

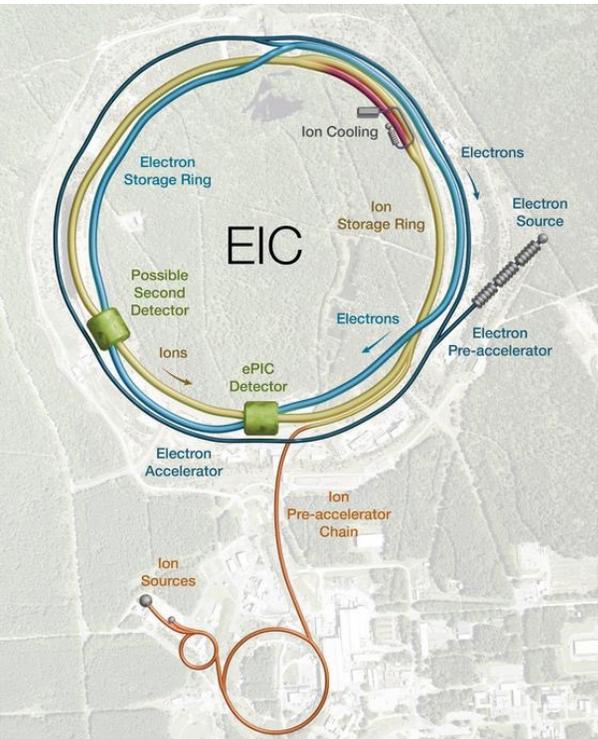
Amplitudes to shockwaves in QCD and gravity in Regge asymptotics

Raju Venugopalan
BNL and CFNS, Stony Brook

Celestial Holography satellite meeting, April 15-16, 2025

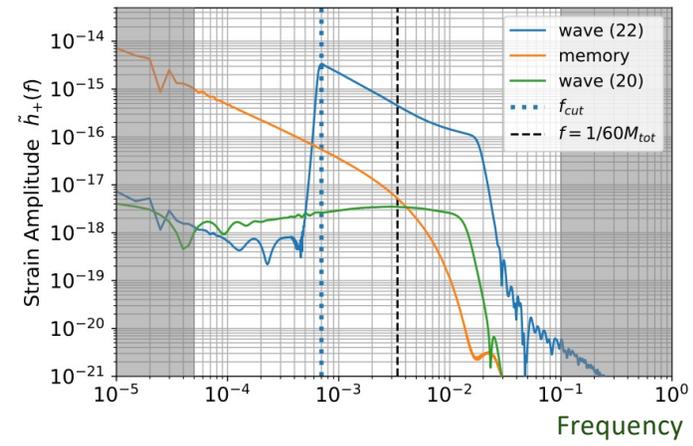
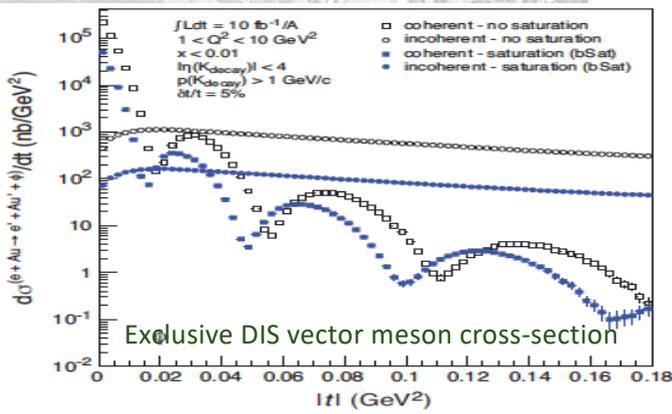
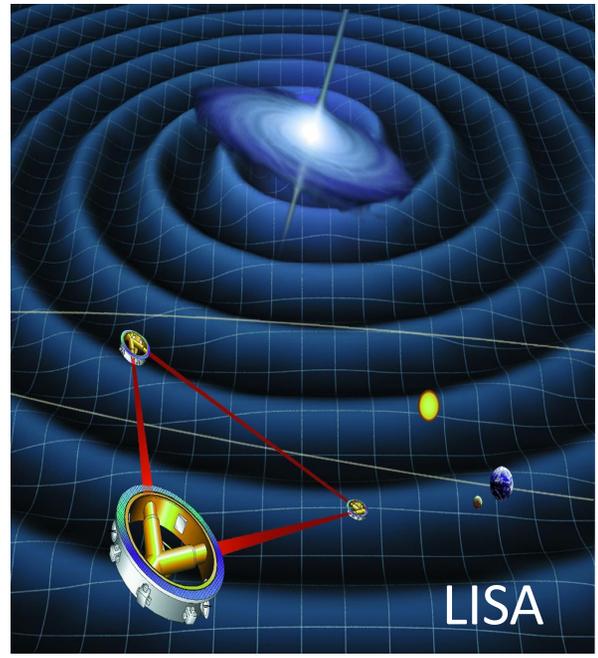


Based in part on recent published work
(arXiv:2311.03463, 2312.03507, 2312.11652, 2406.10483, + **2505.XXXXX in preparation**)
with Himanshu Raj (Simons Confinement+ QCD Strings Collaboration Fellow at Stony Brook)



Color memory

GR memory



COLOR MEMORY

Monica Pate, Ana-Maria Raclariu and Andrew Strominger

Abstract

A transient color flux across null infinity in classical Yang-Mills theory is considered. It is shown that a pair of test ‘quarks’ initially in a color singlet generically acquire net color as a result of the flux. A nonlinear formula is derived for the relative color rotation of the quarks. For weak color flux the formula linearizes to the Fourier transform of the soft gluon theorem. This color memory effect is the Yang-Mills analog of the gravitational memory effect.

arXiv:1707.08016

MEASURING COLOR MEMORY IN A COLOR GLASS CONDENSATE AT ELECTRON-ION COLLIDERS

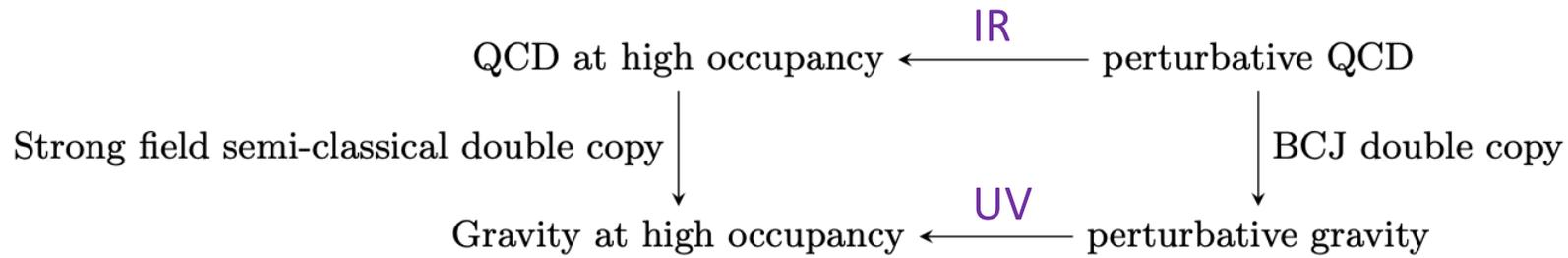
Adam Ball¹, Monica Pate¹, Ana-Maria Raclariu¹, Andrew Strominger¹ and Raju
Venugopalan²

Abstract

The color memory effect is the non-abelian gauge theory analog of the gravitational memory effect, in which the passage of color radiation induces a net relative SU(3) color rotation of a pair of nearby quarks. It is proposed that this effect can be measured in the Regge limit of deeply inelastic scattering at electron-ion colliders.

arXiv:1805.12224

Double Copy: gluon \rightarrow gravitational radiation in shockwave collisions



Monteiro, O'Connell, White, arXiv:1410.0239
Goldberger, Ridgeway, arXiv:1611.03493

Bern, Carrasco, Johansson,
arXiv: 1004.0476

Road map of my talk

I) Why can we measure color memory?

A brief DIS primer. The BFKL equation in multi-Regge asymptotics.

Breakdown of the OPE, classical lumps, their quantum descendants.

Color memory in the Color Glass Condensate.

Precision measurements of the CGC at the Electron-Ion Collider.

II) QCD-Gravity double copy in Regge asymptotics

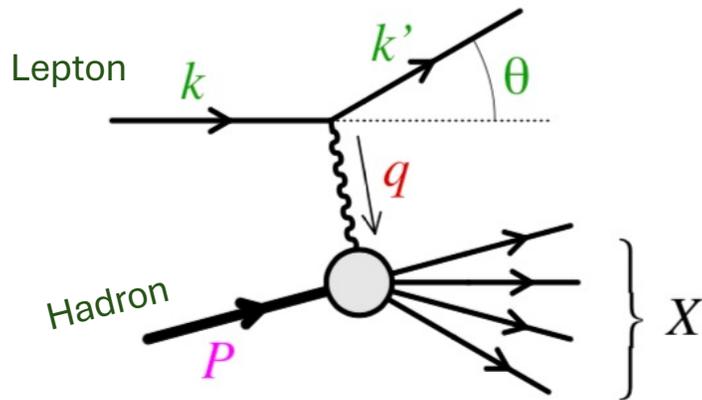
$2 \rightarrow N$ scattering, the Lipatov double copy and reggeization.

From amplitudes to shockwaves.

The Lipatov vertex and shockwave propagators.

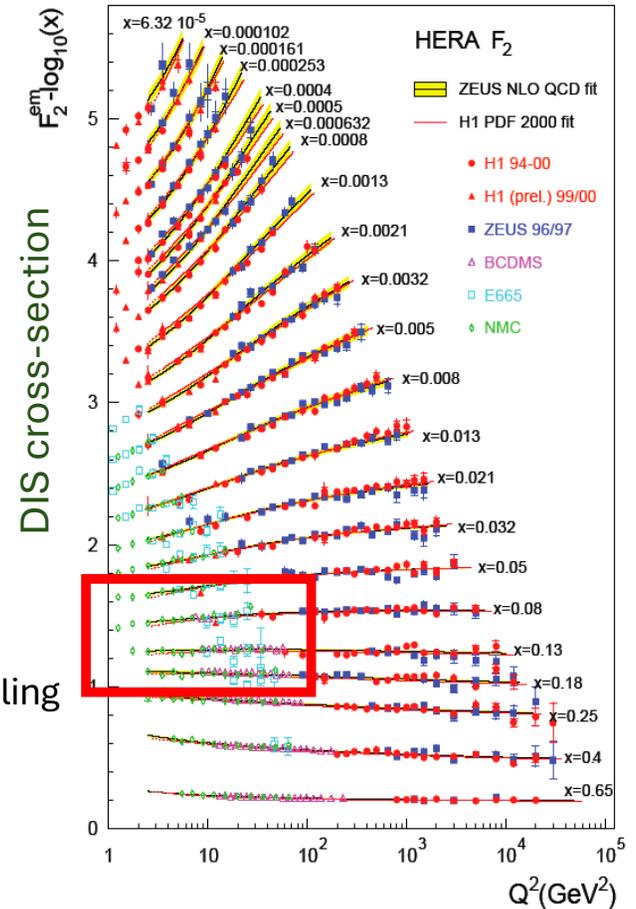
Some consequences.

