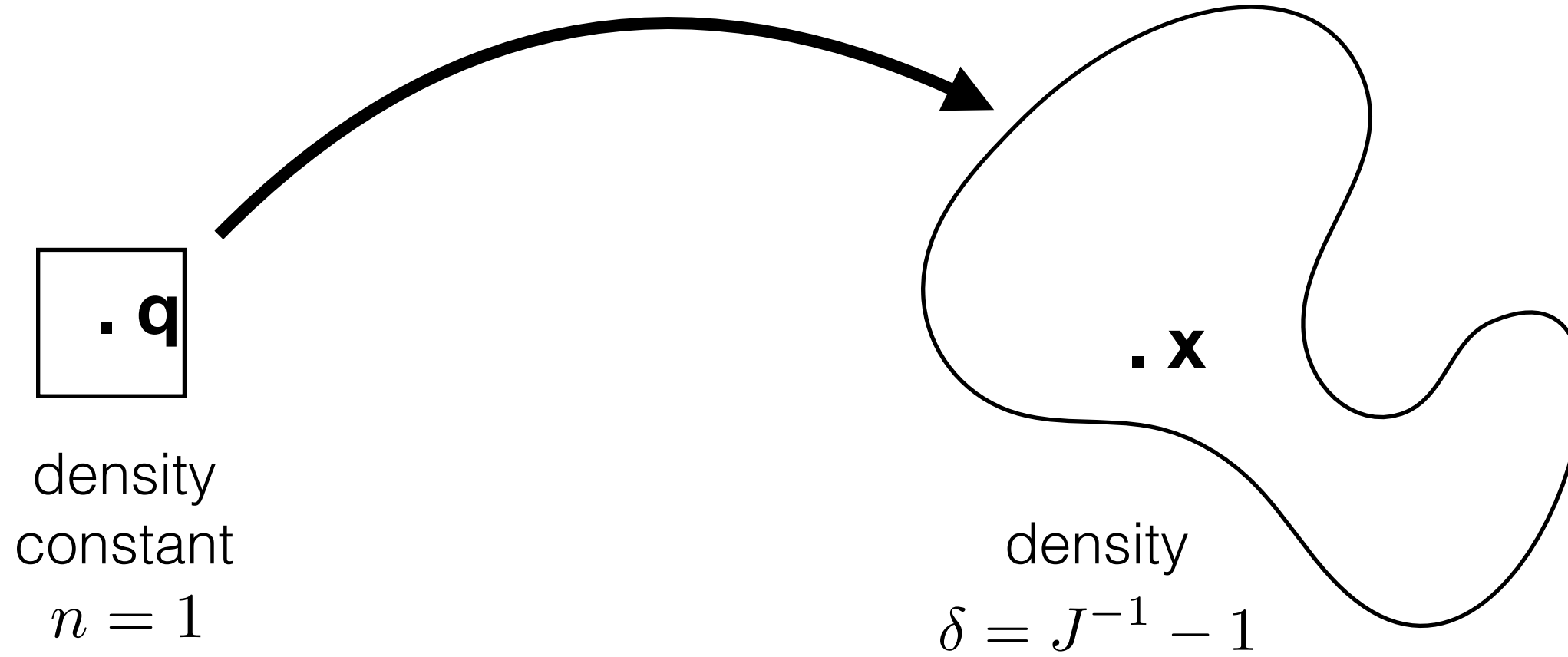
Lagrangian map

$$\mathbf{X}_q(t) = \mathbf{q} + \Psi(\mathbf{q}, t)$$

Overdensity given by Jacobian

$$\delta(\mathbf{x}, t) = \frac{1}{J(\mathbf{q}, t)} - 1$$

$$J := \det \nabla_{\mathbf{q}} \otimes \mathbf{X}_q$$



$$\text{LPT eq: } J (\delta_{ij} + \Psi_{i,j})^{-1} (\Psi''_{i,j} + \mathcal{H} \Psi'_{i,j}) = \frac{3}{2} \mathcal{H}^2 \Omega_m (J - 1)$$

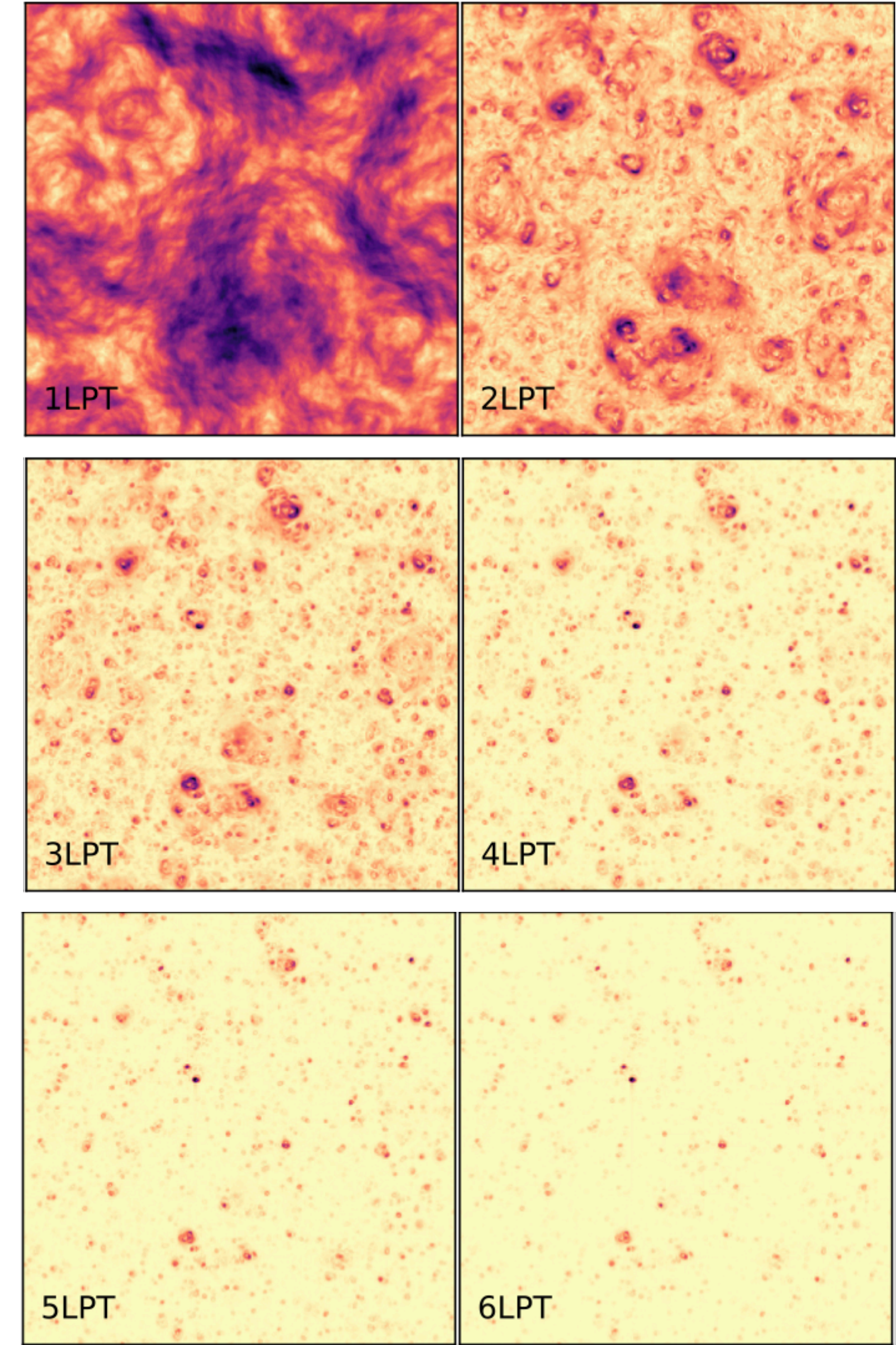
We want to solve this as a perturbative series (D is small parameter)

$$\Psi(\mathbf{q}, \tau) = \sum_{n=1}^{\infty} D(\tau)^n \Psi^{(n)}(\mathbf{q})$$

Buchert (1994), Catelan (1995), Bouchet+(1995), n=3
Rampf (2012), Zheligovsky&Frisch (2014), Matsubara (2015), all order

yields recursion relations to all orders.

For LCDM, actually $D^{(n)}(\tau) \neq D^n(\tau)$, see Rampf, Schobesberger & OH(2022), Fasiello, Fujita, Vlah (2022)



Rampf+OH (2021)

What is the range of applicability of such a theory (aka the 'radius of convergence')?

Perturbative forward model

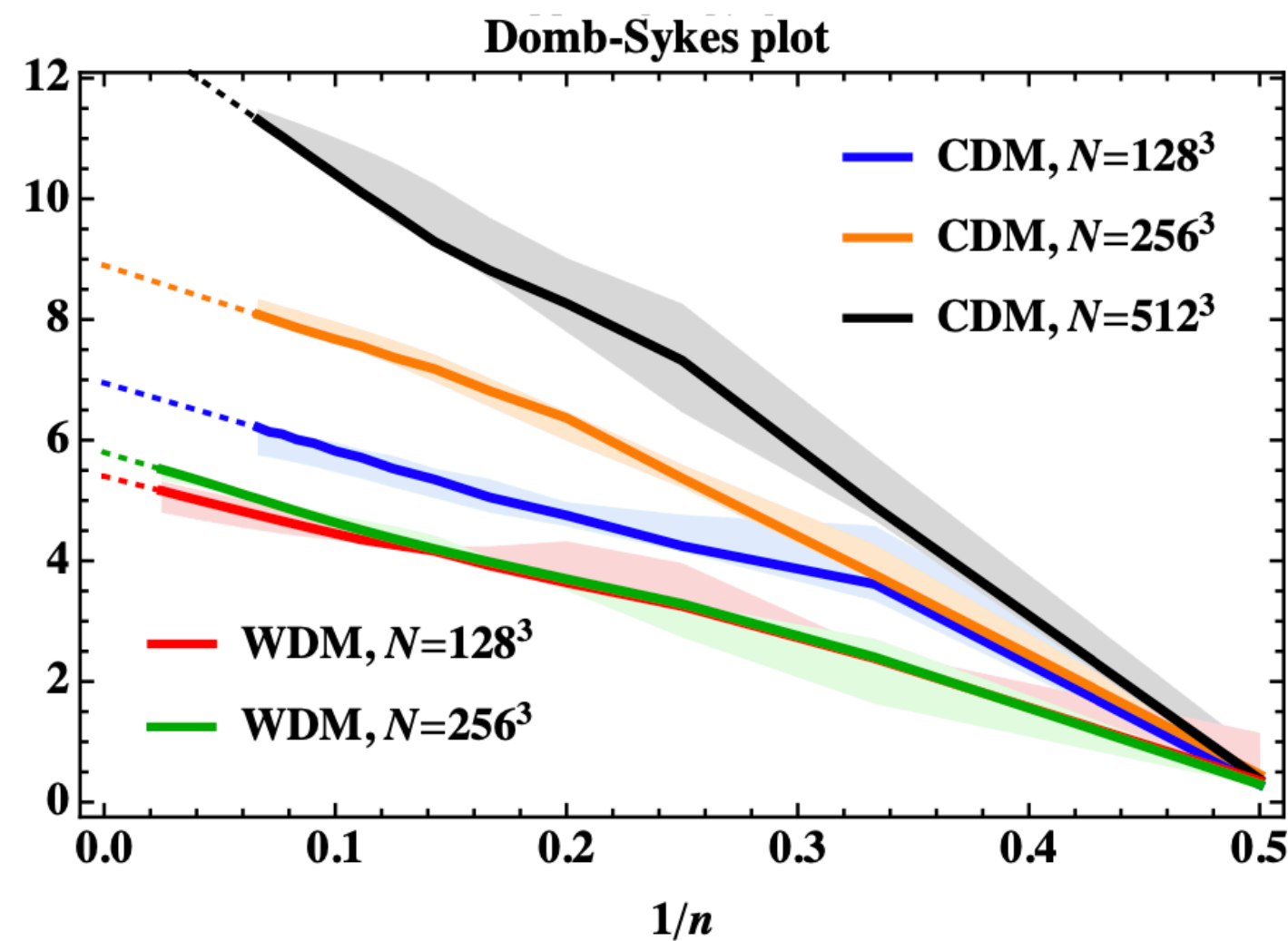
- Advantage: analytical insight
- Lagrangian PT: Extend the asymptotic case to a full Taylor-series

$$\Psi(\mathbf{q}, a) := \mathbf{X}(\mathbf{q}, a) - \mathbf{q} = \sum_{n=1}^{\infty} \psi^{(n)}(\mathbf{q}) a^n$$

truncate at order N

Buchert (1994), Catelan (1995), Bouchet+(1995), N=3
 Rampf (2012), Zeligovsky&Frisch (2014), Matsubara (2015), all order

