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Memory of Robinson-Trautman waves

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Plan of the talk

- ① Robinson-Trautman waves
- ② Memory in asymptotically flat spacetimes
- ③ Making RT waves asymptotically flat
- ④ Memory of RT waves
- ⑤ Application & discussion

in collaboration with Ali Seraj

① RT (1960) spherical waves

$$ds^2 = -2U du^2 + 2du dr + r^2 ds_2^2,$$

conf. flat 2d metric $ds_2^2 = -\frac{2}{p^2} d\zeta d\bar{\zeta}$

$$P(u, \zeta, \bar{\zeta}) = e^{-\varphi} \quad \boxed{\Delta_u \varphi = -\frac{1}{12M} \Delta_2^2 \varphi} \quad U = -r \Delta_u \varphi - \frac{1}{4} R_2 + \frac{M}{r}$$

Tod (1989) 2d Calabi flow $\Delta_u g_2 = \frac{1}{12M} \Delta_2 R_2 g_2$

Calabi (1983)

(1a) NP tetrad $l = J_r$ $m = J_u + U J_r$ $\bar{m} = r^{-1} P \bar{J}$ not round enough

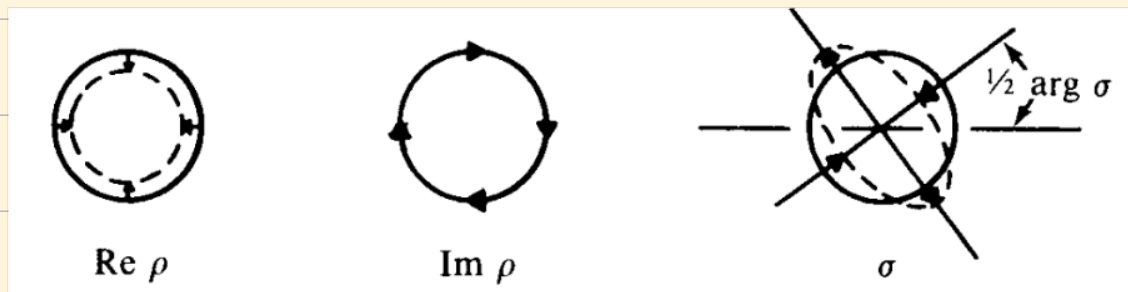
m : covariant derivative $\delta \eta^s = P^{t-s} \bar{J} (P^s \eta^s)$ $\Delta_{\bar{m}} \eta^0 = 2 + \bar{J} \eta^0$

not flat enough $\gamma = \frac{1}{2} J_u \varphi + \mathcal{O}(r^{-1})$, $v = \bar{J} J_u \varphi + \mathcal{O}(r^{-1})$

(1b) algebraically special $\psi_0 = \psi_1 = 0$

l : repeated null direction, generator of null geodesic congruence

twist-free $\rho = \bar{\rho}$ shear-free $\sigma = 0$ expansion $-\text{Re } \rho = \frac{1}{r}$



Sachs (1961)

Chandrasekhar (1983)

FIG. 1. The geometrical interpretation of the optical scalars in terms of the effect of propagation of a small circle perpendicular to the beam.

(1c) trivial constant solution S^2 $P_0 = \frac{1}{\sqrt{2}}(1 + \gamma \bar{\gamma})$ $\varphi_0 = -\ln P_0$

space-time solution Schwarzschild black hole

deviation $\Phi = \varphi - \varphi_0$

$$-3M e^{4\Phi} J_u \Phi = \left[(t_0 \bar{t}_0)^2 + (t_0 \bar{t}_0) \right] \Phi$$

$$-2 t_0 \Phi \bar{t}_0 \Phi - 2 \bar{t}_0 \Phi t_0 \bar{t}_0 t_0 \Phi - 2 t_0 \Phi \bar{t}_0 t_0 \bar{t}_0 \Phi - 2 (t_0 \bar{t}_0 \Phi)^2 + 4 t_0 \Phi \bar{t}_0 \Phi t_0 \bar{t}_0 \Phi$$