DIS and the QCD revolution



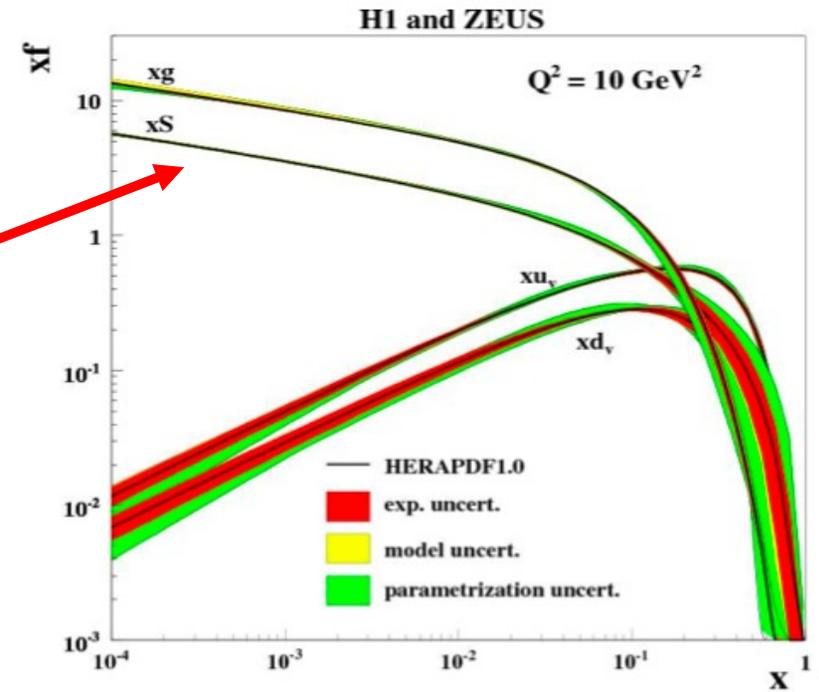
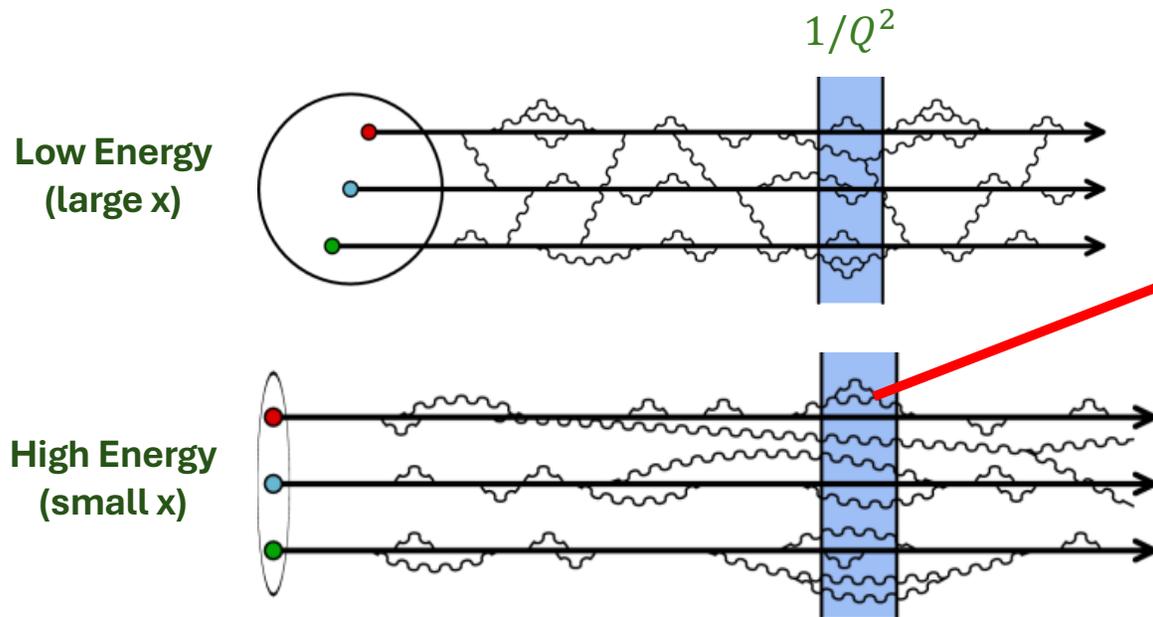
The discovery of Bjorken scaling in DIS led to the parton model, asymptotic freedom, and QCD, shattering existing paradigms...

Bjorken scaling



Bj in memoriam:
<https://indico.slac.stanford.edu/event/9148/timetable/>

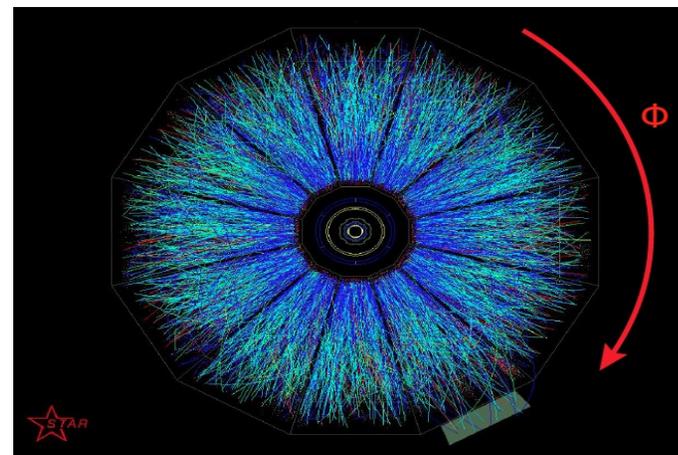
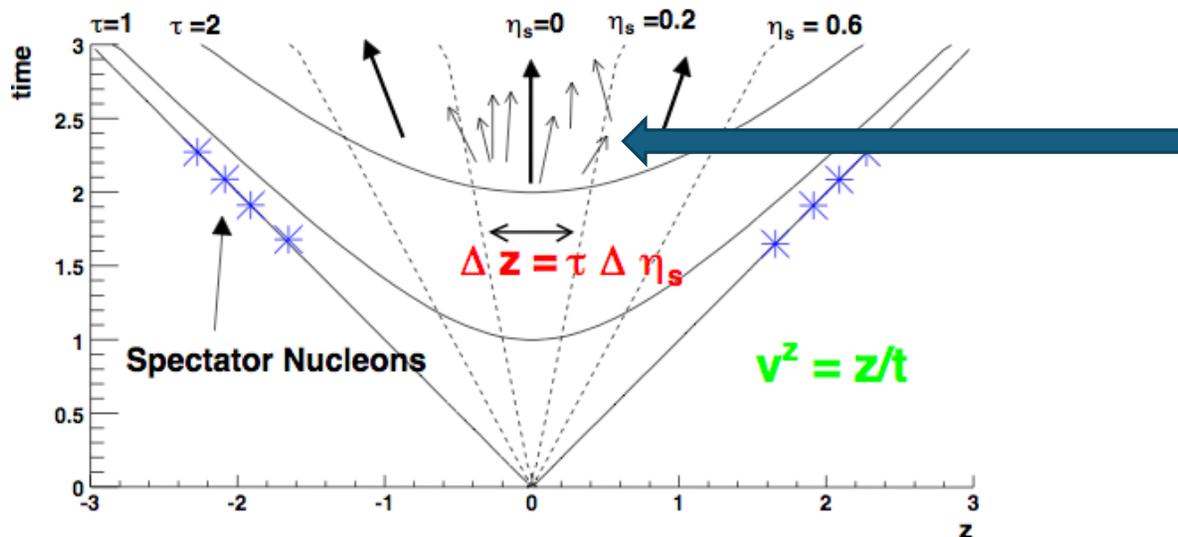
Spacetime picture of wee partons in a hadron



As the proton is boosted, “parton” fluctuations live longer -- released as Bremsstrahlung

Suppression in coupling compensated by large phase space for soft glue: $\alpha_s \ln\left(\frac{1}{x}\right) \sim 1$

Spacetime picture of a high energy hadron-hadron collision



$$\eta_s = \frac{1}{2} \text{Ln} \left(\frac{t+z}{t-z} \right) \approx Y$$

Fast “valence” partons populate fragmentation regions at large rapidities – “leading particle” effect

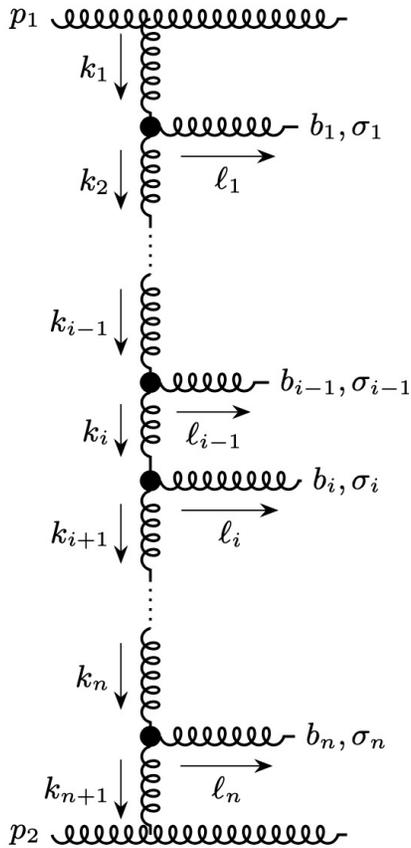
Slow “wee” partons populate central rapidities (mostly gluons and sea-quark pairs)
 – they create a Quark-Gluon-Plasma

QCD thermalization: Ab initio approaches and interdisciplinary connections

Jürgen Berges, Michal P. Heller, Aleksas Mazeliauskas, RV

Rev. Mod. Phys. **93**, 035003 (2021)