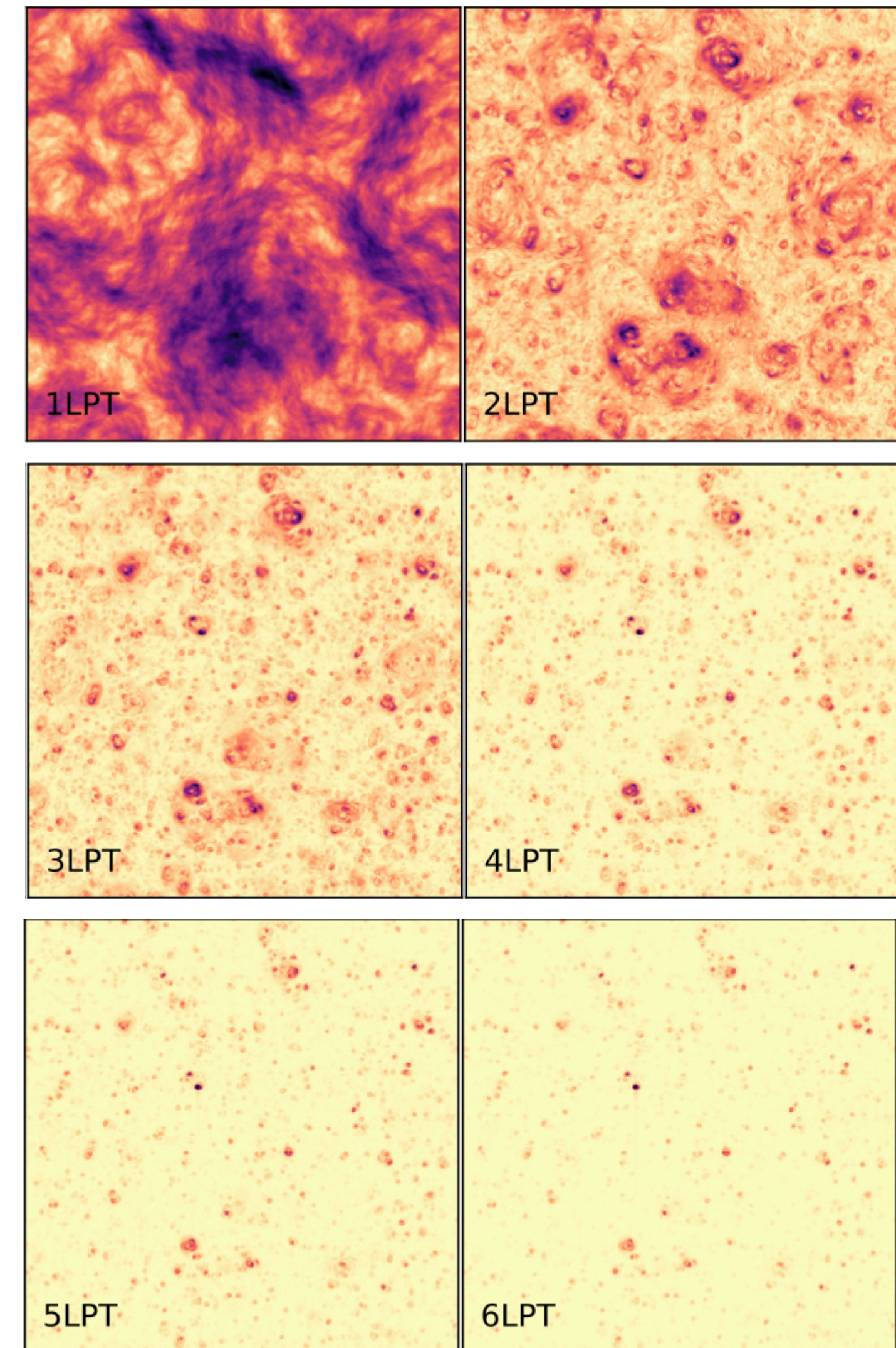
- Strong numerical evidence for convergence prior to shell-crossing:



$$\lim_{n \rightarrow \infty} \frac{\|\psi^{(n)}\|}{\|\psi^{(n-1)}\|} = \frac{1}{D_\star}$$

Numerically $D_\star \geq a_\star$

(Except for measure 0 set)



Rampf+OH (2021)

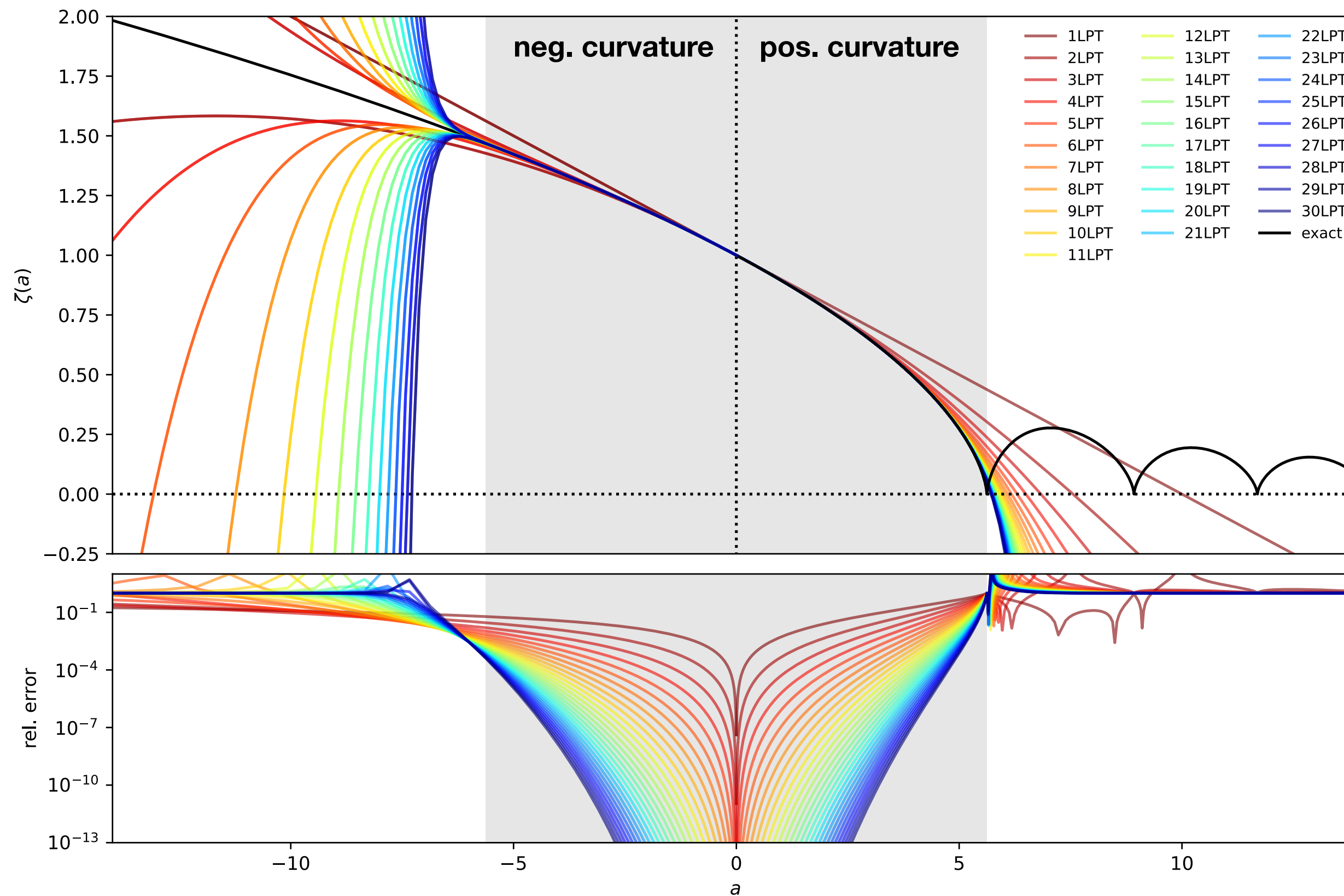
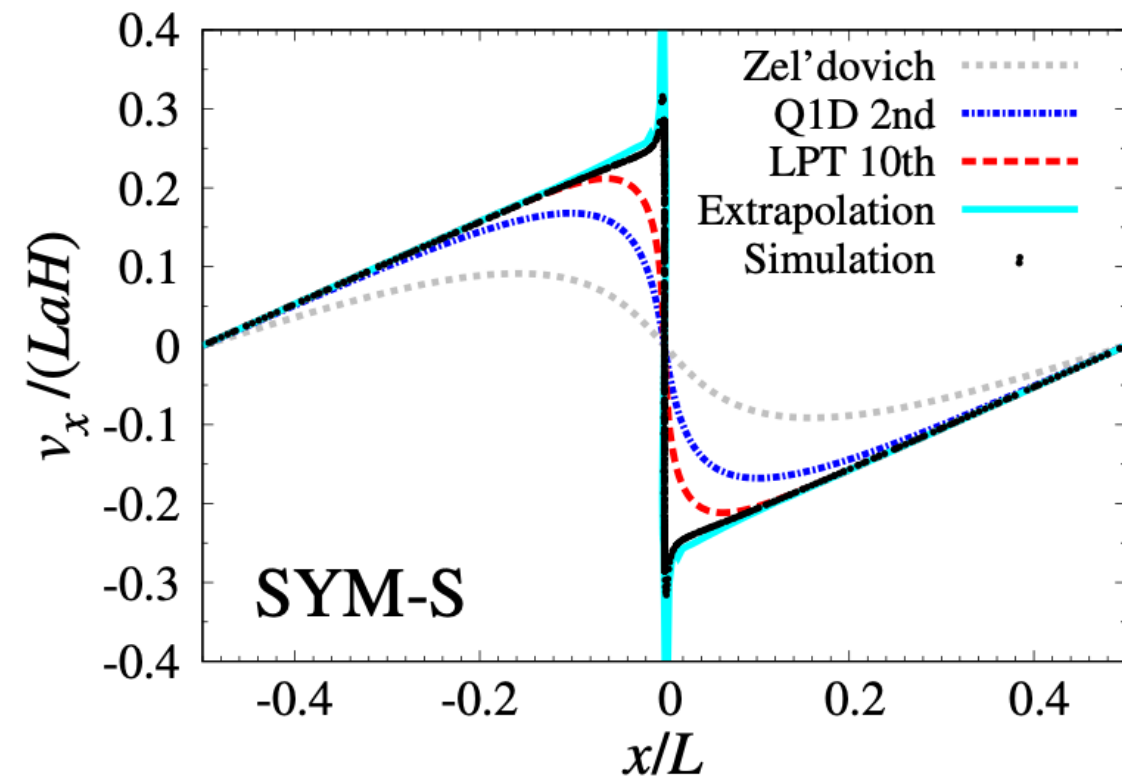
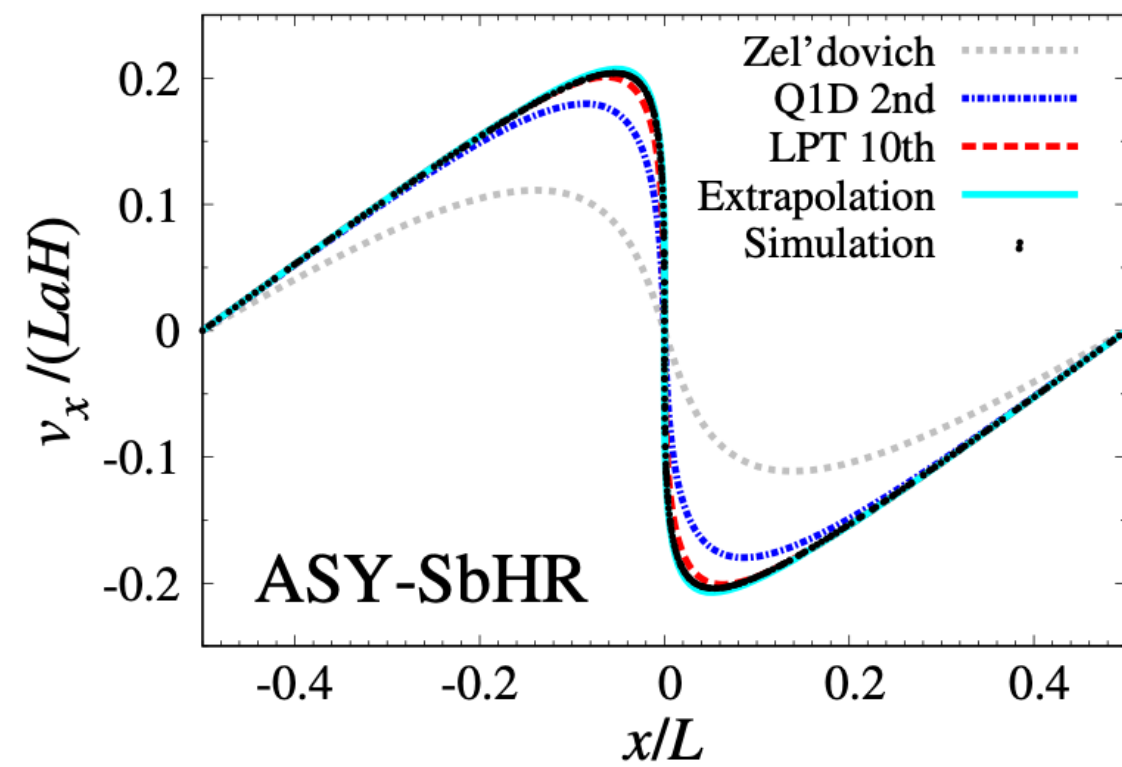
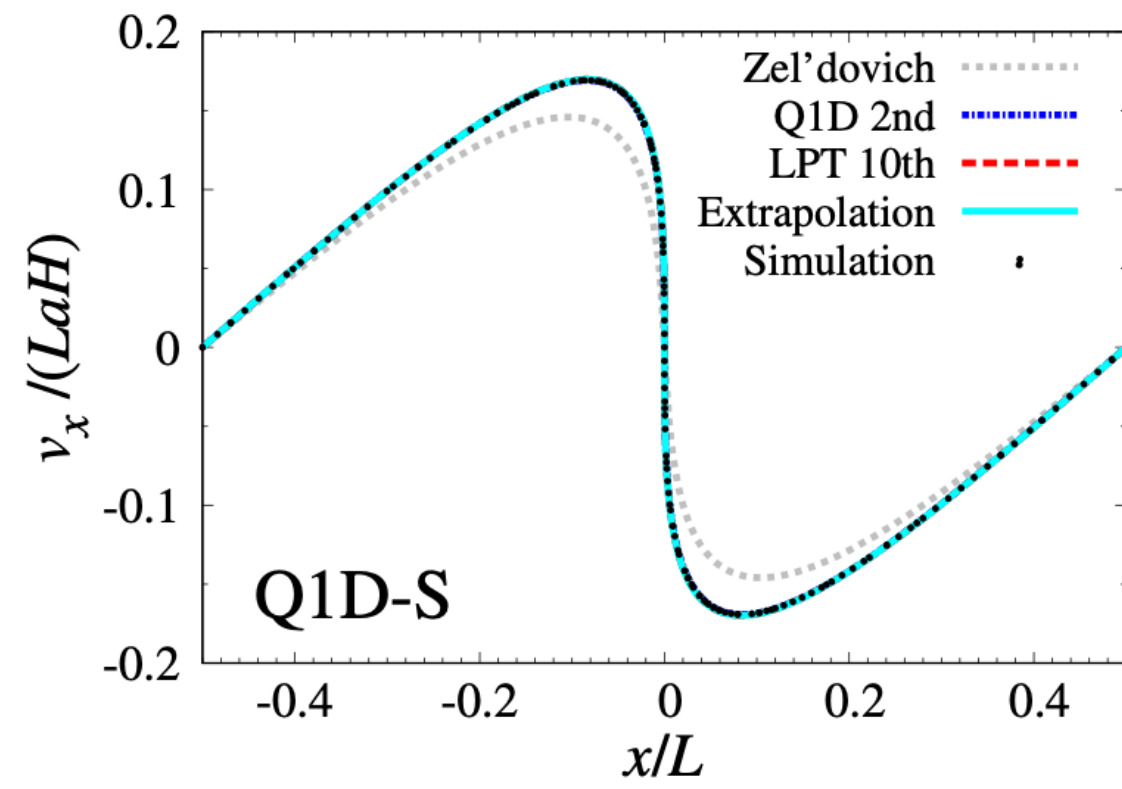
Convergence limiting singularities

