

Boundary conditions and proper-time contours for worldline path integrals

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Workshop on

First Quantisation for Physics in Strong Fields

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- Simple examples for illustration and the tough ones in 'future directions'!

- Solutions to 'wave equations'

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$$\Psi(x) = \int_{\text{Initial Condition}}^{X(\text{final})=x} d[X] d[\text{lag. mult.}] d[\text{ghost}] e^{i(\text{Action})}$$

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- Here¹

- ① X = dynamical variable
 - ② Lagrangian multipliers $\Rightarrow$ constraints + gauge-fixing
 - ③ Ghost $\Rightarrow$ invariant measure for gauge-fixed sub-space
 - ④ Action = $S + S_{\text{gauge-fixing}} + S_{\text{ghost}}$

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1. Quantum Mechanics

Stationary states in Q.M.

$$\Psi_E(x) \stackrel{?}{=} \int_{\text{in.cond.}}^{q(1)=x} \int \mathcal{D}[q] \mathcal{D}[\mathcal{N}] \mathcal{D}[\Pi] \mathcal{D}[\text{ghosts}] \exp [iS_{\text{tot.}}]$$
$$S_{\text{tot.}} = \int_0^1 \left(\frac{m}{2\mathcal{N}} \dot{q}^2 - \mathcal{N}V(x) + E\mathcal{N} + \Pi \dot{\mathcal{N}} + L_{\text{ghost}} \right) dt$$

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- Also, evaluating the action at the above $\mathcal{N}(t)$ gives²

$$S = \pm \int \sqrt{2m(E - V)} \dot{q}^2 \Rightarrow \text{Reparametrization invariance}$$

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- Fortunately, this can be reduced to a much simpler expression.

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Free particle: Q.M.

$$\begin{aligned}\Psi_E(x) &= \int_{-\infty}^{\infty} d\mathcal{N} \int_{m\dot{q}(0)=p}^{q(1)=x} \mathcal{D}[q] \exp[iS] \\ &\propto e^{ip \cdot x} \delta\left(\frac{p^2}{2m} - E\right)\end{aligned}$$

Free particle: scalar QFT

$$\begin{aligned}\psi_p(x) &= \int_{-\infty}^{\infty} d\mathcal{N} \int_{m\dot{q}(0)=p}^{q(1)=x} \mathcal{D}[q] \exp[iS] \\ &\propto \underbrace{e^{ip \cdot x} \delta(p^2 - m^2)}_{\text{mode-function with on-shell measure}}\end{aligned}$$

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- Before going into the details, contrast the real-time approach with

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- This is equivalent to taking the limit $i\mathcal{N} \rightarrow \infty$, **rather than integrating over $\mathcal{N}$** .
- So the question arises: *How are the real-time and Euclidean approaches related?*

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- Complex solutions: $q(t) = e^{-i\omega\mathcal{N}} [xe^{it\omega\mathcal{N}} - ix_0 \sin(\mathcal{N}(t-1)\omega)]$

SHO: the proptime integral

$$\Psi_E(x) = \int_{-\infty}^{\infty} d\mathcal{N} e^{i(E - \frac{\omega}{2})\mathcal{N}} \exp \left[\underbrace{-\frac{m\omega}{2} (x - x_0 e^{-i\omega\mathcal{N}})^2 + \frac{1}{2} x_0^2 e^{-2i\omega\mathcal{N}}}_{\text{Periodic}} \right]$$

- Integrand is periodic for $E = E_n = \omega \left(\frac{1}{2} + n \right)$

SHO: the propertime integral

$$\psi_{E_n}(x) = \oint d\mathcal{N} \underbrace{e^{in\omega\mathcal{N}} \exp \left[-\frac{m\omega}{2} (x - x_0 e^{-i\omega\mathcal{N}})^2 + \frac{1}{2} x_0^2 e^{-2i\omega\mathcal{N}} \right]}_{\equiv \mathcal{I}(e^{-i\omega\mathcal{N}})}$$

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- Integrand is periodic for $E = E_n = \omega \left(\frac{1}{2} + n \right)$
- Motivates the change of variable

$$z = e^{-i\omega\mathcal{N}} \Rightarrow d\mathcal{N} = \frac{i\omega^{-1}dz}{z}$$

Back to ground-state

- For $n = 0$, we get

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- For the SHO, the connection between real-time and Euclidean approaches is the Cauchy's residue theorem!
- Also, not that for a general n , the propertime integral generates the correct $\Psi_{E_n}(x)$ in the form of Cauchy's integral formula.