BFKL: 2 → N QCD amplitudes in Regge asymptotics



Compute multiparticle in multi-Regge kinematics of QCD:

$$y_0^+ \gg y_1^+ \gg y_2^+ \gg \dots \gg y_N^+ \gg y_{N+1}^+ \quad \text{with} \quad \mathbf{k}_i \simeq \mathbf{k}$$

BFKL ladder is ordered in rapidity . Produced partons are wee in longitudinal momentum(“slow”) but hard in transverse momentum – weak coupling Regge regime of QCD

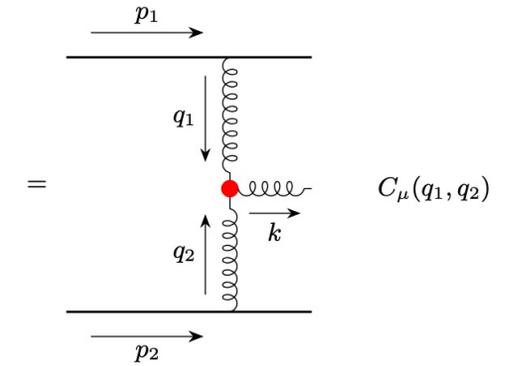
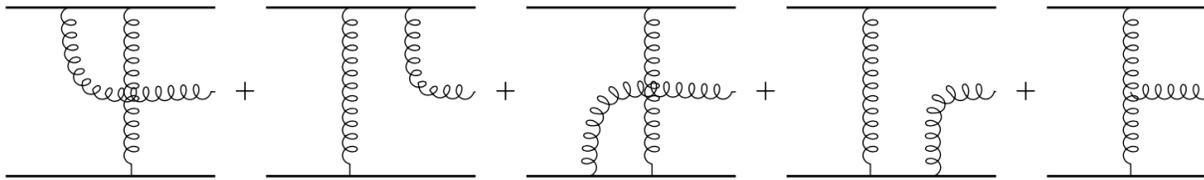
RG description rapidity of evolution given by the BFKL Hamiltonian
Very rapid growth of the amplitude with energy

$$A(s,t) = s^{\alpha(t)} \quad \text{with} \quad \alpha(t) = \alpha_0 + \alpha' |t| \quad \text{BFKL pomeron}$$

BFKL: Balitsky-Fadin-Kuraev-Lipatov (1976-1978)

BFKL: Building blocks

Lipatov effective vertex:

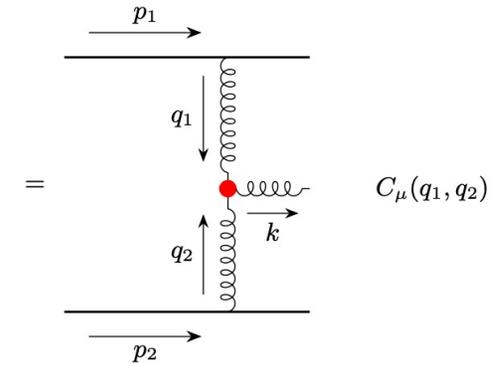
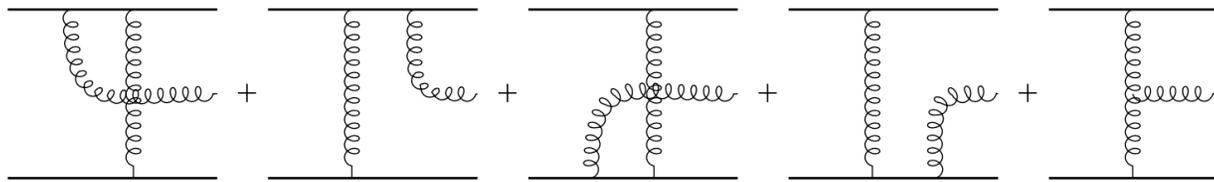


$$C_\mu(\mathbf{q}_1, \mathbf{q}_2) \simeq -\mathbf{q}_{1\mu} + \mathbf{q}_{2\mu} + p_{1\mu} \left(\frac{p_2 \cdot k}{p_1 \cdot p_2} - \frac{q_1^2}{p_1 \cdot k} \right) - p_{2\mu} \left(\frac{p_1 \cdot k}{p_1 \cdot p_2} - \frac{q_2^2}{p_2 \cdot k} \right)$$

Gauge covariant, satisfies $k_\mu C^\mu = 0$

BFKL: Building blocks

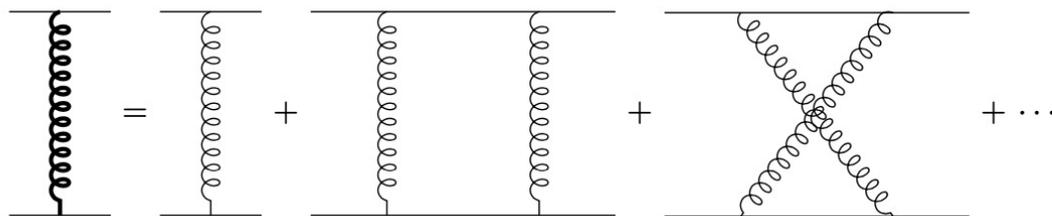
Lipatov effective vertex:



$$C_\mu(\mathbf{q}_1, \mathbf{q}_2) \simeq -\mathbf{q}_{1\mu} + \mathbf{q}_{2\mu} + p_{1\mu} \left(\frac{p_2 \cdot k}{p_1 \cdot p_2} - \frac{q_1^2}{p_1 \cdot k} \right) - p_{2\mu} \left(\frac{p_1 \cdot k}{p_1 \cdot p_2} - \frac{q_2^2}{p_2 \cdot k} \right)$$

Gauge covariant, satisfies $k_\mu C^\mu = 0$

Reggeized gluon:



$$\frac{1}{t_i} \rightarrow \frac{1}{t_i} e^{\alpha(t_i)(y_{i-1} - y_i)}$$

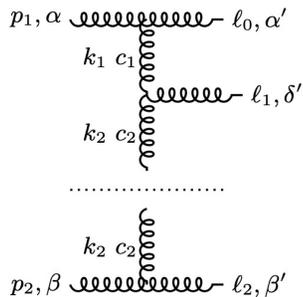
$$\alpha(t) = \alpha_s N_c t \int \frac{d\mathbf{k}}{(2\pi)^2} \frac{1}{\mathbf{k}^2 (\mathbf{q} - \mathbf{k})^2}, \quad t = -\mathbf{q}^2$$

Some features of the BFKL equation and its solution

Key to the construction of the BFKL ladder are multi-Regge asymptotics and dispersive techniques

Example: $2 \rightarrow 3$ amplitude

Building block is the $2 \rightarrow 2$ Born amplitude and three gluon vertex $\mathcal{A}_{2 \rightarrow 2, p_1+p_2 \rightarrow \ell_0+\ell_1}^{\alpha\alpha'\beta\beta'} = \Gamma_{p_1\ell_0}^{\alpha\alpha'c} \frac{s}{t} \Gamma_{p_2\ell_1}^{\beta\beta'c}$  3g vertex



For $2 \rightarrow 3$, pole contribution in k_2 can be expressed in terms of the $2 \rightarrow 2$ amplitude

$$P_{k_2} \mathcal{A}_{2 \rightarrow 2+1}^{\alpha\alpha'\beta\beta'\delta'} = \mathcal{A}_{p_1+(-k_2) \rightarrow \ell_0+\ell_1}^{\alpha\alpha'c_2\delta'} \frac{s}{k_2^2} \Gamma_{p_2\ell_2}^{\beta\beta'c_2} \longrightarrow -ig \frac{s}{k_1^2 k_2^2} \Gamma_{p_1\ell_0}^{\alpha\alpha'c_1} \Gamma_{p_2\ell_2}^{\beta\beta'c_2} f^{c_2\delta'c_1} \times \text{kinematic factors}$$

The k_1 pole can be expressed similarly allowing one to reconstruct the full amplitude

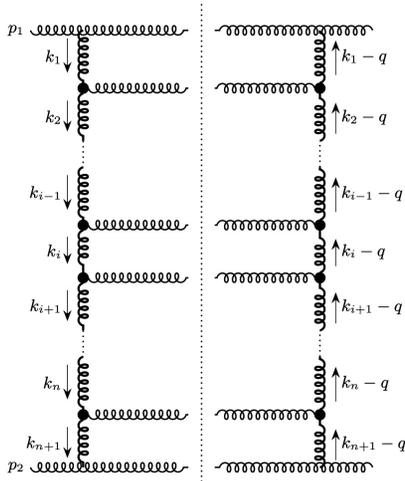
$$\mathcal{A}_{2 \rightarrow 2+1}^{\alpha\alpha'\beta\beta'\delta'} = ig \frac{s}{k_1^2 k_2^2} \Gamma_{p_1\ell_0}^{\alpha\alpha'c_1} \Gamma_{p_2\ell_2}^{\beta\beta'c_2} f^{c_1\delta'c_2} C_\nu(k_1, k_2) \epsilon^\nu(\ell_1)$$

 Lipatov vertex

This process is iterated to all orders to construct the BFKL ladder – the same procedure will apply in gravity...