linearization $\Phi = \epsilon \Phi_1$

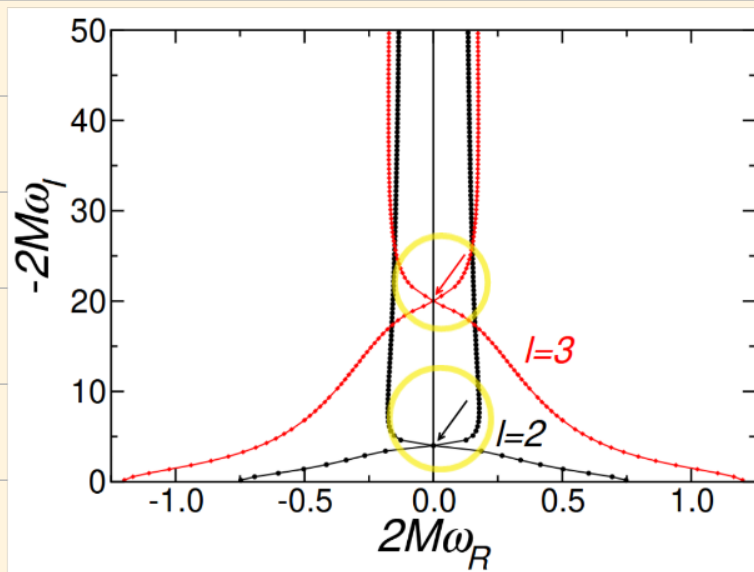
$$-3M J_u \Phi_1 = \left[(t_0 \bar{t}_0)^2 + (t_0 \bar{t}_0) \right] \Phi_1$$

$$\Phi_1 = a^{jm}(u) Y_{jm} + c.c.$$

$$a^{jm}(u) = c^{jm}, \quad j \leq 1 \quad a^{jm}(u) = c^{jm} e^{-\frac{\omega_j}{M} u}, \quad j > 1$$

algebraically special quasi-normal modes

$$\omega_j = \frac{(j-1)j(j+1)(j+2)}{12}$$



Couch & Newman (1973)

Chandrasekhar (1984)

Qi & Schutz (1993)

Berti et al. (2009)

② asymptotically flat: $\rho' = \bar{\rho}'$ $\epsilon' = \pi' - \kappa' = 0$ $\bar{b}' = \bar{\alpha}' + \beta'$

7 Bondi frame $l' = J_{r'}$ $m' = J_{u'} - \frac{1}{2} J_{r'} + \mathcal{O}(r'^{-1})$

coordinate conditions $m' = r'^{-1} (\bar{\rho}_0) \bar{J}' + \mathcal{O}(r'^{-1})$

r' dependence $D' \rho' = \rho'^2 + \nabla' \bar{\sigma}'$ $D' \nabla' = 2 \rho' \nabla' + \bar{\Psi}'_0$

$\bar{\Psi}'_0 = \mathcal{O}(r'^{-5})$ $-\rho' = +\frac{1}{r'} + \mathcal{O}(r'^{-3})$ $\nabla' = \nabla'^0 r'^{-2} + \mathcal{O}(r'^{-2})$

incoming radiation

expansion

shear

news $\lambda'^0 = J_{u'} \bar{\sigma}'^0$ $\psi_3'^0 = -\bar{t}'_0 \bar{\lambda}'^0$ $\psi_4'^0 = -J_{u'} \bar{\lambda}'^0$

evolution $J_{u'} \psi_2'^0 = \bar{t}'_0 \psi_3'^0 + \nabla'^0 \psi_4'^0$

(non-linear) memory Christodoulou (1991) [Fraundtner (1992)]

$$\int_{u'} \dot{t}_0^{12} \bar{\nabla}^{10} = - \int_{u'} \left[\psi_2^{10} + \nabla^{10} \bar{\Delta}^{10} \right] + \boxed{\Delta^{10} \bar{\Delta}^{10}} \quad (x)$$

Bondi mass aspect $\mathcal{M}'_B = -\frac{1}{2}(\psi_2^{10} + \nabla^{10} \bar{\Delta}^{10} + \text{c.c.})$

conservation laws

$$\int_{S^2} (x) \cdot \bar{\Delta}^{10}_{jm} \begin{cases} j \leq 1 \\ j \geq 2 \end{cases} \rightarrow \frac{1}{2}(\Delta^{10} \bar{\nabla}^{10}_{jm} + \text{c.c.}) = \Delta(\dot{t}_0^{12} \mathcal{M}'_B)_{jm} + \int_{u'_i}^{u'_f} d\sigma (\dot{t}_0^{12} |\Delta^{10}|^2)_{jm}$$

generalized mass aspect $\mathcal{M}'_G = -(\psi_2^{10} + \nabla^{10} \bar{\Delta}^{10} + \dot{t}_0^{12} \bar{\nabla}^{10}) = \overline{\mathcal{M}'_G}$

$$(x) \quad \boxed{\int_{u'} \mathcal{M}'_G = - |\Delta^{10}|^2} \quad \text{mass-loss for generalized aspect}$$

③ making RT flat • generalized solution space $P(u, \zeta, \bar{\zeta})$

• compute symmetry group BMS + (complex) Weyl trsf.

