

The Project in a Nutshell

NUclear CLustering Effects in Astrophysical Reactions

NUCLEAR



European Research Council
Established by the European Commission



UK Research
and Innovation

Nucleosynthesis in First Stars and Other Puzzles

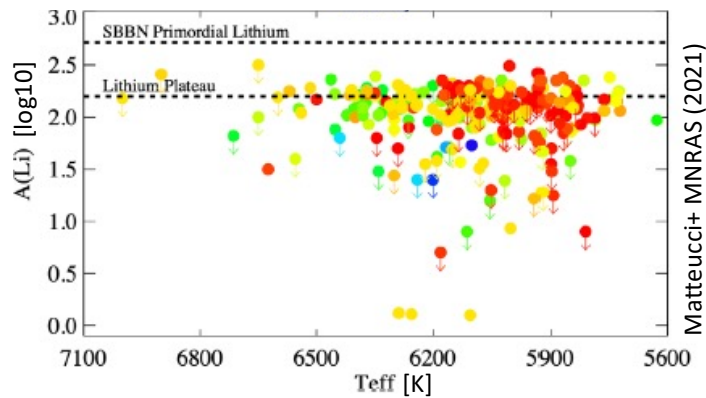


Long-Standing Questions in Nuclear Astrophysics



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Q1. Cosmological Lithium Problem



factor of 3 discrepancy between
observed and predicted Li abundance



Standard Model of Particle Physics
+ Cosmology

Q2. Nucleosynthesis in First Stars

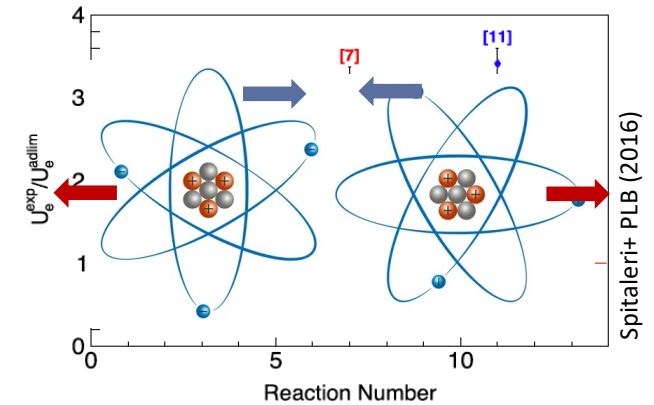


made of pristine H and He
very massive → need CNO nuclei



Chemical Evolution of Early Universe
+ Astronomical Observations (JWST)

Q3. Electron Screening Puzzle



discrepancy between experiment
and theory remains unexplained



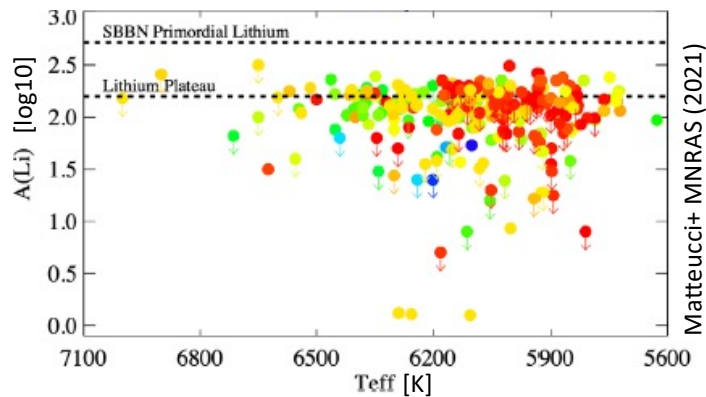
Reactions in Plasmas
Fusion-driven Energy Generation

Long-Standing Questions in Nuclear Astrophysics



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Q1. Cosmological Lithium Problem



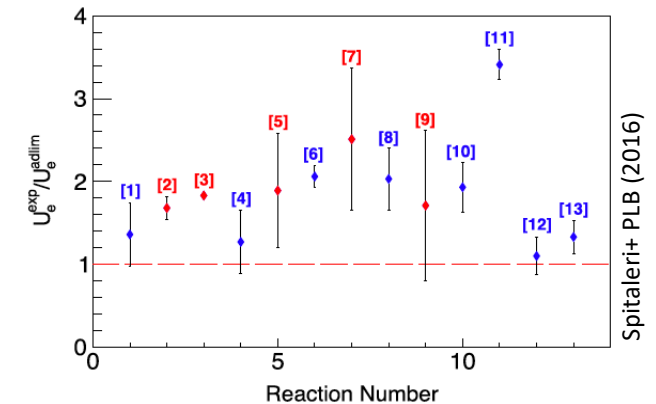
factor of 3 discrepancy between
observed and predicted Li abundance