2 → N + 2 amplitude in the Regge limit: the BFKL equation

BFKL Pomeron: compound color singlet state of two reggeized gluons



The imaginary part of this 2 → N + 2 amplitude simplifies greatly in Mellin space

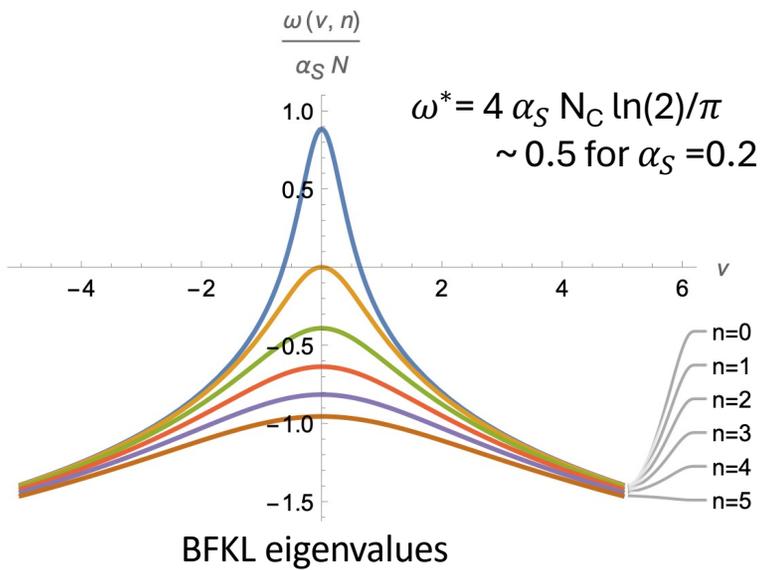
$$\mathcal{M}_\ell(\mathbf{q}^2) \equiv \int_1^\infty d\left(\frac{s}{\mathbf{k}^2}\right) \frac{\text{Im} \mathcal{A}_{2 \rightarrow 2}^{\mu\mu'\nu\nu'}(s, t)}{\mathcal{A}_0^{\mu\mu'\nu\nu'}(s, t)} \left(\frac{s}{\mathbf{k}^2}\right)^{-\ell-1}$$

$$\text{with } \mathcal{M}_\ell(\mathbf{q}^2) = 2\pi \mathbf{q}^2 \alpha_s N_c^2 (N_c^2 - 1) \int \frac{d^2 \mathbf{k}}{(2\pi)^2} \frac{1}{\mathbf{k}^2 (\mathbf{q} - \mathbf{k})^2} f_\ell(\mathbf{k}, \mathbf{q})$$

where $f_\ell(\mathbf{k}, \mathbf{q})$ satisfies the BFKL integral equation

$$(\ell - \alpha(\mathbf{k}^2) - \alpha((\mathbf{q} - \mathbf{k})^2)) f_\ell(\mathbf{k}, \mathbf{q}) = 1 - 2\alpha_s N_c \int \frac{d^2 \mathbf{k}'}{(2\pi)^2} \frac{f_\ell(\mathbf{k}', \mathbf{q})}{\mathbf{k}'^2 (\mathbf{q} - \mathbf{k}')^2} \left(\mathbf{q}^2 - \frac{\mathbf{k}^2 (\mathbf{q} - \mathbf{k}')^2 + \mathbf{k}'^2 (\mathbf{q} - \mathbf{k})^2}{(\mathbf{k} - \mathbf{k}')^2} \right)$$

2 → N + 2 amplitude in the Regge limit: the BFKL equation



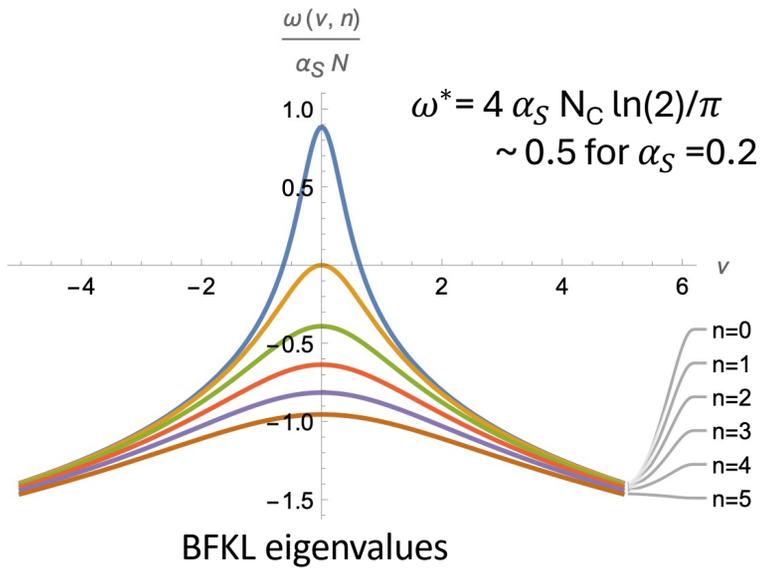
Solved as a simple eigenvalue equation as a function of Fourier conjugate variables -conformal spin ν and azimuthal variable n

Performing the inverse Mellin transform, one obtains

$$\sigma(s) \sim s^{\omega^*} = s^{0.5}$$

a much faster growth than the $\text{Ln}^2(s)$ predicted by Froissart...

2 → N + 2 amplitude in the Regge limit: the BFKL equation



This so-called LLx (leading log in x) result has been extended to NLLx accuracy.

Excellent review of state-of-the art:
 Del Duca, Dixon, arXiv:2203.13026

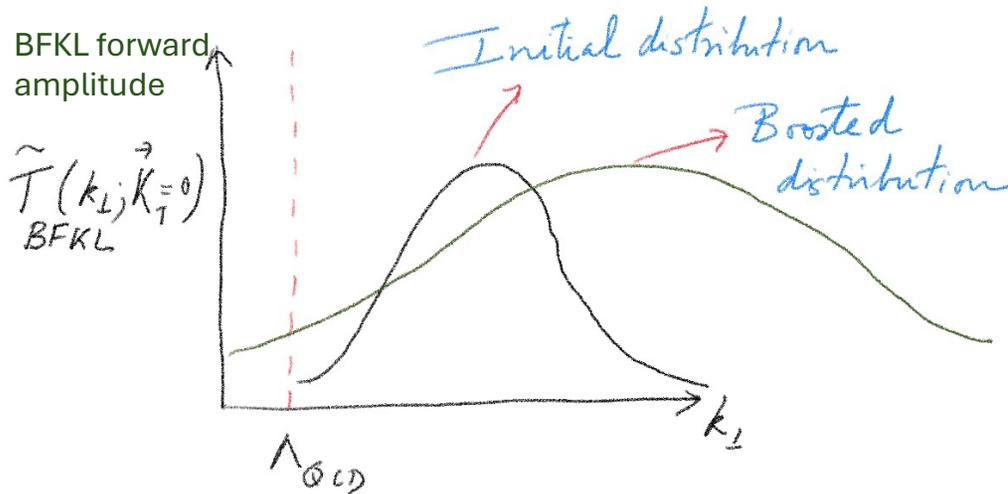
After much sophisticated analysis, this gives

$$\sigma(s) \sim s^{0.3} \text{ - in reasonable agreement with HERA data}$$

Not the full story...

and only a preview to a richer, many-body picture

BFKL: infrared diffusion and gluon saturation



For a fixed large Q^2 there is an $x_0(Q^2)$ such that below x_0 the OPE breaks down...

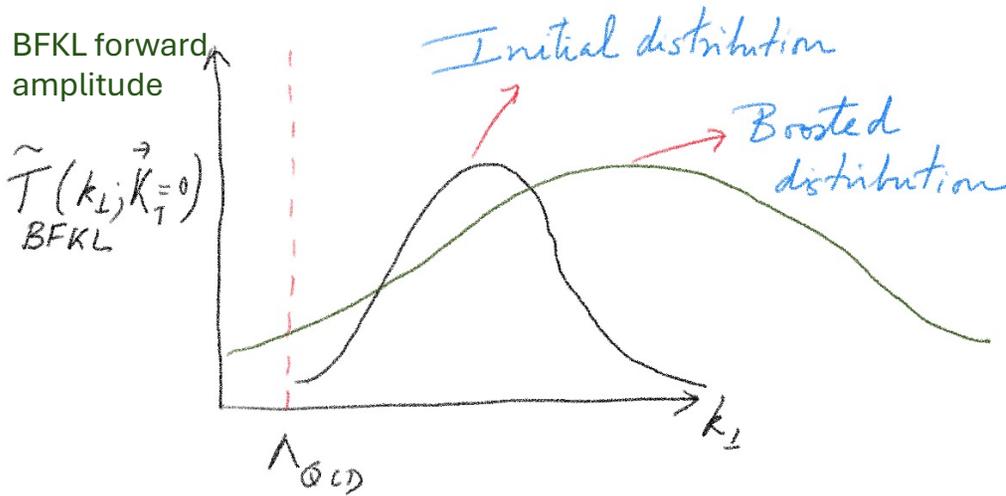
significant nonperturbative corrections in the leading twist coefficient and anomalous dimension functions due to diffusion of gluons to small values of transverse momentum.

A. H. Mueller, PLB 396 (1997) 251

NLLx BFKL does not cure infrared diffusion

Gluon saturation cures infrared diffusion

BFKL: infrared diffusion and gluon saturation



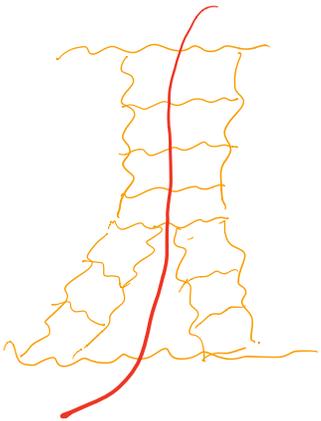
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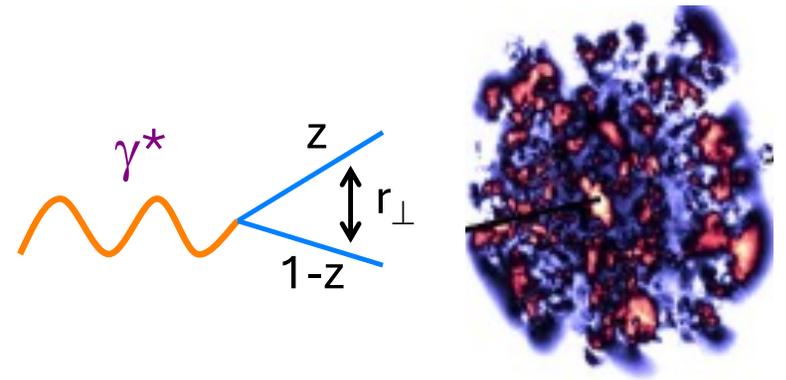
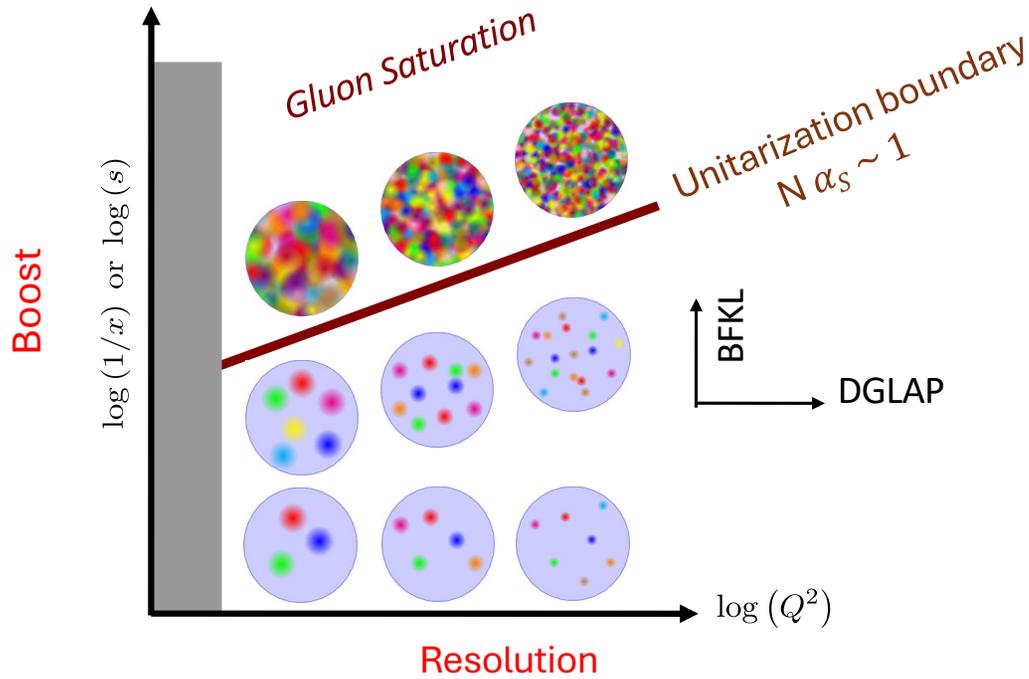
Gluon saturation cures infrared diffusion