Rampf, Frisch, OH '21: For 1D planar initial data, generic non-analyticity appears in the displacement

$$\Psi(\mathbf{q}, a) \propto (a - a_*(\mathbf{q}))^{\frac{2}{3}}$$

for 1D spherically symmetric initial data, this is hardened to (Rampf & OH '23)

$$\Psi(\mathbf{q}, a) \propto (a - a_*(\mathbf{q}))^{\frac{2}{3}}$$

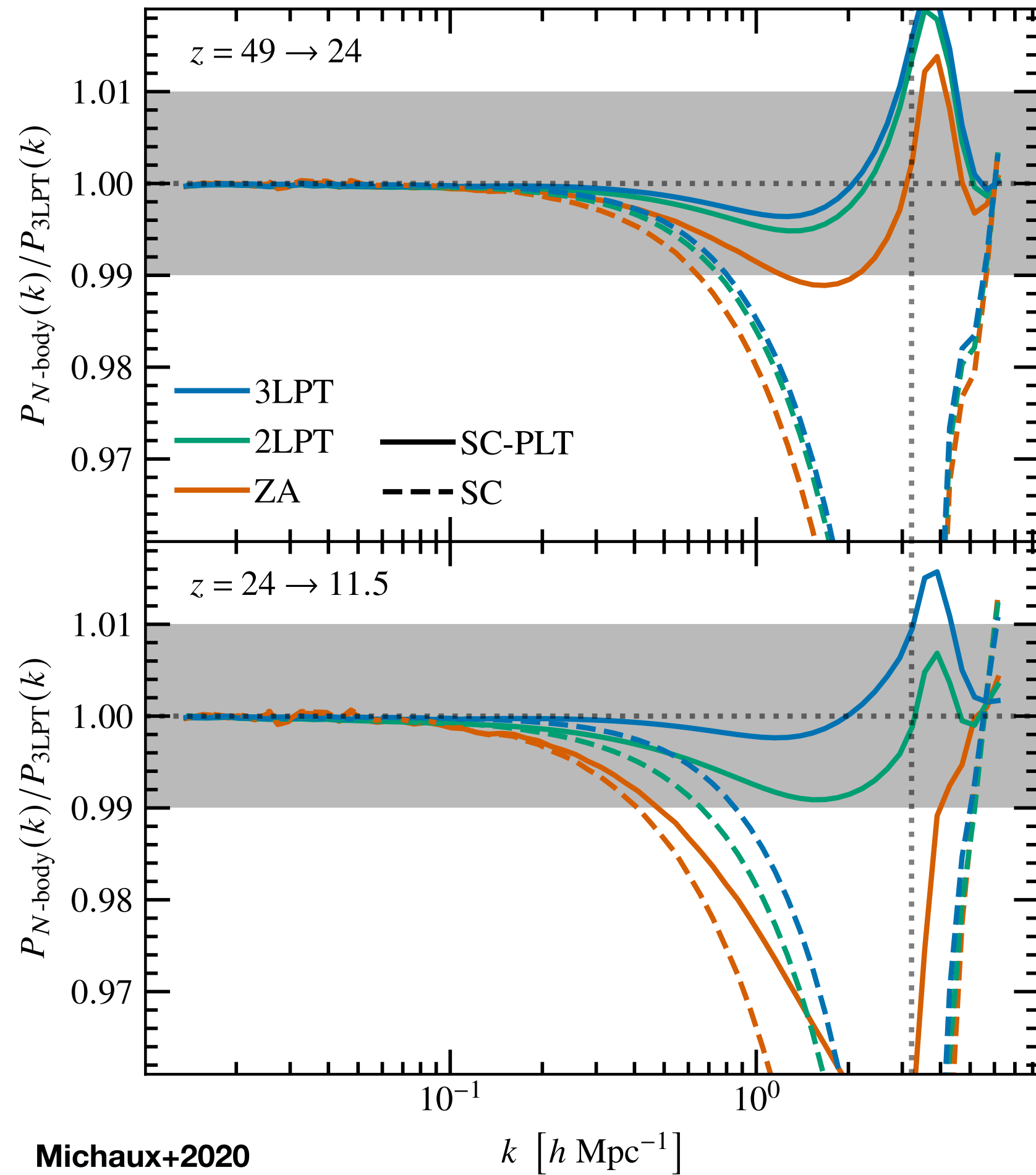


Using N-body to go further:

use finite order, discretised LPT to set up an N-body simulation, then run it

Agreement between N-body and LPT?

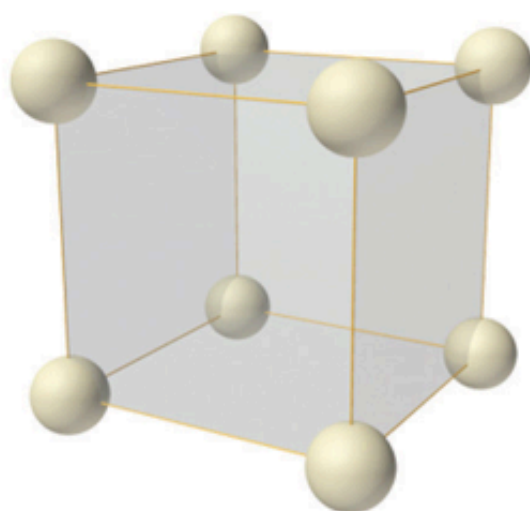
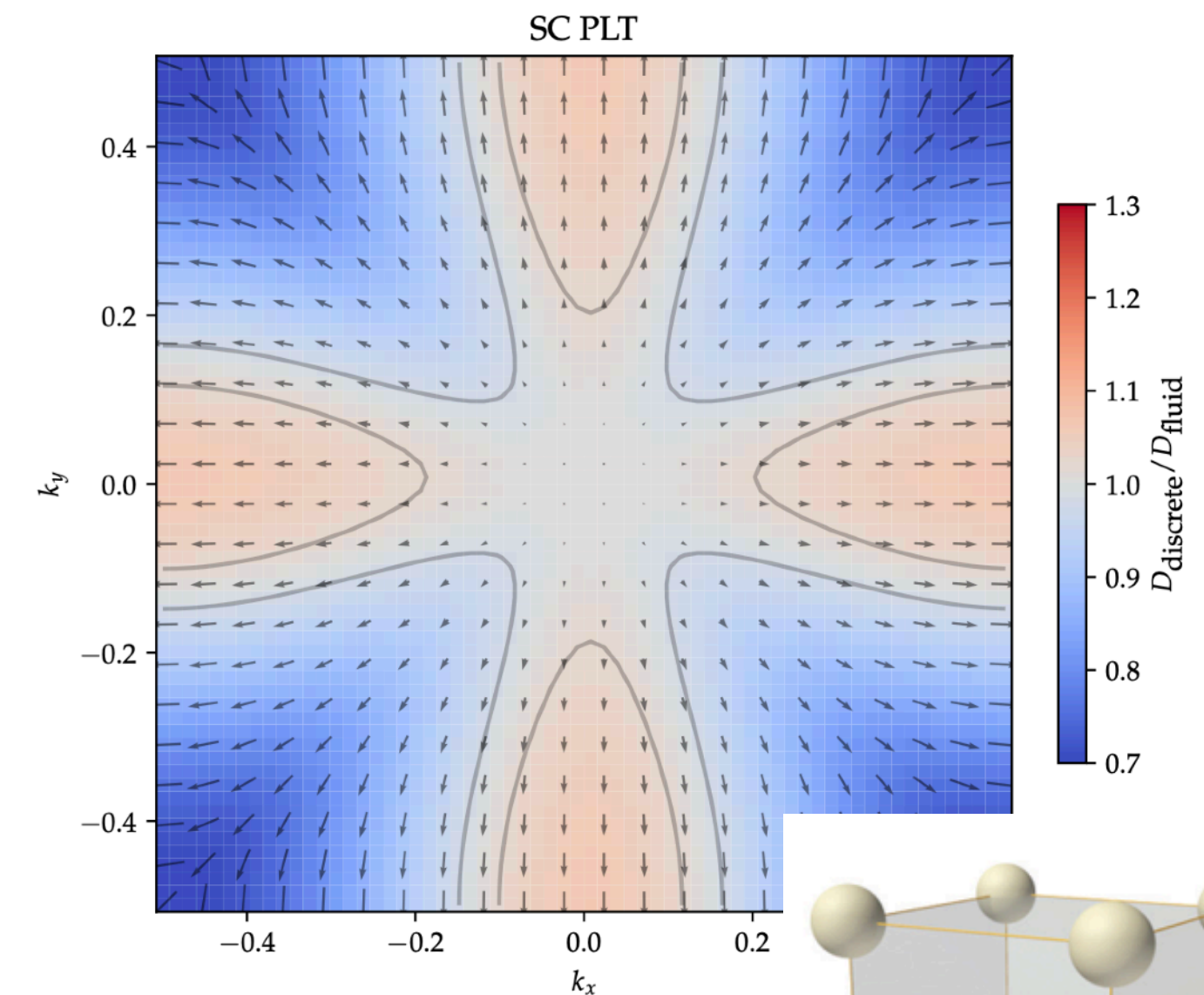
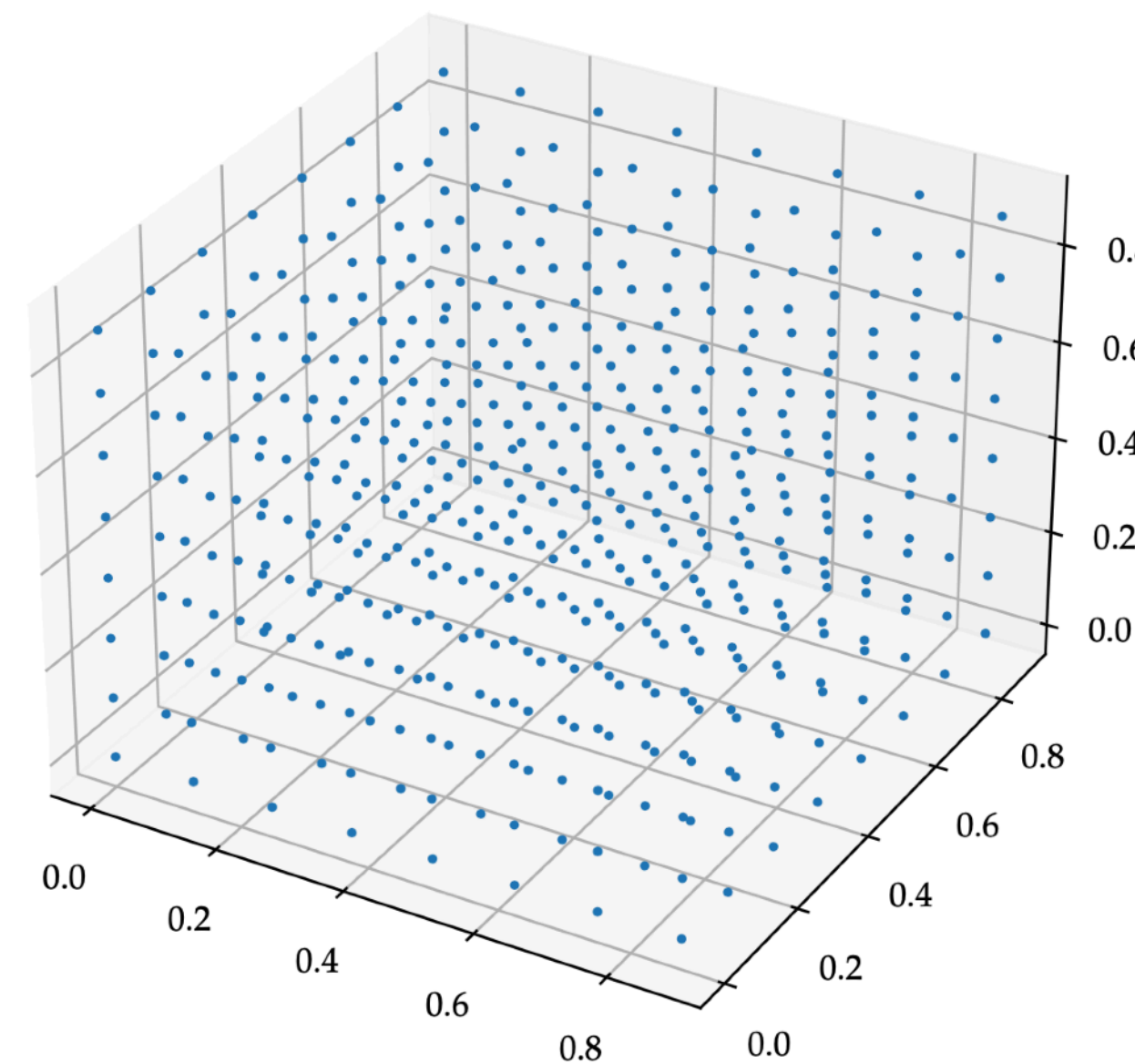
Comparison of weakly evolved power spectra:



How is this possible? They model the same discrete set of modes

Two sources of error:

- 1) the nLPT truncation error (Scoccimarro 1998, Crocce+2006) aka 'transients'
- 2) the N-body discreteness (and force) errors:



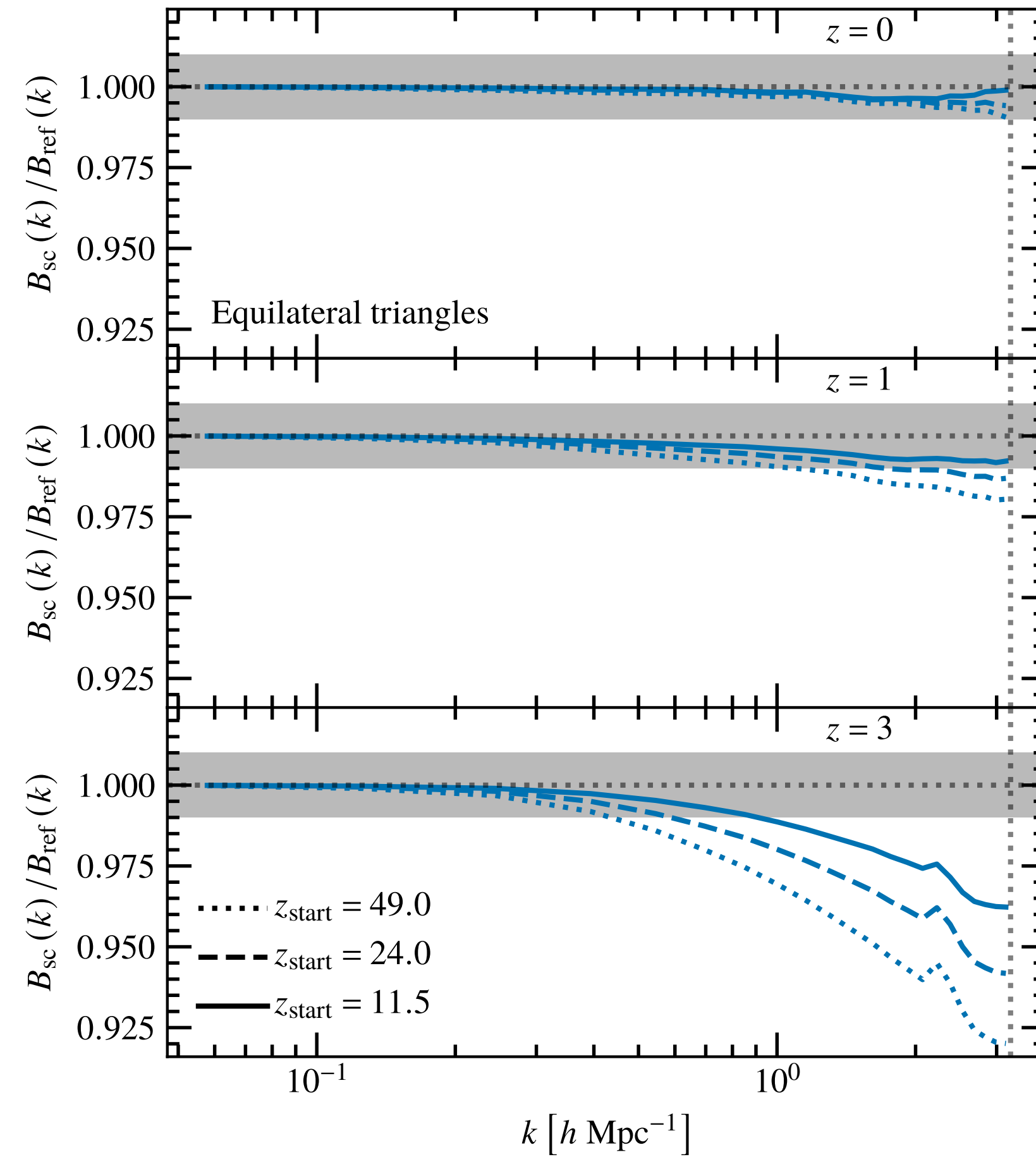
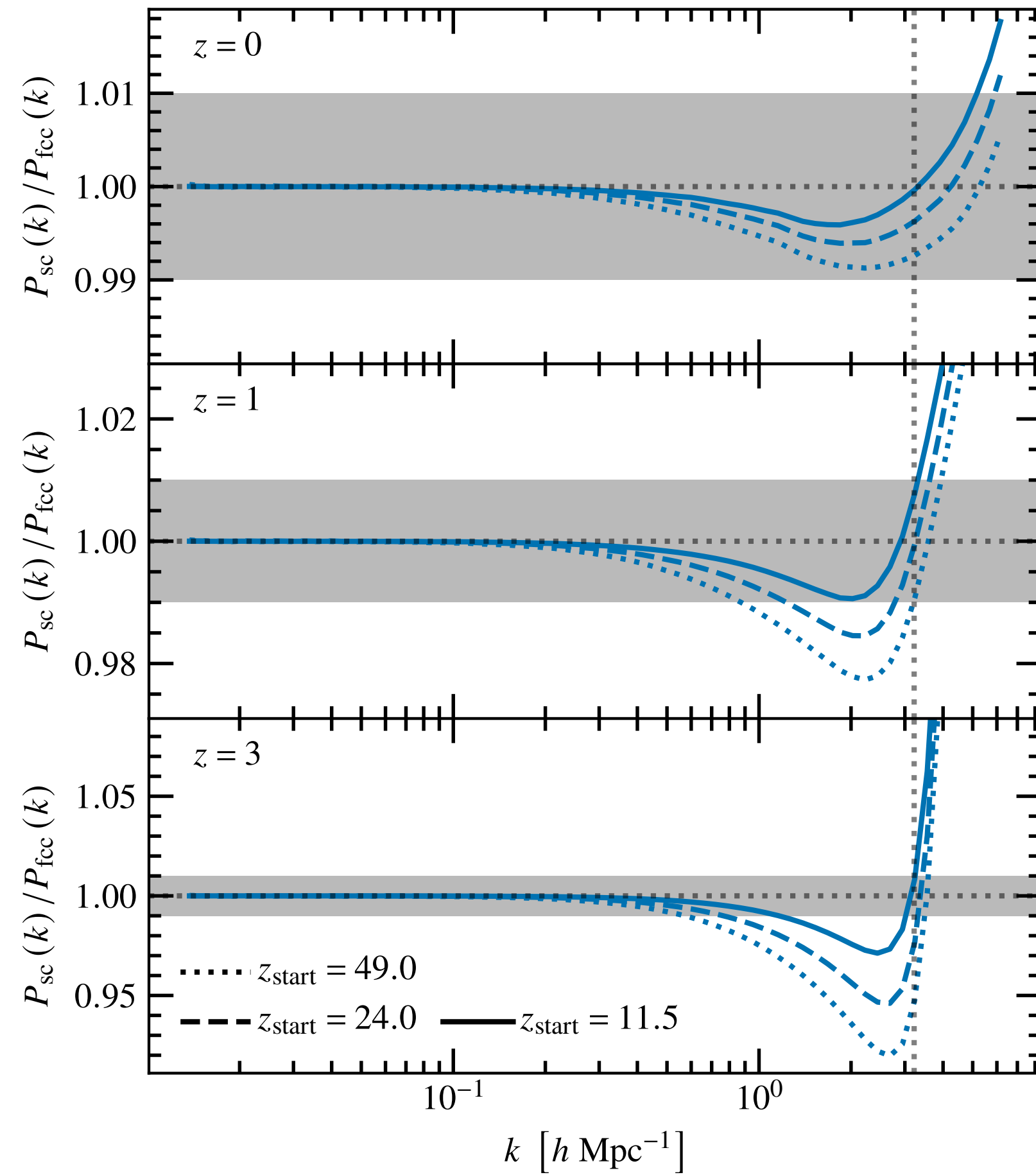
SC
(simple cubic)

cf. PLT of Joyce+2005, Joyce&Marcos 2007, Marcos 2008, but also Garrison+2016

Discreteness – impact on low- z spectra

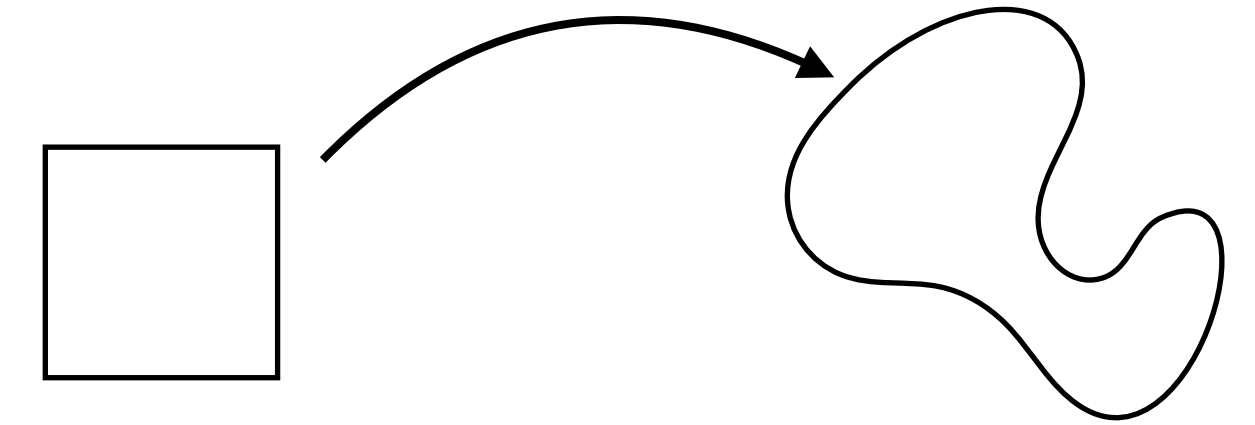
effect at low z wiped out by non-linearity (scale-mixing, asymptotic halo profiles), but not at higher z