Q2. Nucleosynthesis in First Stars



made of pristine H and He
very massive → need CNO nuclei

Q3. Electron Screening Puzzle



discrepancy between experiment
and theory remains unexplained

key idea:

Nuclear Clustering



key to unlock all three puzzles



Nuclear Clustering



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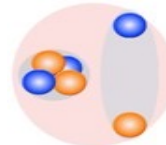
a. He-4 = α particle



p proton
n neutron

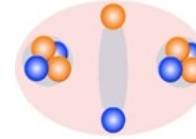
very stable configurations
→ building blocks for other nuclei

${}^6\text{Li}$



$\alpha \oplus d$

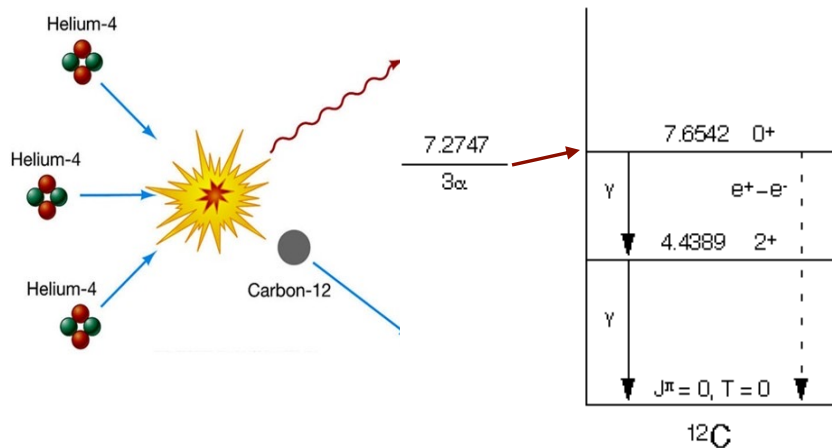
${}^{10}\text{B}$



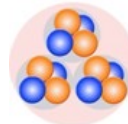
$\alpha \oplus d \oplus \alpha$

nuclear clustering may greatly
enhance fusion probabilities
at low (i.e. astrophysical) energies