+ other higher twist cuts of $O(1)$ when gluon occupancy $N \equiv \frac{xG_A(x, Q_S^2)}{2(N_c^2 - 1)\pi R_A^2 Q_S^2} = \frac{1}{\alpha_S(Q_S)}$

Classicalization when $\alpha_S(Q_S) \ll 1$ for saturation scale $Q_S \gg \Lambda_{QCD}$

Maximal packing of gluons: Gluon saturation



Color screened wee partons
 live on surface of sphere of radius
 $1/k^+ \sim 1/k_{\perp} \sim 1/Q_S(x, b)$

Emergent dynamics of semi-classical lumps that unitarize the cross-section described by the saturation scale $Q_S(x, b)$.

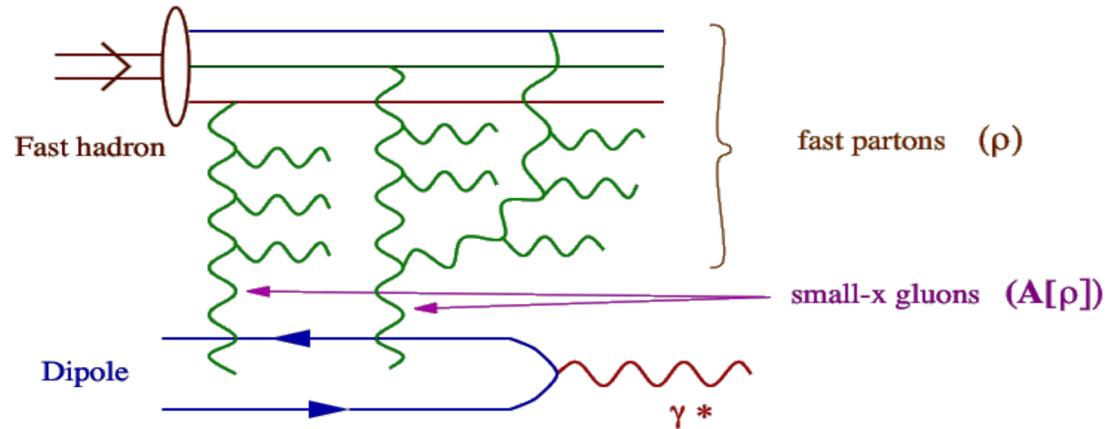
Due to asymptotic freedom, the many-body strong field dynamics of these lumps can be computed in weak coupling

Color Glass Condensate: classical EFT for Regge asymptotics

Born-Oppenheimer separation between fast and slow light-front modes (Galilean EFT)

Large x (P^+) modes: static, strong ($\sim 1/g$) color sources ρ^a

Small x ($k^+ \ll P^+$) modes: fully dynamical gauge fields A_μ^a



$$\mathcal{Z}[j] = \int [d\rho] W_{\Lambda^+}[\rho] \left\{ \frac{\int^{\Lambda^+} [dA] \delta(A^+) e^{iS_{\Lambda^+}[A,\rho] - \int j \cdot A}}{\int^{\Lambda^+} [dA] \delta(A^+) e^{iS_{\Lambda^+}[A,\rho]}} \right\}$$

$W_{\Lambda^+}[\rho]$: nonpert. gauge inv. weight functional defined at initial $x_0 = \Lambda^+ / P^+$

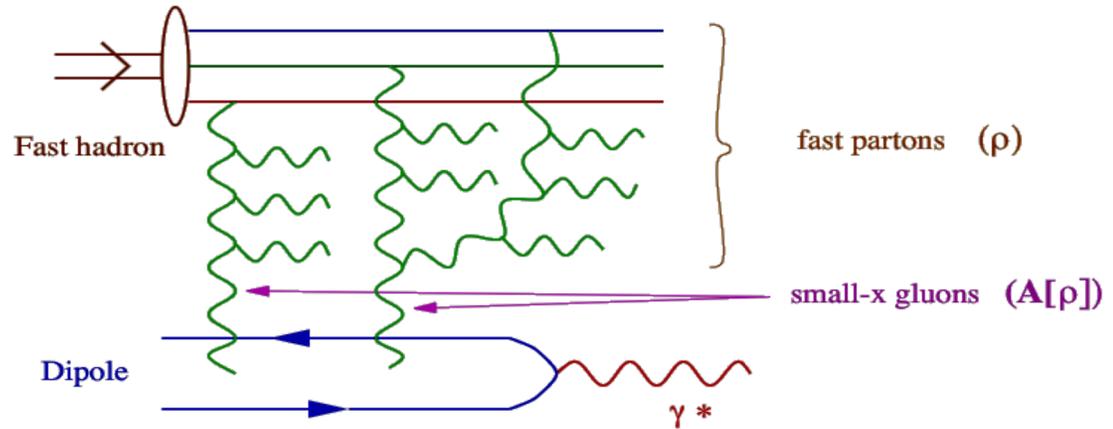
$S_{\Lambda^+}[A, \rho]$: Yang-Mills action + gauge-inv. coupling of sources to fields (Wilson line)

Color Glass Condensate: classical EFT for Regge asymptotics

Born-Oppenheimer separation between fast and slow light-front modes (Galilean EFT)

Large x (P^+) modes: static, strong ($\sim 1/g$) color sources ρ^a

Small x ($k^+ \ll P^+$) modes: fully dynamical gauge fields A_μ^a



Explicit construction for large nuclei (= large number of coherent sources of color charge at small x)

$$W_{\Lambda^+}[\rho] \rightarrow \int [d\rho] \exp \left(- \int d^2x_\perp \left[\frac{\rho^a \rho^a}{2\mu_A^2} - \frac{d_{abc} \rho^a \rho^b \rho^c}{\kappa_A} \right] \right)$$

Pomeron configurations

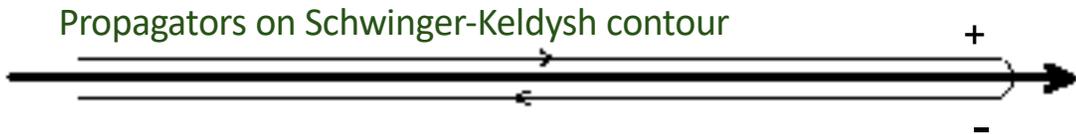
Odderon configurations

For $A \gg 1$, $\mu_A^2 \sim Q_S^2 \propto A^{1/3} \gg \Lambda_{QCD}^2$
 weak coupling EFT for large parton densities!

General all-order formalism: Cutkosky's rules in strong fields

$$2 \operatorname{Im} \sum_{\text{conn.}} V =$$

connected vacuum graphs in $\lambda\phi^3$



Well-known example: Schwinger pair production in strong field QED

Simple understanding of "AGK cutting rules" of Reggeon Field Theory:
 combinatorics of cut and uncut sub-graphs contributing to a given multiplicity

AGK: Abramovsy, Gribov, Kancheli

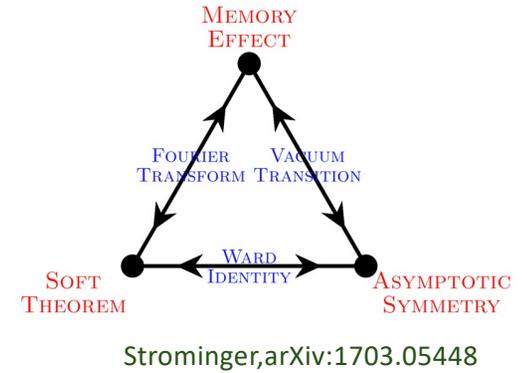
- Very general consequence of unitarity in strong fields
- Independent of language of Pomerons and Reggeons

Color memory in the CGC

Static Yang-Mills shockwave wave solution in LC gauge

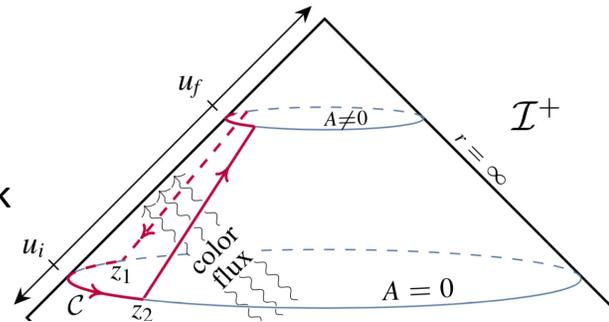
$$A_i = 0 \quad \Big| \quad x^- = 0$$

$$A_i = -\frac{1}{ig} U \partial_i U^\dagger$$



Transverse dynamics can be mapped on to celestial sphere at null infinity:

Kick Q_5 suffered by dipole crossing shock is the color memory effect



$$(r, u, z, \bar{z}) \rightarrow (\lambda r, \lambda^{-1} u, \lambda^{-1} z, \lambda^{-1} \bar{z})$$

$$\text{Map: } x^+ = \sqrt{2r}, \quad x^- = \frac{1}{\sqrt{2}}(u + rz\bar{z}),$$

$$x^1 + ix^2 = 2rz$$

For $\lambda \rightarrow \infty$:

$$r \rightarrow \infty \rightarrow x^+ \rightarrow \infty, x^- \rightarrow 0$$

Pate, Raclariu, Strominger, PRL (2017)

Ball, Pate, Raclariu, Strominger, RV, Ann. Phys. 407 (2019) 15

The Wilson lines $U = P \exp \left(i \int_y^\infty dy' \frac{\rho(x_t, y')}{\nabla_t^2} \right)$ are vertex operators on the celestial sphere