Michaux+2020



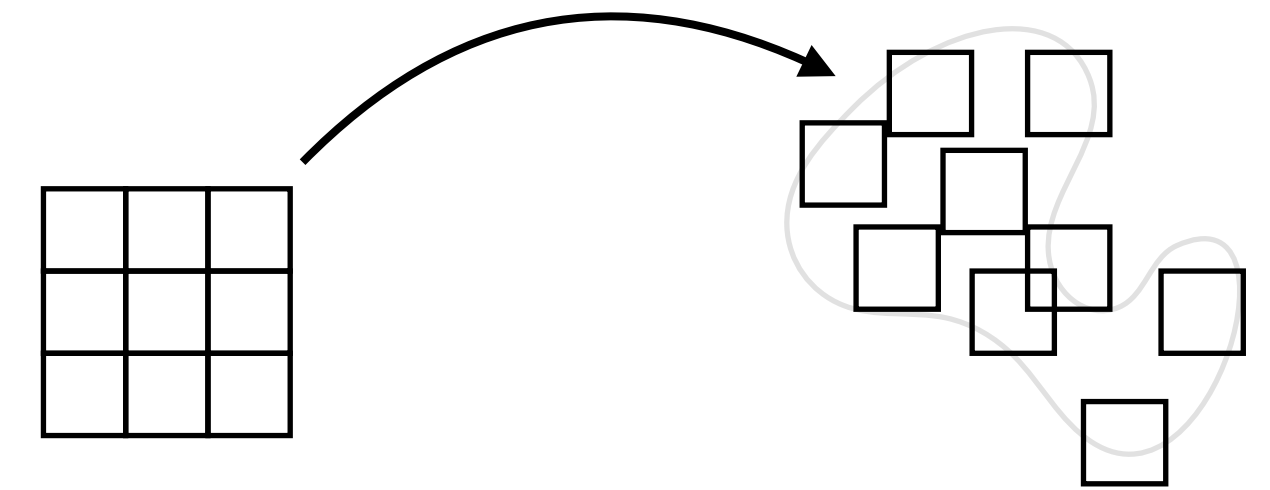
fluid-limit

$$Q \subset \mathbb{R}^3 \rightarrow \mathbb{R}^6 : \mathbf{q} \mapsto (\mathbf{x}_{\mathbf{q}}(t), \mathbf{v}_{\mathbf{q}}(t))$$



discrete case

$$i \in \{1 \dots N\} \mapsto (\mathbf{x}_i, \mathbf{v}_i)$$



should start late for smaller errors (LPT more correct than N-body while valid)

(cf. also Marcos 2008)

**So nLPT wins? use N-body only
to go where nLPT cannot?**

Time integrators, what do they do

Lump all phase space coordinates together $\xi_j := (X, P)_j$

Then Hamiltonian EoMs are a first order operator equation

$$\dot{\xi}_j = \hat{\mathcal{H}}(t) \xi_j \quad \text{with} \quad \hat{\mathcal{H}}(t) := \{\cdot, \mathcal{H}(t)\} = \{\cdot, \alpha T\} + \{\cdot, \beta V\} =: \hat{D}(t) + \hat{K}(t).$$

With formal solution

$$\xi_j(t) = \mathcal{T} \exp \left[\int_0^t dt' \hat{\mathcal{H}}(t') \right] \xi_j(0)$$

Now apply Strang operator splitting to find coefficients consistent with an expansion to order m

$$\mathcal{T} \exp \left[\int_t^{t+\epsilon} dt' \hat{\mathcal{H}}(t') \right] \simeq \exp [\epsilon_n \hat{K}] \cdots \exp [\epsilon_3 \hat{D}] \exp [\epsilon_2 \hat{K}] \exp [\epsilon_1 \hat{D}] + \mathcal{O}(\epsilon^m).$$

Finally expand operator exponentials into generators

$$\xi_j(\tilde{t} + \epsilon) = \left(I + \frac{\epsilon}{2} \hat{D} \right) (I + \epsilon \hat{K}) \left(I + \frac{\epsilon}{2} \hat{D} \right) \xi_j(\tilde{t})$$

Better time integrators for LSS studies

Definition 1 (Canonical DKD integrator).

$$\begin{aligned}X_i^{n+1/2} &= X_i^n + \alpha(\tau_n, \tau_{n+1})P_i^n, \\P_i^{n+1} &= P_i^n + \beta(\tau_n, \tau_{n+1})A\left(X_i^{n+1/2}\right), \\X_i^{n+1} &= X_i^{n+1/2} + \gamma(\tau_n, \tau_{n+1})P_i^{n+1},\end{aligned}$$

Given some conditions on the coefficients, the integrator is globally 2nd order, if A is sufficiently smooth

Consider also more restrictive form (List&OH 23):

Definition 2 (Π -integrator).

$$\begin{aligned}X_i^{n+1/2} &= X_i^n + \frac{\Delta D}{2}\Pi_i^n, \\ \Pi_i^{n+1} &= p(\Delta D, D_n)\Pi_i^n + q(\Delta D, D_n)A(X_i^{n+1/2}), \\ X_i^{n+1} &= X_i^{n+1/2} + \frac{\Delta D}{2}\Pi_i^{n+1},\end{aligned}$$

Do N-body integrators reproduce LPT?

Should be able to get Zeldovich in **one** time step

Proposition 4 (Characterisation of Zel'dovich consistency). *A Π -integrator is Zel'dovich consistent if and only if $p(\Delta D, D_n)$ and $q(\Delta D, D_n)$ satisfy the following relation:*

$$\frac{1 - p(\Delta D, D_n)}{q(\Delta D, D_n)} = \frac{3}{2} \Omega_m D_{n+1/2} \stackrel{\text{EdS}}{\approx} \frac{3}{2} a_{n+1/2}.$$

Example (FastPM). FastPM method of Feng et al. (2016), but this is the DKD version, while they used KDK

$$\begin{aligned} \mathbf{X}_i^{n+1/2} &= \mathbf{X}_i^n + \frac{D_{n+1/2} - D_n}{F(D_n)} \mathbf{P}_i^n && \stackrel{\text{EdS}}{\approx} \mathbf{X}_i^n + \frac{\Delta a}{2a_n^{3/2}}, \\ \mathbf{P}_i^{n+1} &= \mathbf{P}_i^n + \frac{F(D_{n+1}) - F(D_n)}{G(D_{n+1/2})} \mathbf{A}(\mathbf{X}^{n+1/2}) && \stackrel{\text{EdS}}{\approx} \mathbf{P}_i^n + \frac{2}{3} \frac{a_{n+1}^{3/2} - a_n^{3/2}}{a_{n+1/2}} \mathbf{A}(\mathbf{X}^{n+1/2}), \\ \mathbf{X}_i^{n+1} &= \mathbf{X}_i^{n+1/2} + \frac{D_{n+1} - D_{n+1/2}}{F(D_{n+1})} \mathbf{P}_i^{n+1} && \stackrel{\text{EdS}}{\approx} \mathbf{X}_i^{n+1/2} + \frac{\Delta a}{2a_{n+1}^{3/2}}, \end{aligned}$$

FASTPM is the unique Π -integrator that is both Zel'dovich consistent and symplectic.

Proof of symplecticity, consistency, and convergence order see List&OH 23 (not done in Feng+12)

Counterexample. All integrators typically used for N-body simulations are not Zel'dovich consistent.

Can we do even better?

Spoiler: yes, but have to give up symplecticity

Second order LPT (2LPT) can be written

$$X_i(D) = X_i^n + [D - D_n]\psi_i^{n,(1)} + [D - D_n]^2\psi_i^{n,(2)}$$

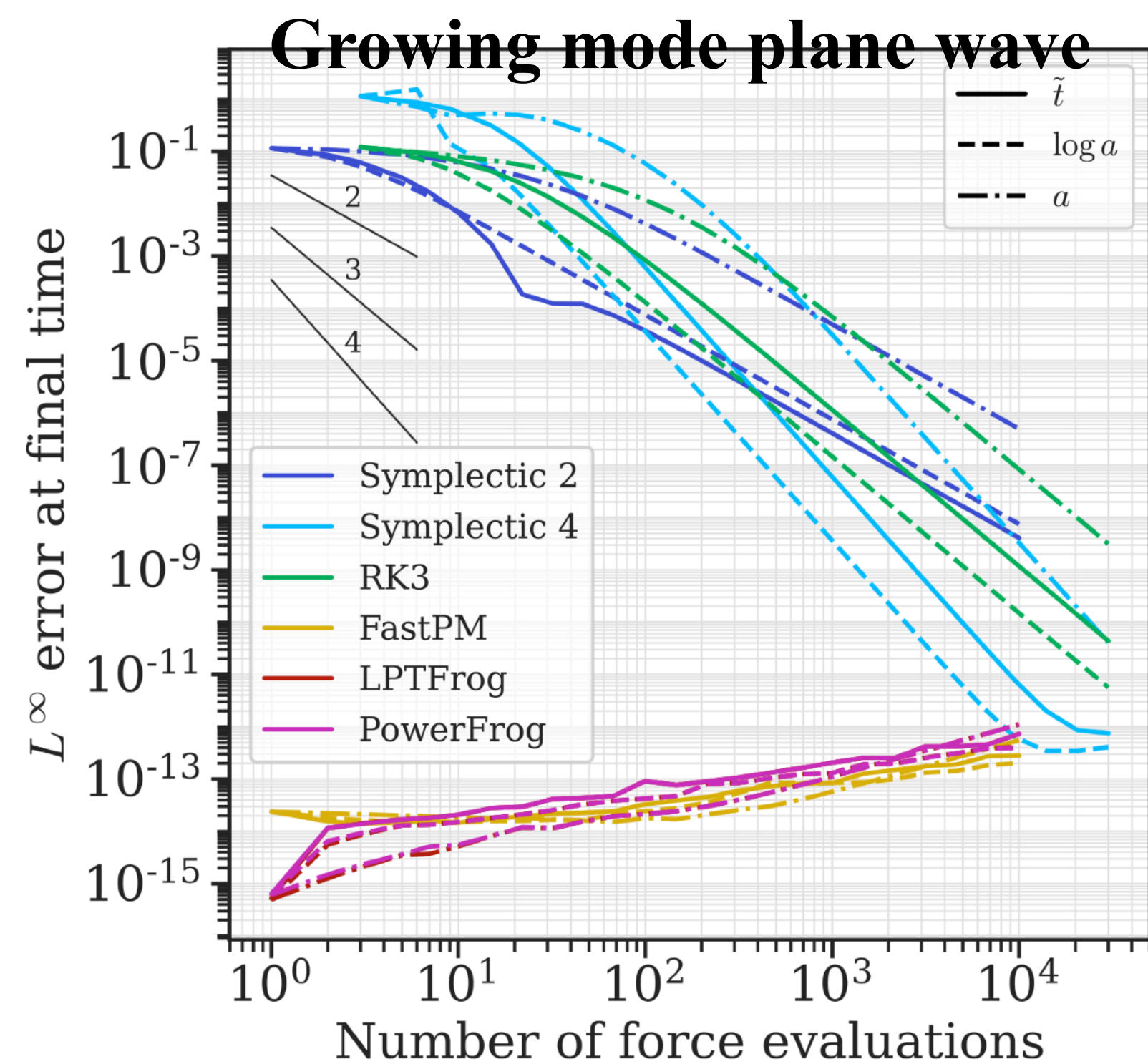
$$d_D X_i(D) = \psi_i^{n,(1)} + 2[D - D_n]\psi_i^{n,(2)},$$

$$d_D^2 X_i(D) = 2\psi_i^{n,(2)} = \text{const.}$$

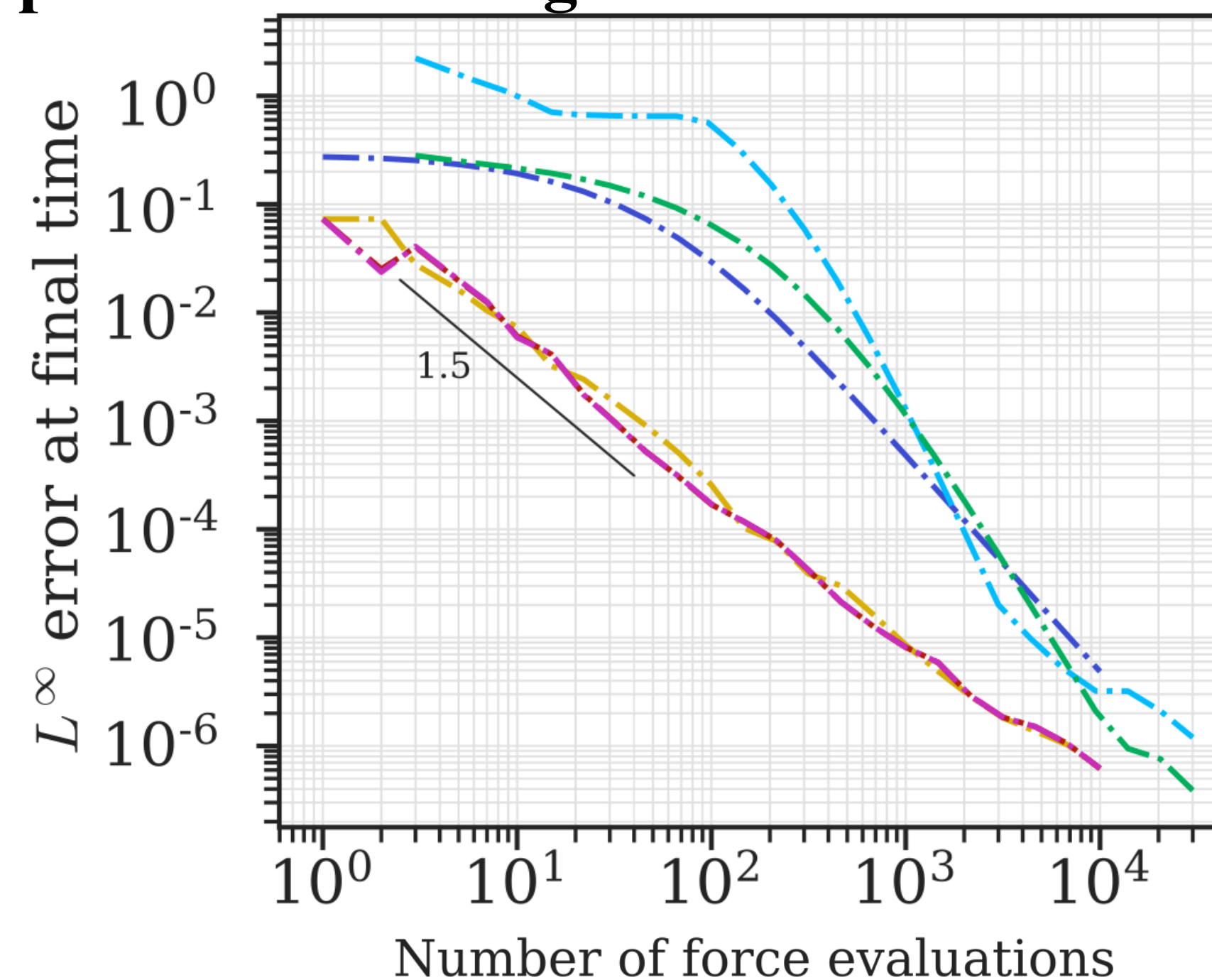
i.e. acceleration is constant

This can be matched in various ways to yield new integrators that should reproduce 2LPT, not only Zeldovich.