triple alpha process



Hoyle state



$\sim 10^7$ times faster

electron screening puzzle



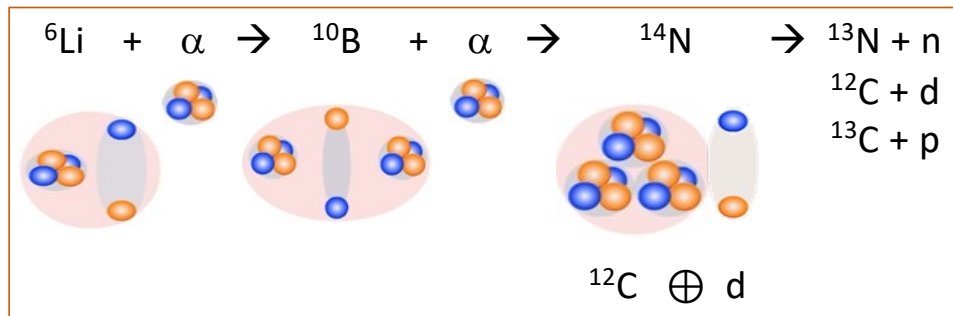
lower Coulomb barrier → enhanced fusion

Idea and Current Status



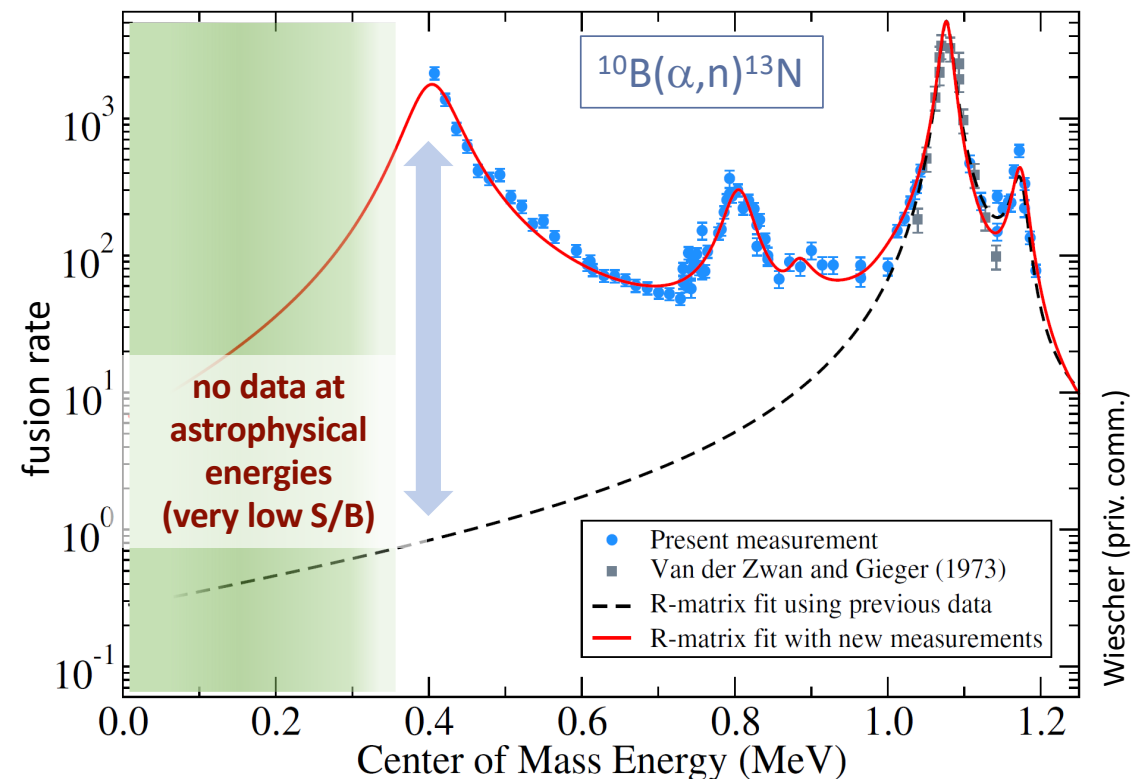
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α -induced reactions on Li and B proposed to produce CNO and destroy Li if enhanced by clustering



hypothesis untested both experimentally and theoretically until recently

tantalizing new evidence of cluster structures enhancing fusion rates in α -induced reactions on Li and B



Work Programme and Outcomes



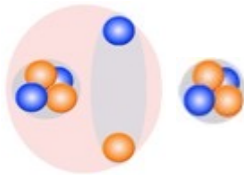
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Theoretical Programme

WP2a: PDRA2



G Hupin



- cluster structures at low energies
- impact on astrophysical reactions & electron screening



RJ deBoer

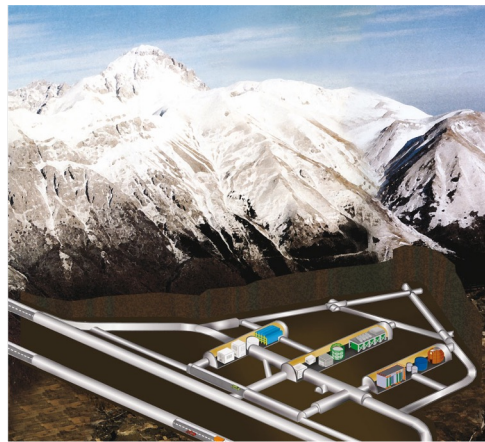
WP2b: **PI**, PDRA2

- improved R-matrix capabilities
- robust reaction rates

Experimental Programme

WP1: **PI**, PDRA1, PhD1, PhD2

Laboratory for Underground Nuclear Astrophysics

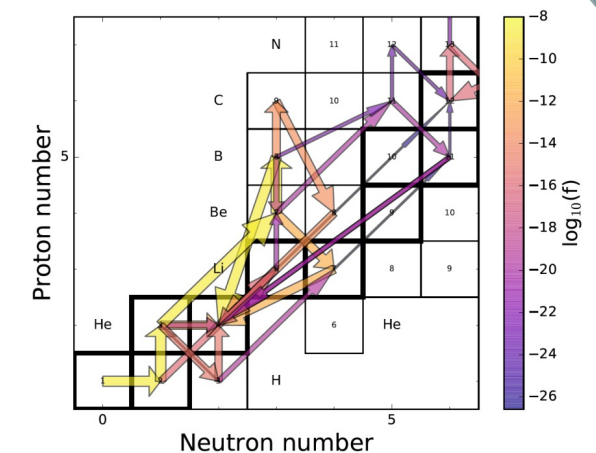


- α +Li and α +B reactions (Q1-Q3)
- ultra-low background @LUNA
- lowest-energy data (world best)