Satisfy a 2-D conformal Kac-Moody algebra

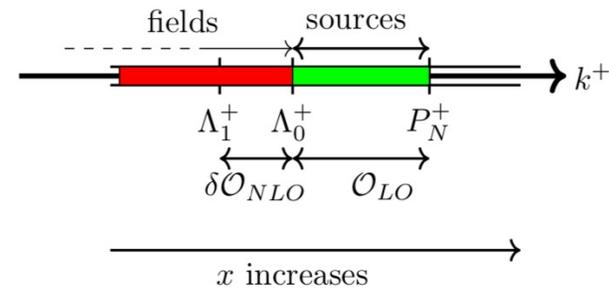
V.P. Nair (1988)
He, Mitra, Strominger (2015)

Boosting memory in CGC EFT: RG hierarchy of many-body correlators

$$\frac{\partial}{\partial Y} \langle \mathcal{O}[\rho] \rangle_Y = \frac{1}{2} \left\langle \int_{x,y} \frac{\partial}{\partial \rho^a(x)} \chi_{x,y}^{ab} \frac{\partial}{\partial \rho^b(y)} \mathcal{O}[\rho] \right\rangle_Y$$

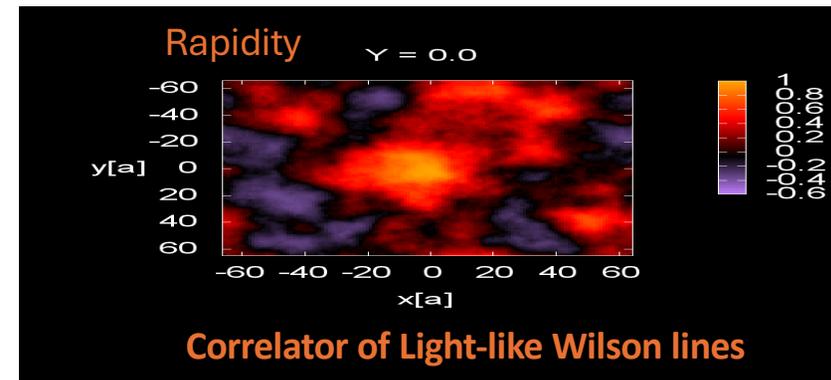
↓
Rapidity → “time”

↓
“diffusion coefficient”: retarded Green function
in strong field background



Langevin diffusion of “wee” partons
in functional space of color fields

B-JIMWLK hierarchy of n-point Wilson line correlators



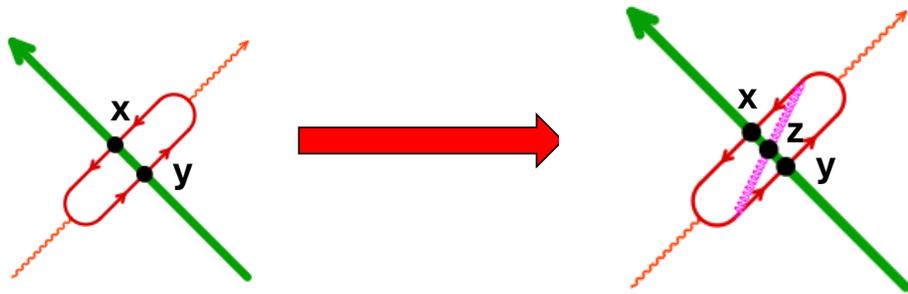
Dumitru, Jalilian-Marian, Lappi, Schenke, RV
PLB706 (2011)219

Balitsky, hep-ph/9509348

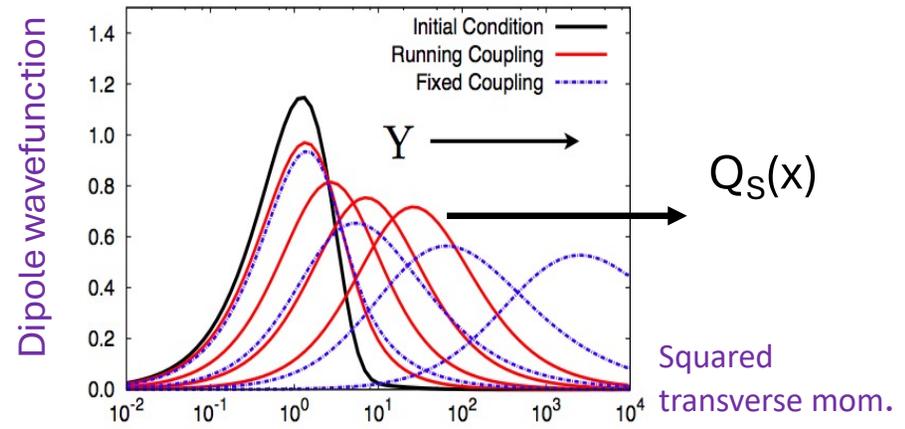
Jalilian-Marian, Kovner, Leonidov, Weigert, hep-ph/9706377

Iancu, Leonidov, McLerran, hep-ph/0011241

How memory evolves in inclusive DIS



Evolution of dipole correlators of Wilson lines in shockwave background

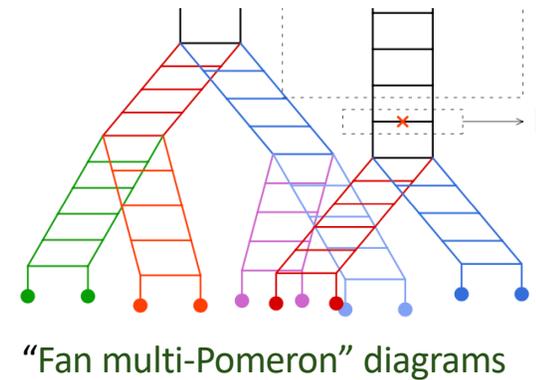


$$\frac{\partial}{\partial Y} \langle \text{Tr}(V_x V_y^\dagger) \rangle_Y = -\frac{\alpha_S N_c}{2\pi^2} \int_{z_\perp} \frac{(x_\perp - y_\perp)^2}{(x_\perp - z_\perp)^2 (z_\perp - y_\perp)^2} \langle \text{Tr}(V_x V_y^\dagger) - \frac{1}{N_c} \text{Tr}(V_x V_z^\dagger) \text{Tr}(V_z V_y^\dagger) \rangle_Y$$

$Y = \text{Ln}(1/x)$

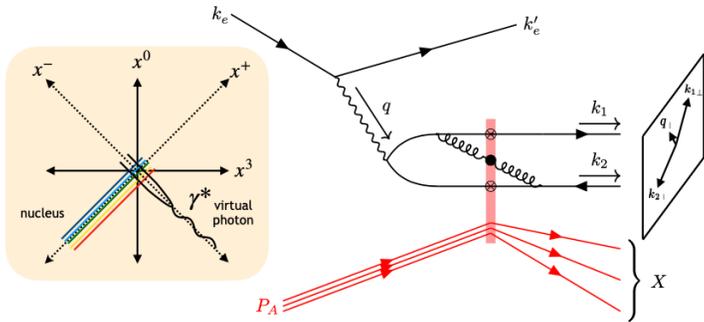
Closed form expression for $A \gg 1, N_c \rightarrow \infty$: non-linear Balitsky-Kovchegov (BK) eqn.

BFKL equation for dipoles is the “low density” linearized form of this equation!



“Fan multi-Pomeron” diagrams

Precision measurements of memory at the EIC

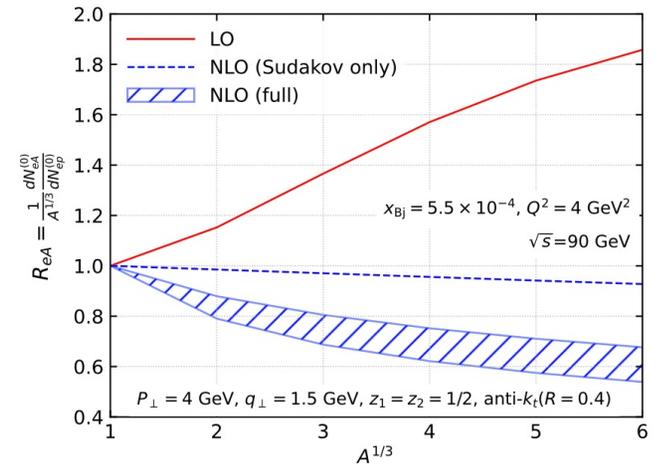
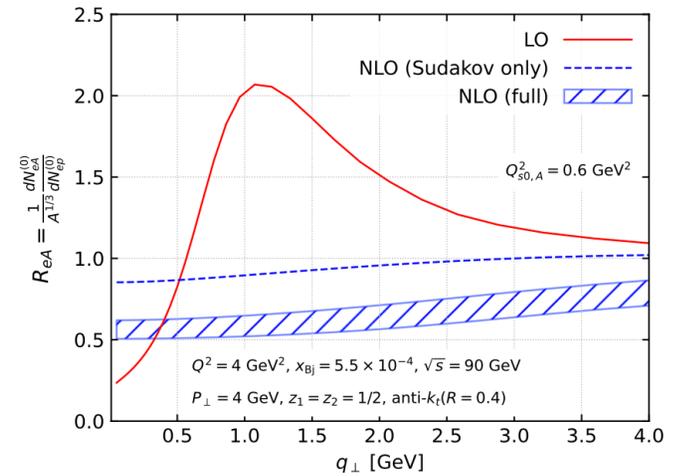


Measure imbalance of back-to-back di-jets in DIS

Caucal, Salazar, Schenke, Stebel, RV, PRL 132 (2024) 8, 081902

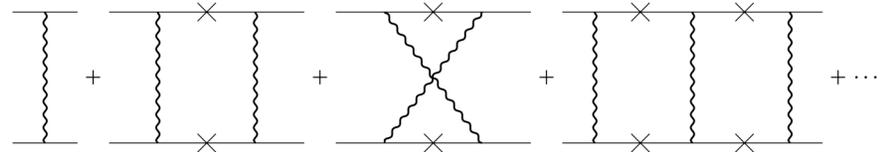
Global analyses to extract “universal” features from p+A collisions at the LHC and e+A collisions from the EIC