• work out action on solution space G.B. & Troessent (2016)

subleties combined coord. trsf + frame rotation, 2nd order

$$x^\mu = x^\mu(x'^\nu) \quad e_a'^{\nu} \frac{\partial x^\mu}{\partial x'^\nu} = \Lambda_a^\mu e_\mu^\nu$$

finite Weyl trsf $P(u, \zeta, \bar{\zeta}) \rightarrow P_0(\zeta', \bar{\zeta}')$

Bondi time $u = u_0(u', \zeta', \bar{\zeta}') + O(r'^{-1})$ inverse $u'_0(u, \zeta, \bar{\zeta}) = \int_0^u \frac{P}{P_0} dv$

subleading class II rotation $b = r'^{-1} b_0 + O(r'^{-2})$ $b_0 = \zeta'_0 u'_0 /$

Results for RT

$$\sigma'^0 = j_0'^2 u'_0$$

$$\psi_4'^0 = \dots \quad \psi_3'^0 = \dots \quad \psi_2'^0 = -[e^{3\Phi} M + 2b_0 \bar{\psi}_3'^0 + b_0^2 \bar{\psi}_4'^0] \quad \bar{\psi}_1'^0 = \dots$$

$$\bar{\psi}_0'^0 = -[4b_0 \psi_0'^0 + 6b_0^2 \psi_2'^0 + 4b_0^3 \psi_3'^0 + b_0^4 \bar{\psi}_4'^0] \quad \text{Adamo et al. (2009)}$$

$$M'_\alpha = [e^{3\Phi} M - [(j_0 \bar{j}_0)^2 + (j_0 \bar{j}_0)] u'_0] \quad \text{Tafel (2000)}$$

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$$\sigma'^0 = [-\bar{j}_0^2 \Phi + (\bar{j}_0 \Phi)^2]$$

back to natural RT coordinates

$$\ln M_\alpha = -e^{-\Phi} |j_0^2 \Phi - (j_0 \Phi)^2|^2$$

$\Leftrightarrow$ RT equation

decreasing $M_\alpha(u_f) \leq M_\alpha(u_i) \quad u_f > u_i$

useful for non-linear convergence to Schwarzschild?

Rendall, Schmidt, Singleton, Chrusciel (1988-92)

⑤ No news

$$\ddot{\varphi} + (\dot{\varphi})^2 = e^{-2\Phi} [\dot{\varphi}_0^2 \bar{\Phi} - (\dot{\varphi}_0 \bar{\Phi})^2] = e^{-2\varphi} [\bar{J}^2 \varphi - (\bar{J}\varphi)^2] = 0 \quad + c.c.$$

7 parameter family of constant "vacuum" solutions

$$\varphi_\nu = -\ln \sqrt{\frac{\mu}{4}} - \ln [|a\zeta + b|^2 + |c\bar{\zeta} + d|^2], \quad a, b, c, d \in \mathbb{C} \quad ad - bc = 1$$

$$J\bar{J}\varphi_\nu = -\frac{\mu}{4} e^{2\varphi_\nu} \quad \varphi_0 = \varphi_\nu (\mu=2, a=1=d, b=0=c)$$

vacuum sector of Liouville $S = \int d\zeta d\bar{\zeta} \left[\frac{1}{2} J\phi_L \bar{J}\phi_L - \frac{\mu}{2} e^{\phi_L} \right] \quad \phi_L = \frac{1}{2}\varphi_\nu$

⑥ Summary

- explicit map of RT to asymptotically flat
- expressions for news and memory
- new (?) way of writing RT equation
- success of flat holography?

RT solutions with $\Lambda \neq 0$ $U \rightarrow U + \frac{\Lambda r^2}{6}$

same effective 2d dynamics Bakas & Skenderis (2014)