Tests for 1D collapse

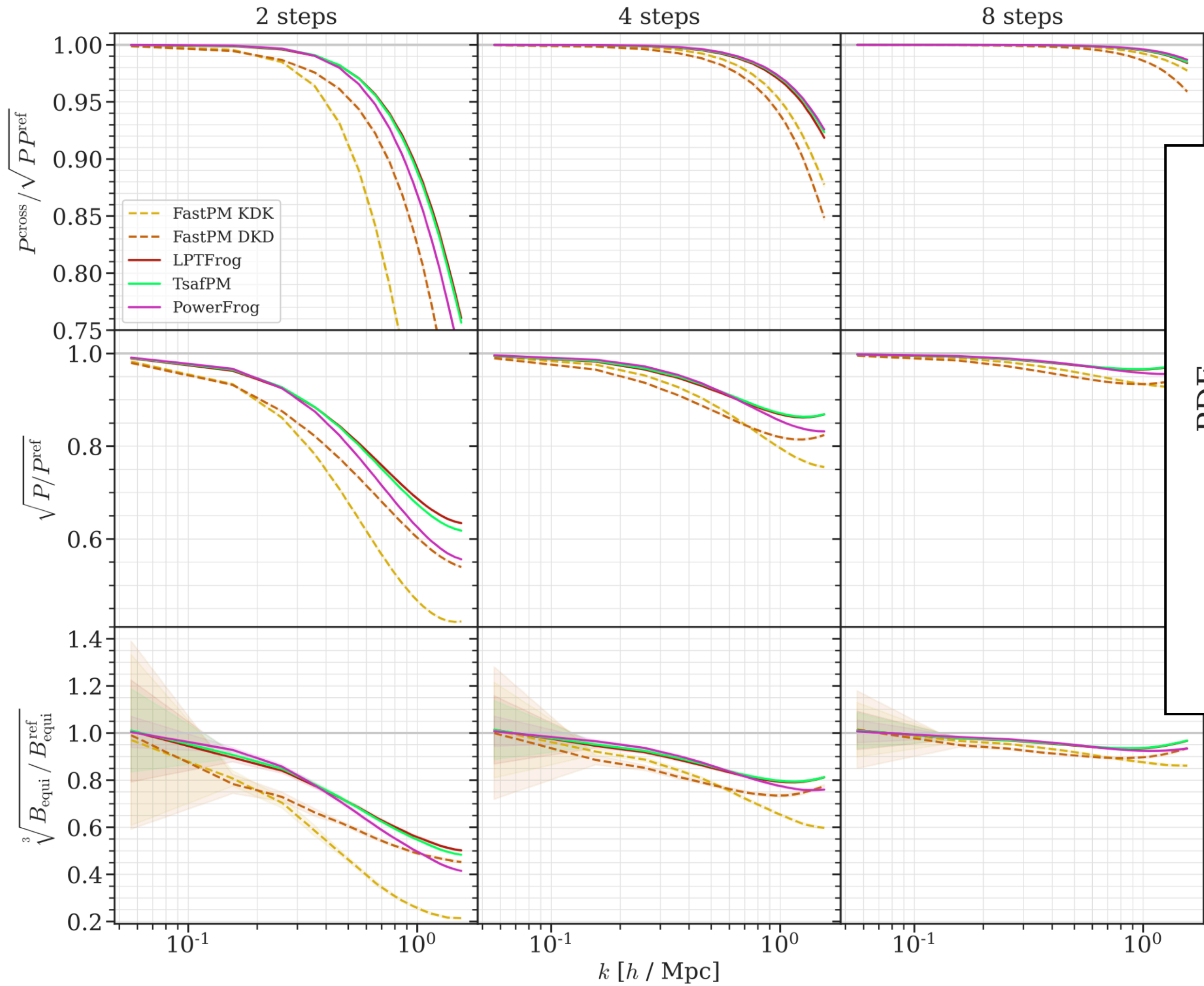


post shell-crossing: a

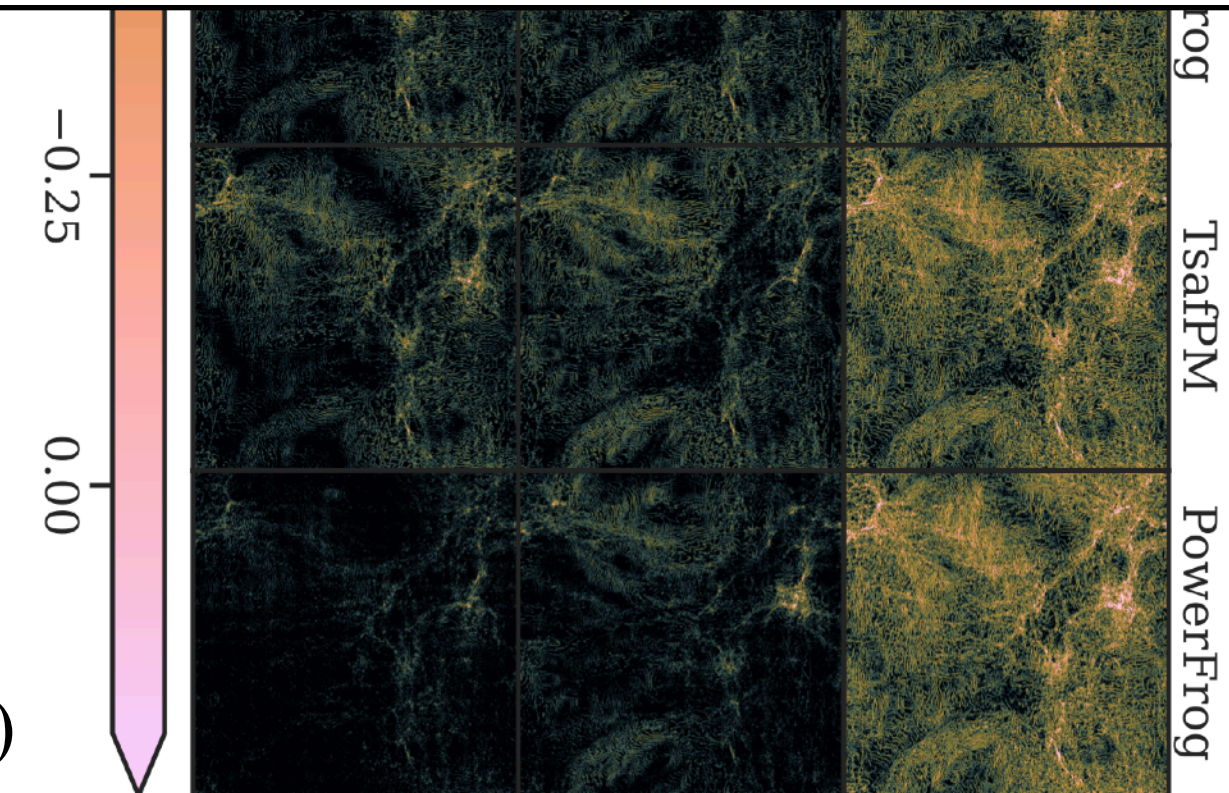
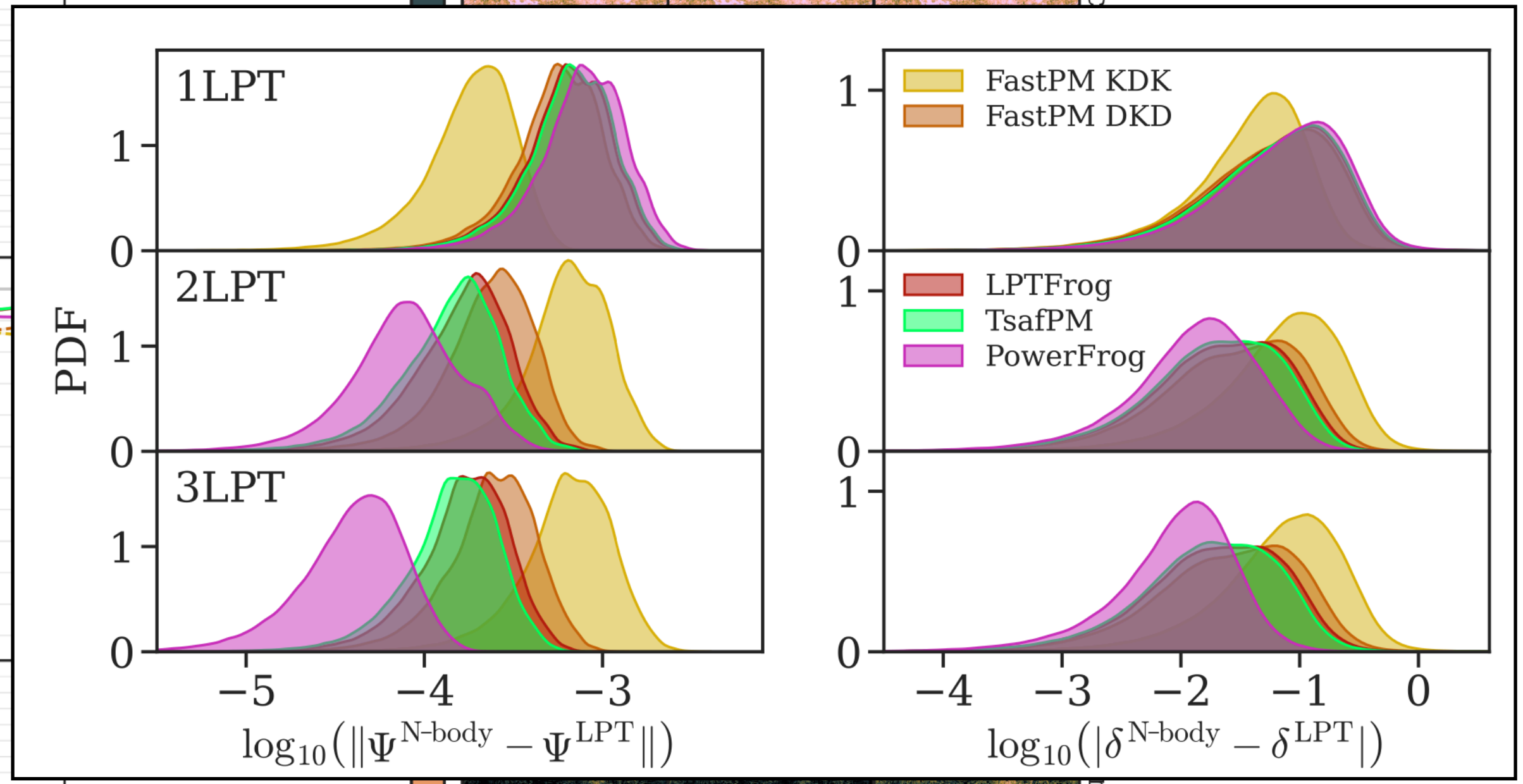
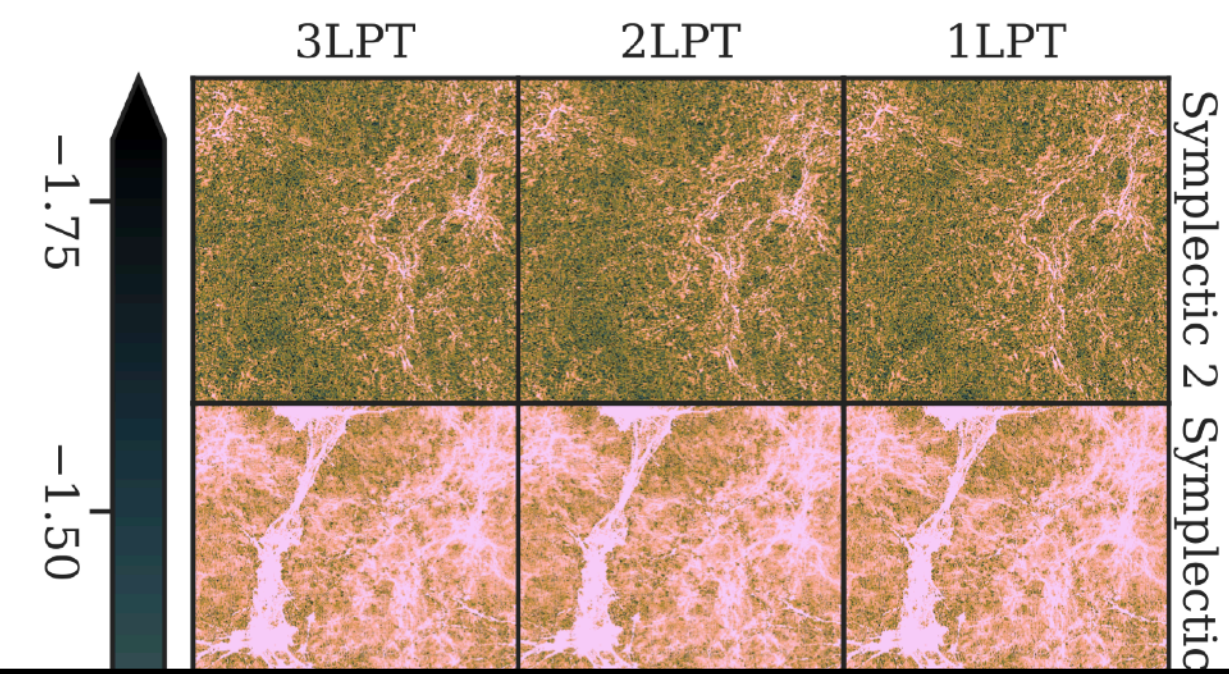