Large # of inclusive, semi-inclusive, exclusive and diffractive final states

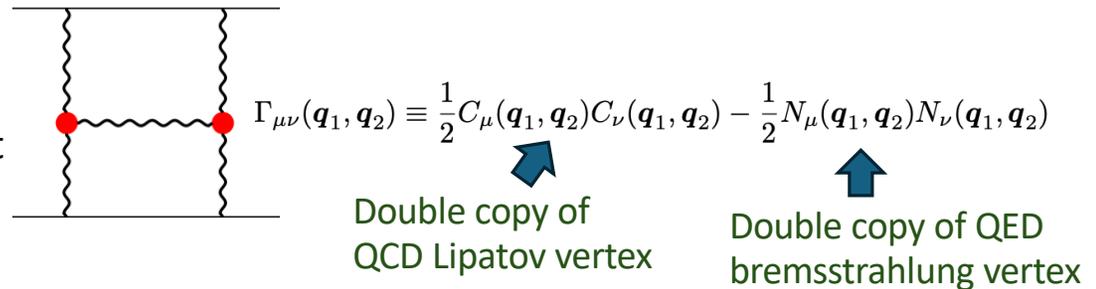


Multiparticle production in gravity: amplitudes to shockwave collisions

A) In gravity, the dominant contribution at large impact parameters is Eikonal scattering



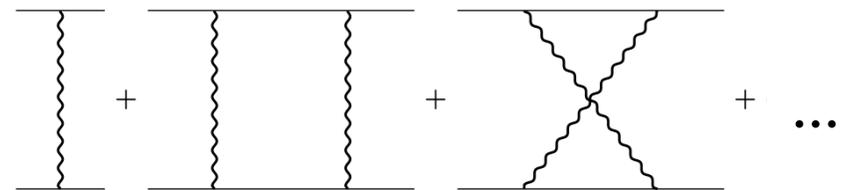
B) Gravitational Lipatov vertex is dominant radiative correction in Regge kinematics – smoothly matches to Weinberg in the ultrarelativistic soft limit



Lipatov, PLB 116B (1982); JETP 82 (1982)

C) Reggeization contributions suppressed except for $b \rightarrow R_s$

$$\mathcal{M}^{(1)} \sim \frac{\kappa^2}{8\pi^2} \left(\underbrace{-i\pi s \log\left(\frac{-t}{\Lambda^2}\right)}_{\text{Eikonal phase}} + \underbrace{t \log\left(\frac{s}{-t}\right) \log\left(\frac{-t}{\Lambda^2}\right)}_{\text{graviton Regge trajectory}} \right)$$



Resummation a la BFKL – but additional ladder and non-ladder $\ln^2(s)$ contributions

Bartels, Lipatov, Sabio-Vera, arXiv:1208.3423
Melville, Naculich, Schnitzer, White, arXiv:1306.6019

Lipatov vertex from shockwave collisions

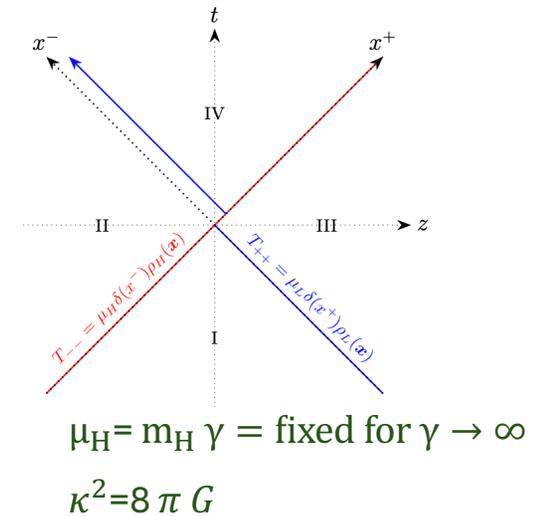
We will now sketch how the Lipatov vertex is recovered in shockwave collisions

Aichelburg-Sexl shockwave metric

$$ds^2 = 2dx^+ dx^- - \delta_{ij} dx^i dx^j + f(x^-, \mathbf{x}) (dx^-)^2$$

$$\text{with } f(x^-, \mathbf{x}) = 2\kappa^2 \mu_H \delta(x^-) \frac{\rho_H(\mathbf{x})}{\square_\perp} = \frac{\kappa^2}{\pi} \mu_H \delta(x^-) \int d^2 \mathbf{y} \ln \Lambda |\mathbf{x} - \mathbf{y}| \rho_H(\mathbf{y})$$

Soln of Einstein's eqns sourced by the EM tensor $T_{\mu\nu} = \delta_{\mu-} \delta_{\nu-} - \mu_H \delta(x^-) \rho_H(\mathbf{x})$



Shockwave collisions: single shock background

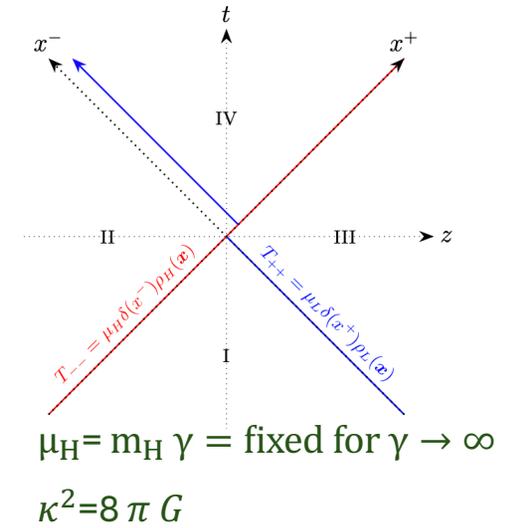
Linearizing around the metric $g_{\mu\nu} = \bar{g}_{\mu\nu} + \kappa h_{\mu\nu}$

fixing light cone gauge $h_{\mu+}=0$, find

$$h_{ij}(x^+, x^-, \mathbf{x}) = V(x^-, \mathbf{x}) h_{ij}(x^+, x^- = x_0^-, \mathbf{x})$$

with the gravitational Wilson line $V(x^-, \mathbf{x}) \equiv \exp\left(\frac{1}{2} \int_{x_0^-}^{x^-} dz^- \bar{g}_{--}(z^-, \mathbf{x}) \partial_+\right)$

Exactly analogous to the QCD case
with $A_- \rightarrow g_{--}$ and $T^a \rightarrow \partial_+$



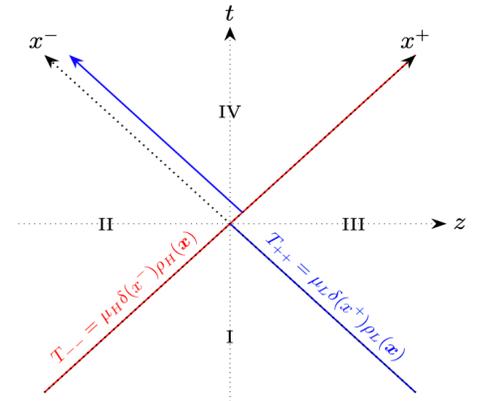
Shockwave collisions: “dilute-dense” approximation

Now consider the interaction of the “dilute” source ρ_L with the dense ρ_H shockwave:

$$T_{\mu\nu} = \delta_{\mu-}\delta_{\nu-}\mu_H\delta(x^-)\rho_H(\mathbf{x}) + \delta_{\mu+}\delta_{\nu+}\mu_L\delta(x^+)\rho_L(\mathbf{x})$$

Solve for metric in region IV – forward lightcone

$$g_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu} \quad \bar{g}_{--} = 2\kappa\mu_H\delta(x^-)\frac{\rho_H(\mathbf{x})}{\square_{\perp}}$$



$$\mu_H = m_H \gamma = \text{fixed for } \gamma \rightarrow \infty$$

$$\kappa^2 = 8\pi G$$

We decompose the perturbation $h_{\mu\nu}$ into a term linear in ρ_L and one bi-linear in $\rho_L\rho_H$ (dilute-dilute limit)

Linearized Einstein’s equations in light-cone gauge ($h_{+\mu}=0$) take the form

$$\bar{g}_{--}\partial_+^2\tilde{h}_{ij} - \square_{\perp}\tilde{h}_{ij} = \kappa^2 \left[\left(2\partial_i\partial_j - \square_{\perp}\delta_{ij}\right)\frac{1}{\partial_+^2}T_{++} + 2T_{ij} - \delta_{ij}T - \frac{2}{\partial_+} \left(\partial_iT_{+j} + \partial_jT_{+i} - \delta_{ij}\partial_kT_{+k}\right) \right]$$

$$\tilde{h}_{ij} \equiv h_{ij} - \frac{1}{2}\delta_{ij}h \text{ where } h = \delta_{ij}h_{ij}$$

Shockwave collisions in general relativity: geodesics

Unlike QCD case, sub-eikonal contributions T_{+i}, T_{ij} are required for consistency of equations of motion

These are not uniquely fixed by energy-momentum conservation, the dynamics of the sources is needed to fix this. In the point particle approximation,

$$T^{\mu\nu}(x) = \frac{\mu_L}{\sqrt{-g}} \int_{-\infty}^{\infty} d\lambda \dot{X}^\mu \dot{X}^\nu \delta^{(4)}(x - X(\lambda))$$

Solution of the corresponding null geodesic equations $\ddot{X}^\mu + \Gamma_{\nu\rho}^\mu \dot{X}^\nu \dot{X}^\rho = 0$, $g_{\nu\rho} \dot{X}^\nu \dot{X}^\rho = 0$

in shockwave background given by $X^- = \lambda$, $X^i = b^i - \kappa^2 \mu_H X^- \Theta(X^-) \frac{\partial_i \rho_H(\mathbf{b})}{\square_\perp}$

$$X^+ = -\kappa^2 \mu_H \Theta(X^-) \frac{\rho_H(\mathbf{b})}{\square_\perp} + \frac{\kappa^4 \mu_H^2}{2} X^- \Theta(X^-) \left(\frac{\partial_i \rho_H(\mathbf{b})}{\square_\perp} \right)^2$$

These geodesic solutions allow us to reconstruct the required components of the stress-energy tensor

Shockwave collisions in general relativity: Lipatov vertex

Solving eqns of motion, taking the Fourier transform, and putting the graviton momenta on-shell, one obtains

Gravitational
radiational field

$$\tilde{h}_{ij}^{(2)}(k) = \frac{2\kappa^3 \mu_H \mu_L}{k^2 + i\epsilon k^-} \int \frac{d^2 \mathbf{q}_2}{(2\pi)^2} \Gamma_{ij}(\mathbf{q}_1, \mathbf{q}_2) \frac{\rho_H}{\mathbf{q}_1^2} \frac{\rho_L}{\mathbf{q}_2^2}$$