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- Bound state wave functions in QFT?

2. Schwinger effect

The Schwinger effect: intro

- The phenomenon of decay of QED vacuum in the presence of a constant external electromagnetic field via pair production.³

³F. Sauter(1931),W. Heisenberg and H. Euler (1936),J. Schwinger(1951) 

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The Schwinger effect: intro

- The phenomenon of decay of QED vacuum in the presence of a constant external electromagnetic field via pair production.³
- Robust non-perturbative prediction in QED.
- In this talk, we restrict to **scalar QED**.

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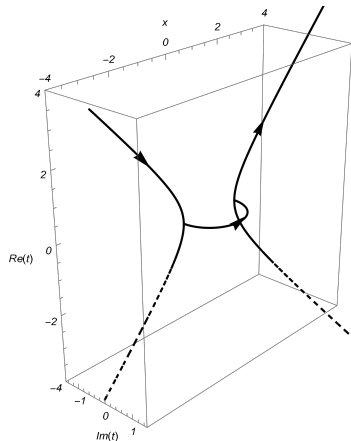
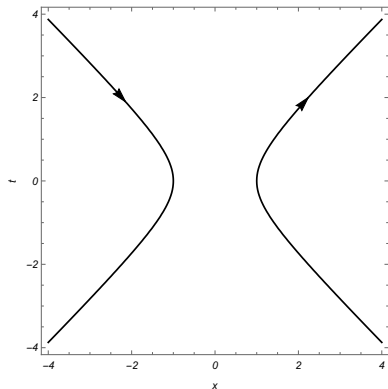
Worldline for Schwinger

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- Open worldline $\rightarrow$ particle spectrum [Barut and Duru (1990), Srinivasan and Padmanabhan (1990), K.R (2021), Esposti and Torgrimsson (2022)]

Heuristic picture⁴



⁴Srinivasan and Padmanabhan (1999)

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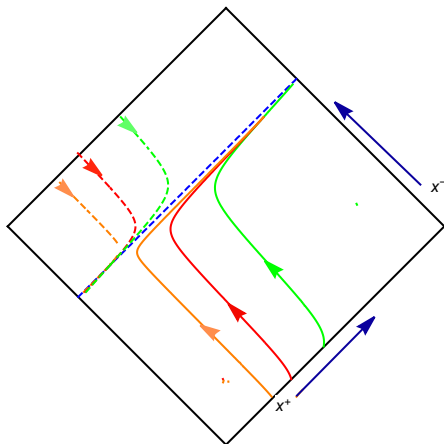
- In the conventional background-QFT approach, one works with positive(negative) energy solutions of the background Klein-Gordon equation.
- In most frequently used gauges, these mode functions involve parabolic cylinder functions.
- However, in a class of gauges, including the so-called lightfront gauge below, the solutions are simple functions.

$$eA_\mu = (eA_+, eA_-, eA_\perp) = (eEx^-, 0, 0) \quad ; \quad eE > 0$$

$$x^\pm = \frac{1}{\sqrt{2}}(x^0 \pm x^1) \quad ; \quad x^\perp = (x^2, x^3)$$

Classical physics

Particle and antiparticle solutions for fixed p .



KG equation and the solutions

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$$(\partial_\mu \partial^\mu + m^2)\phi + ieE \left(1 + 2x^- \frac{\partial}{\partial x^-}\right) \phi = 0$$

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$$\phi_p^+(x) = e^{-ip_+ x^+ - ip_\perp x^\perp} \underbrace{e^{-\left(\frac{1}{2} - i \frac{m^2 + p_\perp^2}{2eE}\right) \log\left(1 - eE \frac{x^-}{p_+}\right)}}_{\sum f_n(p; x) (eE)^n}$$

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- Note the 'horizon' at $x^- = p^+/(eE)$.