Tests in 2D and 3D

Quijote simulation at $z=0$



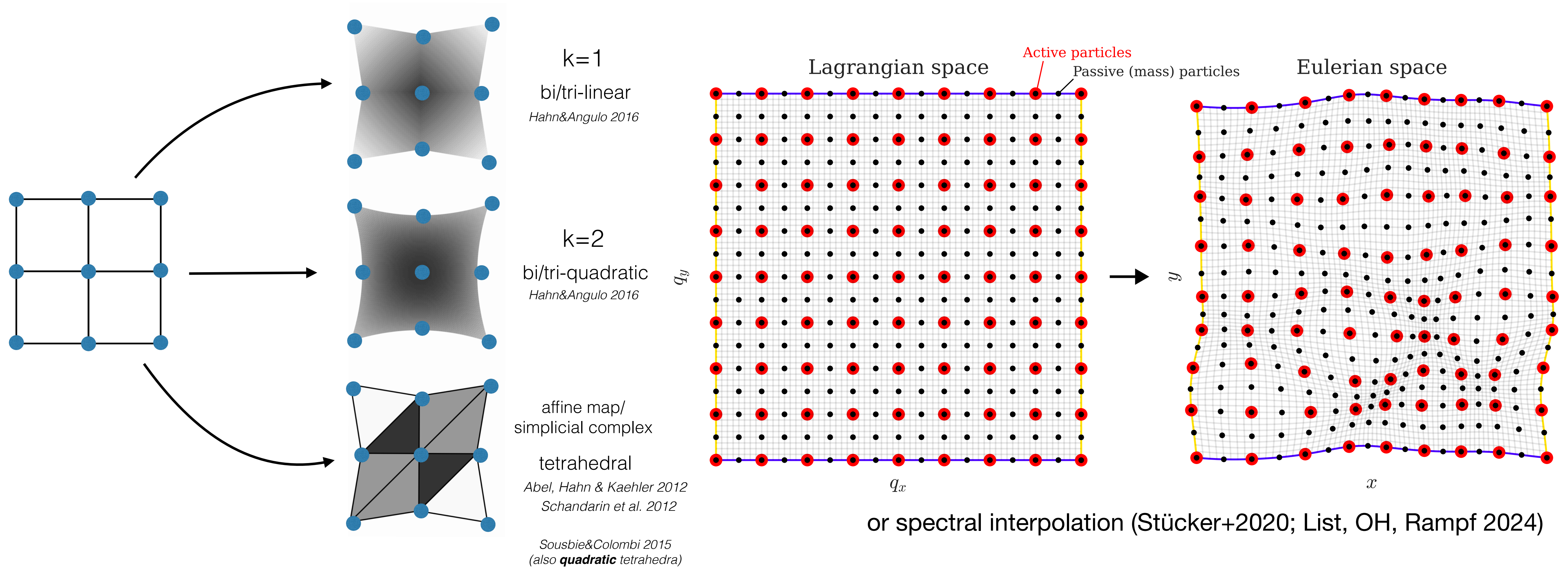
Comparison to nLPT



What do we need nLPT for now?

**integrator is 2.5LPT,
but can we beat discreteness?**

Beyond N-body - tessellation/resampling methods

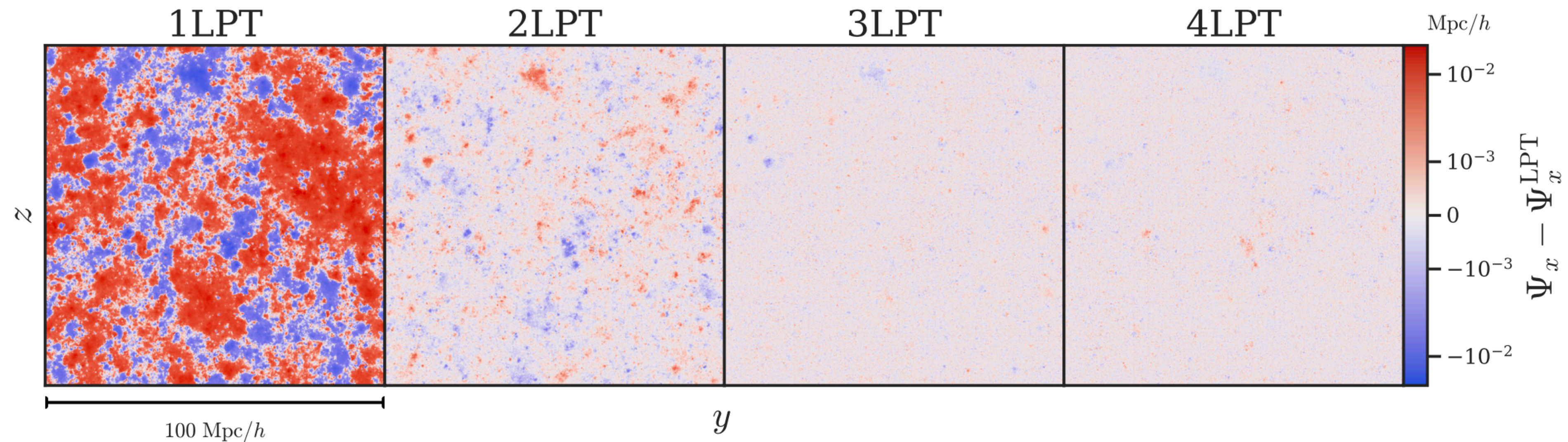


Unifying LPT and N-body: ICs without “LPT”

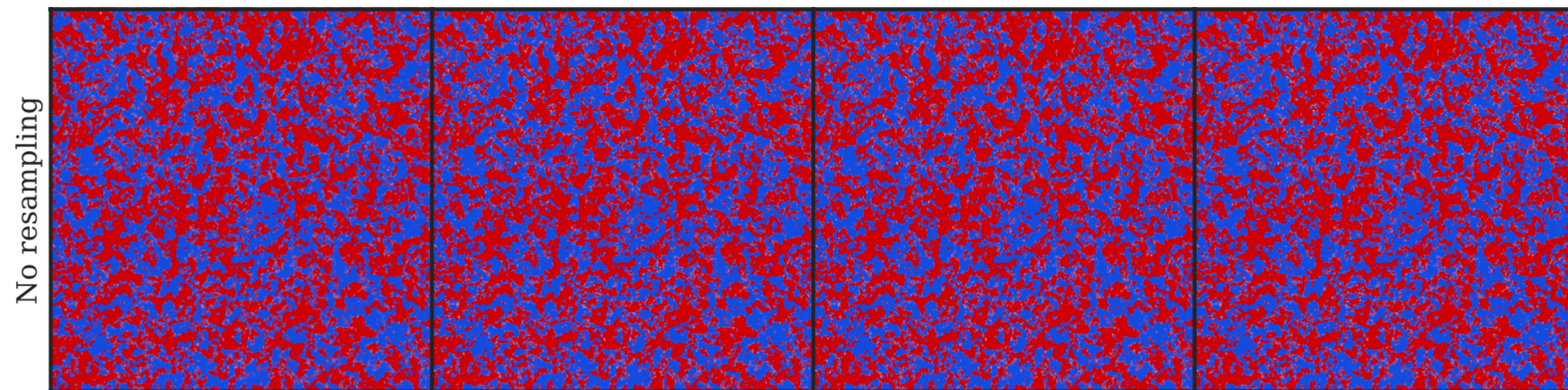
List, OH & Rampf, PRD 2024

PowerFrog integrator is asymptotically consistent with 2LPT for $a \rightarrow 0$, **can start at $a=0$** as we do in LPT

Residual of **single PowerFrog step** from $a=\infty$ to $a=0.05$, wrt. nLPT:

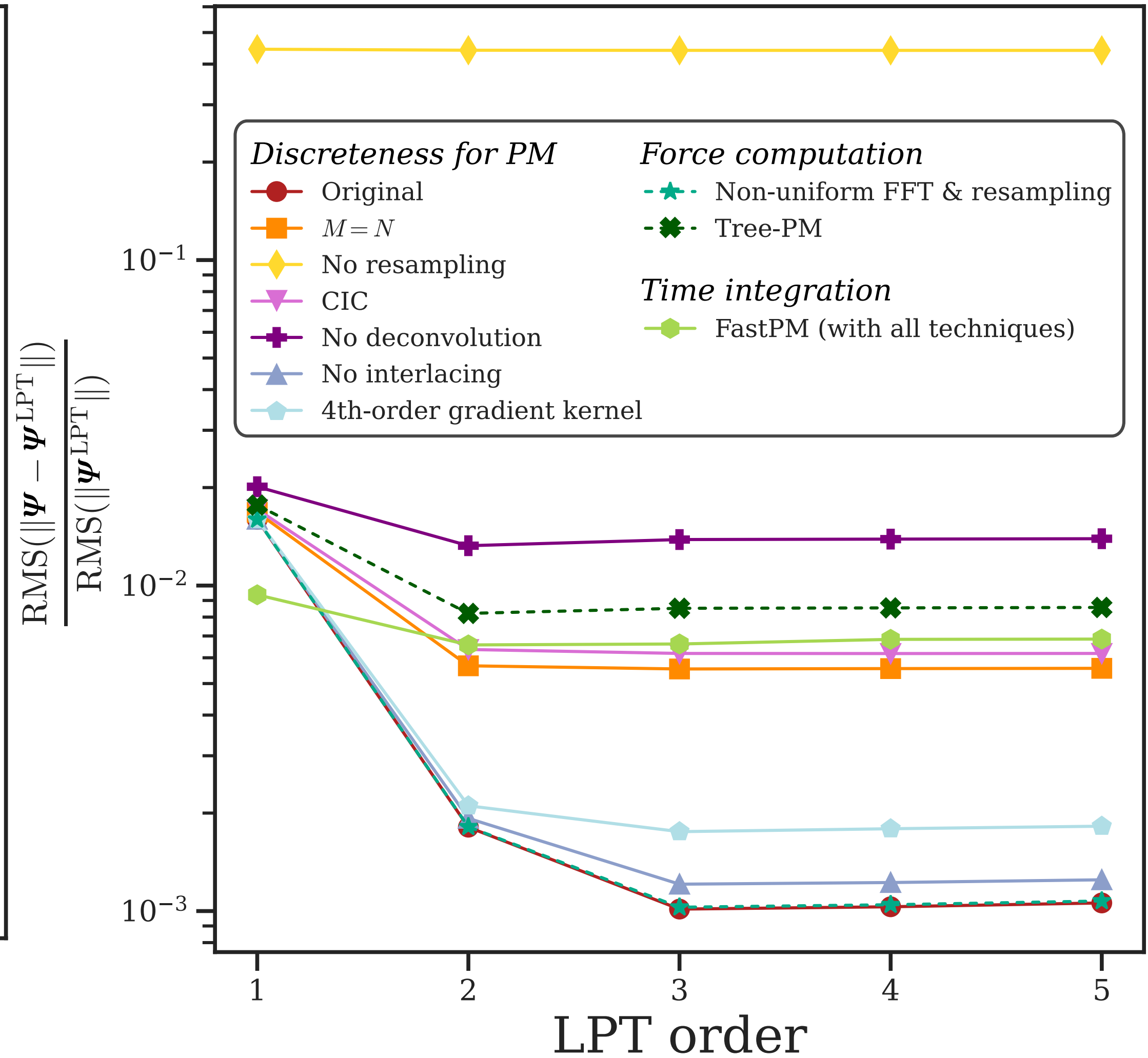
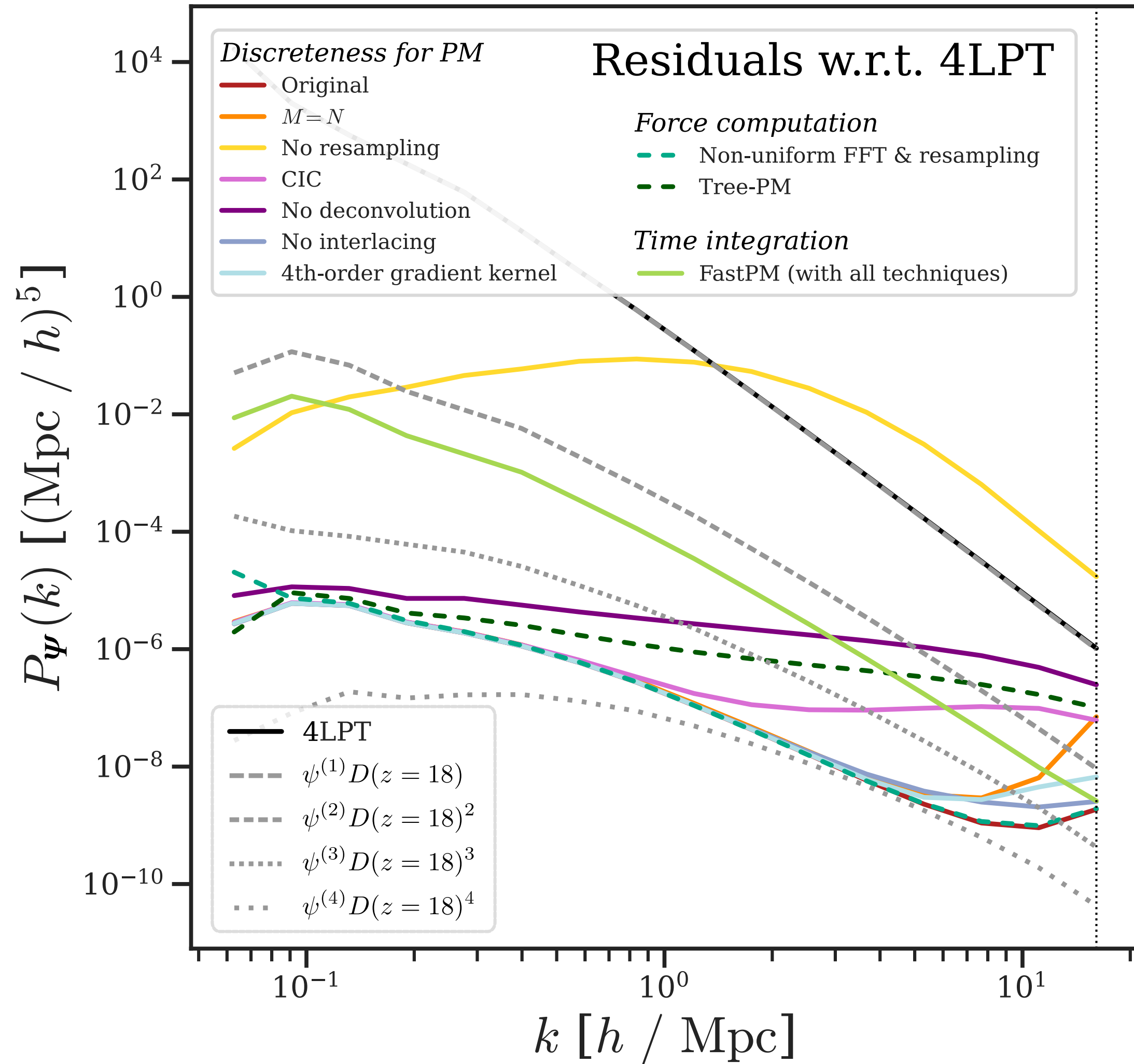


This is only possible after controlling **all** discreteness effects and approx. errors in the N-body simulation, otherwise:



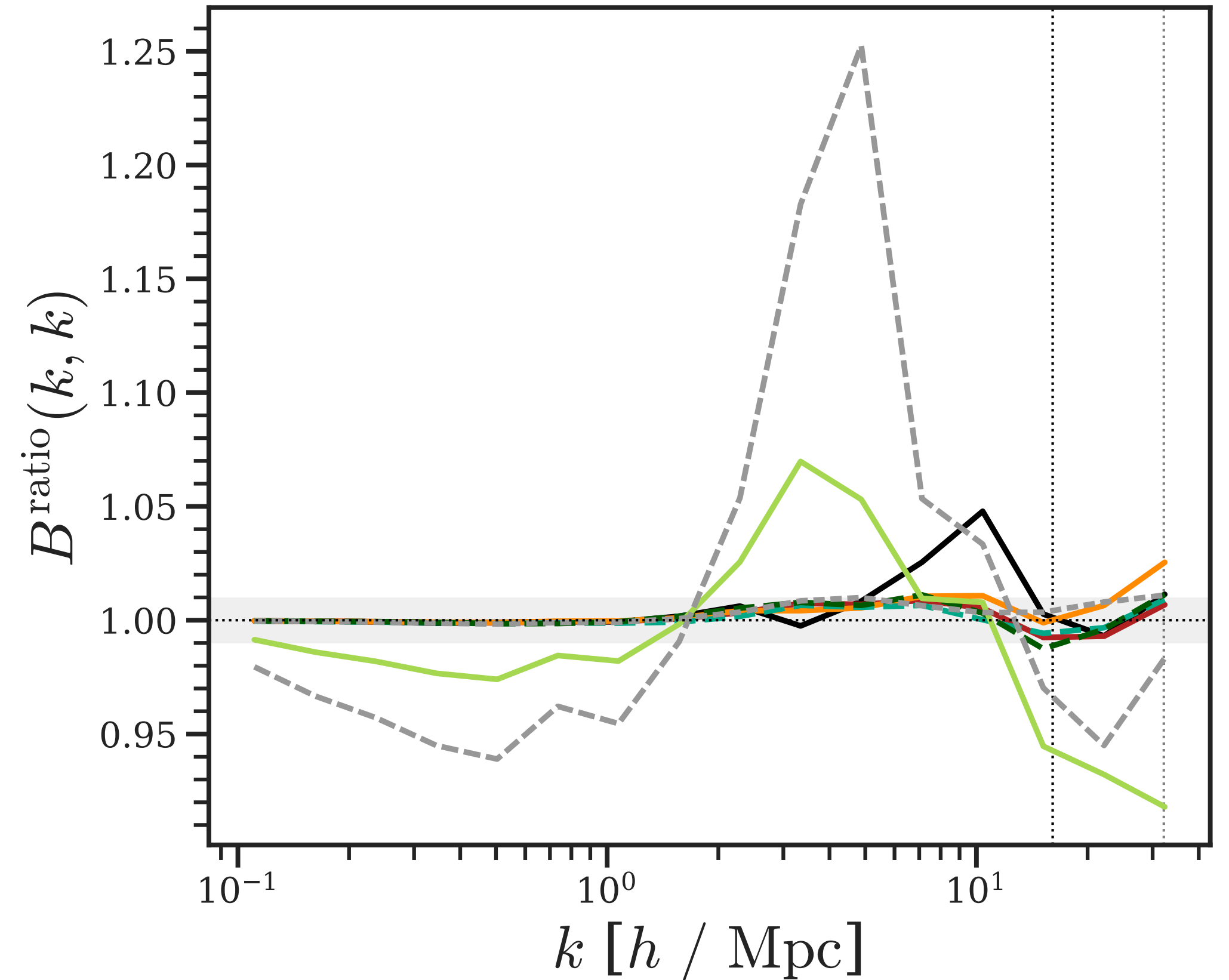
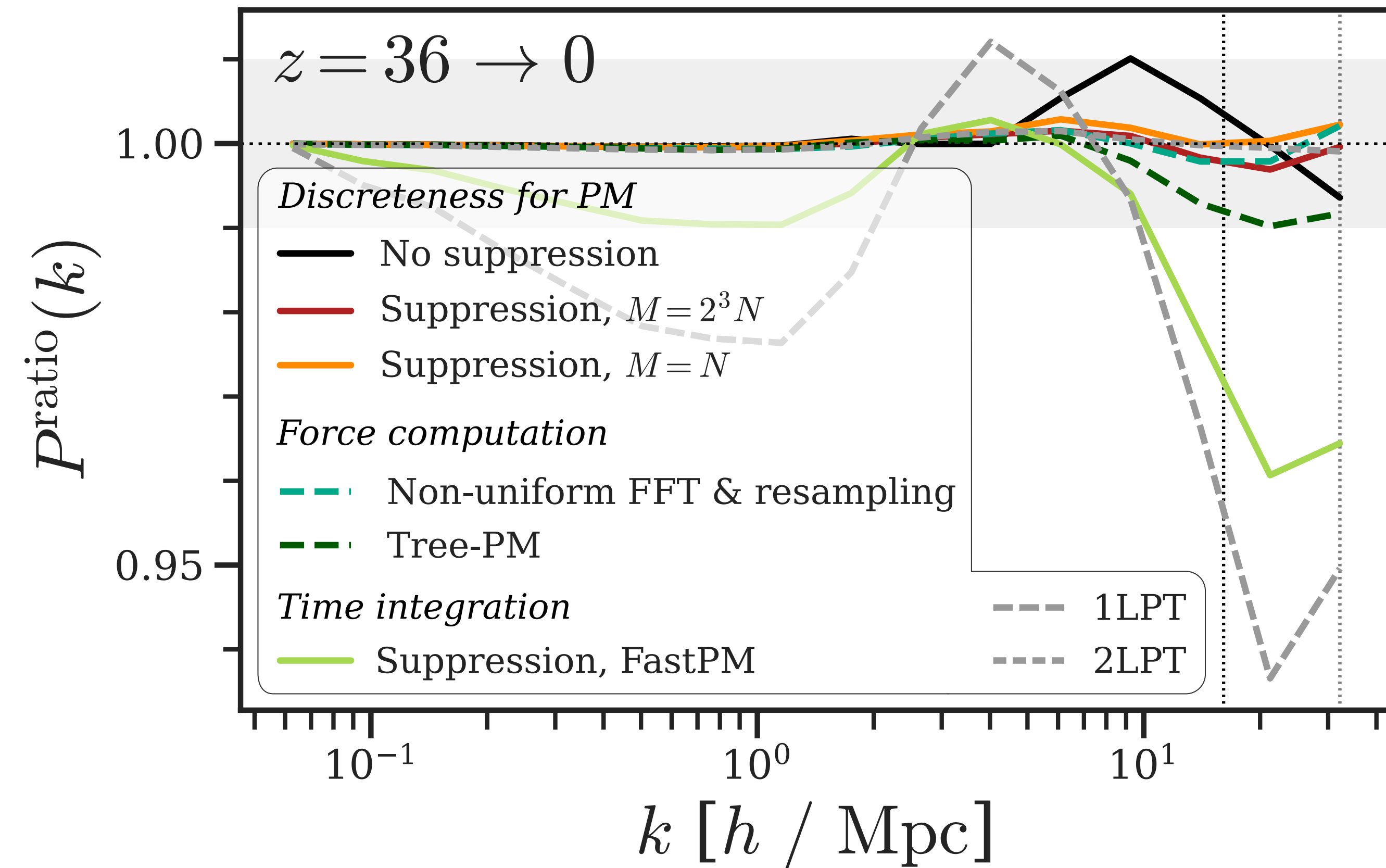
Unifying LPT and N-body: ICs without “LPT”

Errors at “IC” time ($a=0.05$)



Unifying LPT and N-body: ICs without “LPT”

Errors at final time (powerfrog step to $z=36$, then standard Gadget-4 N-body run)



SUMMARY

- LPT has key role in ICs for cosmological simulations, or as a field-level forward model
- Demonstrated convergence of LPT beyond shell-crossing, slow convergence due to SC singularities
- 3LPT needed for precision era N-body simulations, push to late starts to reduce errors
- new LPT inspired integrators (beyond 'FastPM') for fast simulations
- can now replace LPT initial conditions with N-body

SUMMARY

- LPT has key role in ICs for cosmological simulations, or as a field-level forward model
- Demonstrated convergence of LPT before shell-crossing, slow convergence due to SC singularities
- 3LPT needed for highest precision N-body simulations, push to late starts to reduce errors
- many new LPT inspired integrators (beyond 'FastPM') possible for fast simulations
- now possible to replace nLPT initial conditions with N-body, particularly interesting for inference

- n-LPT recursion
- all fast time integration methods ('ultimate' Rampf, List, OH, in prep.)
- discreteness suppression/sheet interpolation, NUFFT,...
- GPU based (run 512^3 in seconds)
- forward and backward differentiable
- has Einstein-Boltzmann module included (OH, List & Porqueres 2024)
- easy interface to inference
- **stay tuned...**