Gravitational Lipatov vertex



likewise for other components, recovering $\Gamma_{\mu\nu}(\mathbf{q}_1, \mathbf{q}_2) \equiv \frac{1}{2} C_\mu(\mathbf{q}_1, \mathbf{q}_2) C_\nu(\mathbf{q}_1, \mathbf{q}_2) - \frac{1}{2} N_\mu(\mathbf{q}_1, \mathbf{q}_2) N_\nu(\mathbf{q}_1, \mathbf{q}_2)$

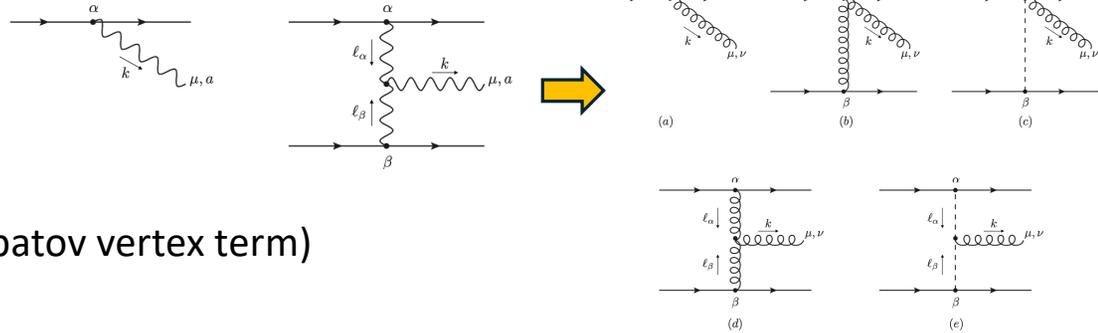
Compare to gauge theory
radiation field

$$a_i(k) = \frac{g^3}{k^2 + i\epsilon k^-} \int \frac{d^2 \mathbf{q}_2}{(2\pi)^2} C_i(\mathbf{q}_1, \mathbf{q}_2) \frac{\rho_H \cdot T}{\mathbf{q}_1^2} \frac{\rho_L}{\mathbf{q}_2^2}$$

$$-if^{abc} T_b T_c C_\mu(\mathbf{q}_1, \mathbf{q}_2) \quad \xleftrightarrow[\text{CK relation?}]{\text{Is there a}} \quad s\Gamma_{\mu\nu}(\mathbf{q}_1, \mathbf{q}_2)$$

Lipatov vertex from classical color-kinematic duality

Consider pert. solutions of Yang-Mills radiation field in collision of colored charges c_α^a (Wong equations)



Taking ultrarelativistic limit (keeping sub-eikonal terms, beyond leading QCD Lipatov vertex term) and making replacements

$$c_\alpha^a \rightarrow p_\alpha^\mu$$

$$if^{a_1 a_2 a_3} \rightarrow \Gamma^{\nu_1 \nu_2 \nu_3}(q_1, q_2, q_3) = -\frac{1}{2}(\eta^{\nu_1 \nu_3}(q_1 - q_3)^{\nu_2} + \eta^{\nu_1 \nu_2}(q_2 - q_1)^{\nu_3} + \eta^{\nu_2 \nu_3}(q_3 - q_2)^{\nu_1})$$

$$g \rightarrow \kappa$$

recovers our previous result for the radiation field in terms of the gravitational Lipatov vertex

Raj,RV, arXiv:2312.03507

We can also show that the soft limit of the gravitational Lipatov vertex gives the ultrarelativistic limit of the Weinberg soft graviton emission vertex

This is consistent with the observation that the soft limit of the classical double copy recovers the Weinberg emission vertex

P.V. Athira,A. Manu, arXiv:1907.10021

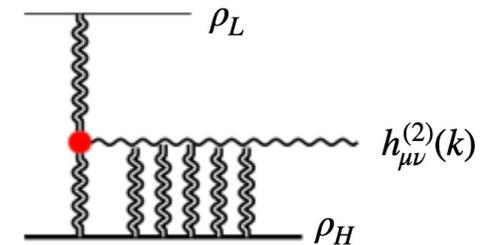
Goldberger, Ridgway, arXiv:1611.03493

RG description in gravity a la CGC EFT in QCD?

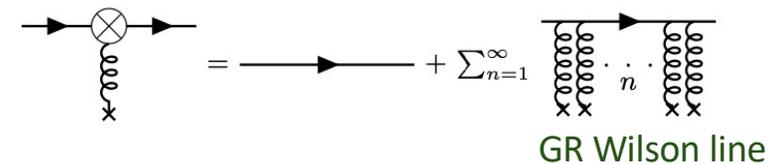
In QCD, the CGC describes a semi-classical lump of size $1/Q_s$ that perturbatively unitarizes cross-sections. It corresponds to the non-trivial IR fixed point of an RG flow where the UV fixed point is BFKL

Can one obtain a similar RG picture in gravity in the shockwave framework? Reggeization must of single ladders must break down at low impact parameter ($t \sim s$).

Instead, multiple scattering in the “dilute-dense” framework plays a central role In a systematic treatment in this framework, we see that geodesic focusing (a la Raychaudhuri) of wee partons must play a role



Another key element in the RG description are shockwave propagators - Again in GR, it may be important to keep non-eikonal contributions



Raj, RV, arXiv:2406.10483

Qualitative ideas of a CGC- Black Hole correspondence as overoccupied gluon/graviton states have been discussed by myself and Gia Dvali

Dvali, RV, arXiv:11989



Geodesic congruence: the geometry of quantum information

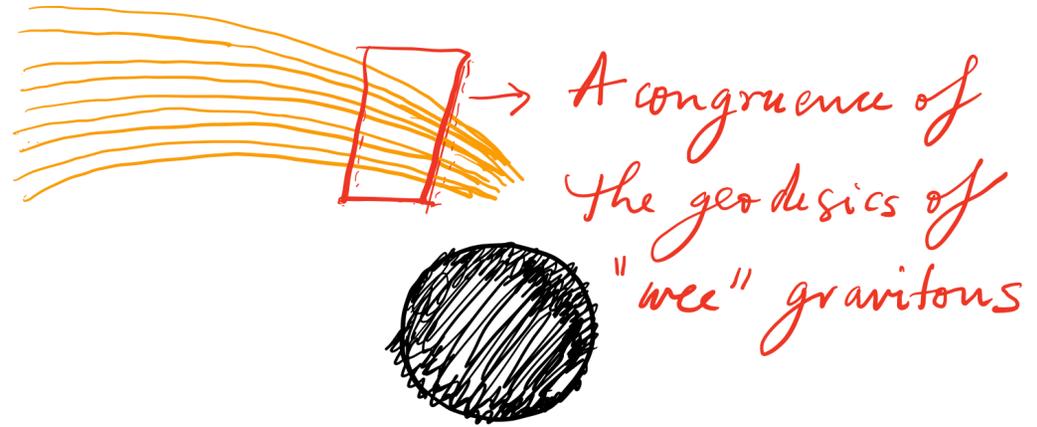
The Raychaudhuri equation

- key in Hawking-Penrose singularity theorems :

Volume change of geodesic congruence

$$\begin{aligned} \dot{\theta} &= -\Omega^i_j \Omega^j_i + K^i_i \\ &= -\frac{1}{3}\theta^2 - \sigma_{ij}\sigma^{ij} + \omega_{ij}\omega^{ij} + K^i_i. \end{aligned}$$

Bulk scalar Shear tensor Rotation tensor Includes Ricci curvature + stochastic graviton noise



H.-T. Cho and B.-L Hu, arxiv:2301.06325
M. Parikh, F. Wilczek, G. Zaharade, PRL (2021)

Remarkably, the Raychaudhuri equation can be rephrased as the Bishop-Gromov upper bound on the “complexity volume in D-1 dimensions” of gate complexity

– realizing Nielsen’s geometric picture in quantum information theory?

A. R. Brown, arXiv:2112.05724

Pure speculation: Can this complexity picture provide further insight into the RG fixed point of BH formation?

2 → N + 2 amplitudes in trans-Planckian gravitation scattering: from wee partons to Black Holes

HIGH-ENERGY SCATTERING IN QCD AND IN QUANTUM GRAVITY AND TWO-DIMENSIONAL FIELD THEORIES

L.N. LIPATOV*

We construct effective actions describing high-energy processes in QCD and in quantum gravity with intermediate particles (gluons and gravitons) having the multi-Regge kinematics. The S -matrix for these effective scalar field models contains the results of the leading logarithmic approximation and is unitary. It can be expressed in terms of correlation functions for two field theories acting in longitudinal and transverse two-dimensional subspaces.

Effective action and all-order gravitational eikonal at planckian energies

AMATI, CIAFALONI, VENEZIANO **NPB403 (1993)707**

Building on previous work by us and by Lipatov, we present an effective action approach to the resummation of all semiclassical (i.e. $O(\hbar^{-1})$) contributions to the scattering phase arising in high-energy gravitational collisions. By using an infrared-safe expression for Lipatov's effective action, we derive an eikonal form of the scattering matrix and check that the superstring amplitude result is reproduced at first order in the expansion parameter R^2/b^2 , where R , b are the gravitational radius and the impact parameter, respectively. If rescattering of produced gravitons is neglected, the longitudinal coordinate dependence can be explicitly factored out and exhibits the characteristics of a shock-wave metric while the transverse dynamics is described by a reduced two-dimensional effective action. Singular behaviours in the latter, signalling black hole formation, can be looked for.

The World as a Hologram

LEONARD SUSSKIND

Wee partons, by contrast, are not subject to Lorentz contraction. This implies that in the Feynman Bjorken model, the halo of wee partons eternally "floats" above the horizon at a distance of order $10^{-13}cm$ as it transversely spreads. The remaining valence partons carry the various currents which contract onto the horizon as in the Einstein Lorentz case.

By contrast, both the holographic theory and string theory require all partons to be wee. No Lorentz contraction takes place and the entire structure of the string floats on the stretched horizon. I have explained in previous articles how this behavior prevents the accumulation of arbitrarily large quantities of information near the horizon of a black hole. Thus we are led full circle back to Bekenstein's principle that black holes bound the entropy of a region of space to be proportional to its area.

***J.Math.Phys.* 36 (1995) 6377; > 4 K cites !**

In Acknowledgements:

Finally I benefitted from discussions with Kenneth Wilson and Robert Perry, about boosts and renormalization fixed points in light front quantum mechanics and Lev Lipatov about high energy scattering.



**30+ years of work by ACV et al. exploring
gravitational shockwave collisions in this 2-D EFT**