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$$\text{Opt. 2: } e^{-ip_+ x^+ - ip_\perp x^\perp} e^{-\left(\frac{1}{2} - i \frac{m^2 + p_\perp^2}{2eE}\right) \log\left(1 - \frac{x^-}{x_h^-} + i0^+\right)}$$

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- Option 1: Consistent with **classical intuition**.
- Option 2: **positive energy w.r.t x^-** .

First quantization

$$\phi_p^+(x) = \int_{-\infty}^{\infty} d\mathcal{N} \int_{\Pi(0)=p}^{q(1)=x} \mathcal{D}[q] e^{i\mathcal{S}_{\text{WL}}[q;\mathcal{N}]} \quad ; \text{ (Gauge fixed and exorcised)}$$
$$\mathcal{S}_{\text{WL}} = - \int_0^1 \left(\frac{\dot{x}^2}{4\mathcal{N}} + \mathcal{N} e A \cdot \dot{x} \right) d\tau \quad \underbrace{-p \cdot q(0)}_{\text{Boundary term}}$$

- Thanks to the quadratic action:

$$\phi_p^+(x) = \int_{-\infty}^{\infty} d\mathcal{N} \underbrace{e^{eE\mathcal{N}}}_{\text{fluctuation factor}} \times \underbrace{e^{i\mathcal{S}_{\text{cl.}}}}_{\text{semiclassical=exact}}$$

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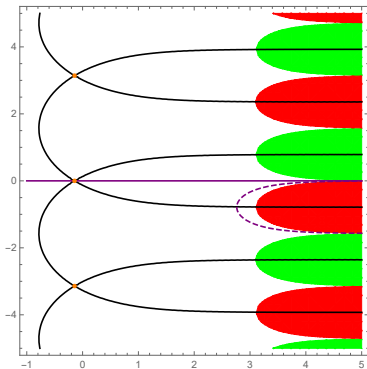
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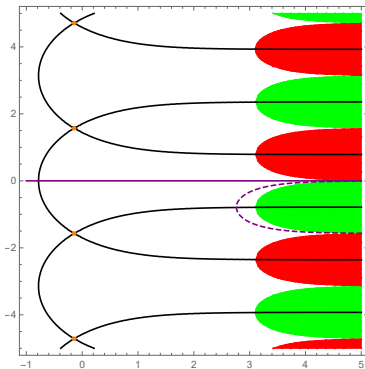
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- Note:** This is a sum of both *space-like* and time-like worldlines!
- On-shell solution is described by the saddle-point: $\partial_{\mathcal{N}_s} \mathcal{S}_{\text{cl.}} = 0$

Picard–Lefschetz



(a) Real saddle point exists.



(b) Real saddle point doesn't exist.

The mode function

- Case 1: $x^- < x_h^-$, Shift $\mathcal{N} = \mathcal{N}_s + \bar{\mathcal{N}}$

$$\phi_p^+(x) = e^{-ip_+ x^+ - ip_\perp x^\perp} e^{-\left(\frac{1}{2} - i \frac{m^2 + p_\perp^2}{2eE}\right) \log\left(1 - eE \frac{x^-}{p_+}\right)} \times$$
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- Case 2: $x^- > x_h^-$, Shift $\mathcal{N} \rightarrow \bar{\mathcal{N}} + \underbrace{\Re[\mathcal{N}_s] - i\pi/(2eE)}_{\text{instanton!}}$

$$\phi_p^+(x) = e^{-\frac{p_\perp^2 + m^2}{2eE}} e^{-i\frac{\pi}{2}} e^{-ip_+x^+ - ip_\perp x^\perp} e^{-\left(\frac{1}{2} - i\frac{m^2 + p_\perp^2}{2eE}\right) \log\left(eE\frac{x^-}{p_+} - 1\right)} \times$$

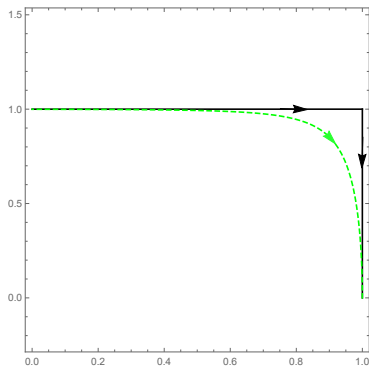
$$\underbrace{\int_{-\infty}^{\infty} d\bar{\mathcal{N}} \text{ function}(\bar{\mathcal{N}})}_{\text{constant}}$$

More on instanton

- Real time + complex $\mathcal{N}$ + gauge transformation = Complex time!

$$\left(q_{\text{cl.}}^-(\tau) - \frac{p_+}{eE}\right) = \left(x^- - \frac{p_+}{eE}\right) e^{2eE(\mathcal{N}_s - \tau)} \quad ; \quad 0 < \tau < 1$$

$$\left(q_{\text{cl.}}^-(\bar{\tau}) - \frac{p_+}{eE}\right) = \left(x^- - \frac{p_+}{eE}\right) e^{2eE\mathcal{N}_s - eE\bar{\tau}} \quad ; \quad \mathcal{N}_s\tau = \bar{\tau} \in \mathbb{C}$$



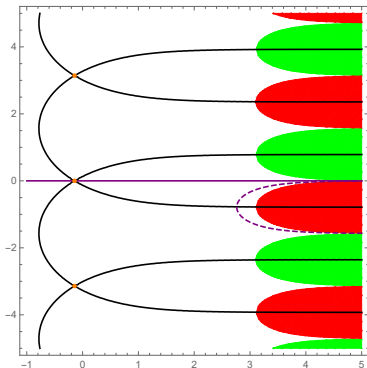
Back to $+i0^+$ vs $\Theta(\dots)$

- The *real-time* worldline approach naturally leads to positive-energy solution, hence the $+i0^+$ prescription for the $\log\left(1 - \frac{x^-}{x_h^-}\right)$

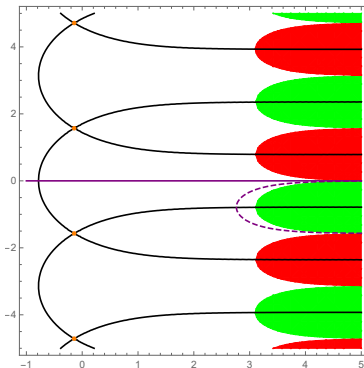
Back to $+i0^+$ vs $\Theta(\dots)$

- The *real-time* worldline approach naturally leads to positive-energy solution, hence the $+i0^+$ prescription for the $\log\left(1 - \frac{x^-}{x_h^-}\right)$
- How to get the other solution, with $\Theta(x_h^- - x^-)$?

Picard–Lefschetz (again)



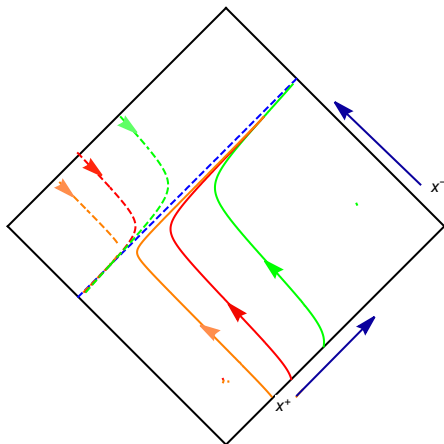
(a) Real saddle point exists.



(b) Real saddle point doesn't exist.

The classical-like solution

$$\int_{\infty}^{\infty - i\frac{\pi}{2eE}} d\mathcal{N} \int_{\Pi(0)=p}^{z(1)=x} \mathcal{D}[q] e^{i\mathcal{S}_{\text{WL}}[q, \mathcal{N}]} \propto \phi_p^+(x) \Theta(x_h^- - x^-)$$



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 - ① Amplitudes vs Bogoluibov co-efficients $\Rightarrow$ (contours)
 - ② Horizon: Schwinger and Hawking effect connection; W.L. approach?

$$\psi_E(\nu) \sim e^{-iE\nu} e^{4GMi \log\left(\frac{\nu_h - \nu}{\mu}\right)}$$

THANK YOU!