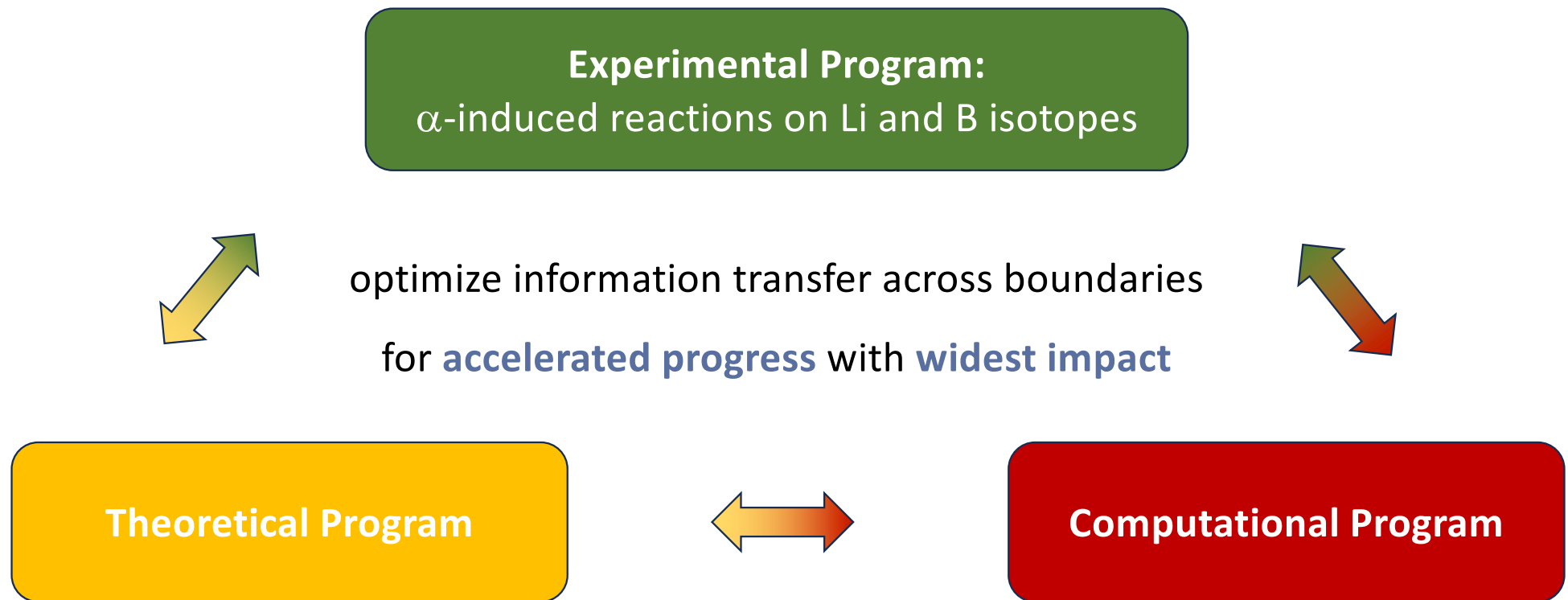
Computational Programme

WP3: **PI**, PDRA3

M Pignatari



- stellar models for first stars (MESA)
- nucleosynthesis networks (NuGRID)
- impact on Q1 and Q2



Timeline



		NUCLEAR: NUClear CLustering Effects in Astrophysical Reactions									
		PI: Marialuisa Aliotta									
Task	Assigned to	Year 1		Year 2		Year 3		Year 4		Year 5	
WP1: Experimental program		S1	S2	S3	S4	S5	S6	S7	S8	S9	S10
T1a	PhD1, PDRA1, PI				M1a						
T1b	PhD2, PDRA1, PI							M1b			
WP2: Theoretical program											
T2a.1	PDRA2, GH				M2a.1						
T2a.2	PDRA2, GH					M2a.2					
T2b.1	PDRA2, JdB										
T2b.2	PDRA2, JdB									M2b	
T2b.3	PDRA2, JdB, PI										M2b
WP3: Computational program											
T3.1	PDRA1, PDRA3, PI					M3.1					
T3.2	PDRA3, MP							M3.2			
T3.3	PDRA3, MP										M3.3
Direction and Oversight											
	PI										

Milestones

M1a: low-energy cross sections data for 10B+a reactions

M1b: low-energy cross sections data for 6,7Li+a radiative captures

M2a.1: theoretical evaluation of reaction cross sections with N isotopes as compound nuclei

M2a.2: complete theoretical evaluation of radiative capture cross sections

M2b: evaluation of uncertainty in R-matrix fits and extrapolation with inclusion of improved potential models and electron screening

M3.1: initial stellar reaction rates evaluations from available and new experimental data (from WP1) as they become available

M3.2: nucleosynthesis calculations in first-generation stars (zero metallicity) with input from from T3.1

M3.3: complete nucleosynthesis simulations for largest progenitor stellar masses

Grant Start Date: 1 December 2024



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<https://www.erc-nuclear.uk>



Meet the Team



Michael Wiescher
(Deus ex Machina)



Marialuisa Aliotta
PRINCIPAL INVESTIGATOR



Guillaume Hupin
TEAM MEMBER



Richard James deBoer
TEAM MEMBER



Marco Pignatari
TEAM MEMBER

Other Team Members

PhD Students



Jamie Jones, PhD
(September 2024)



Rhys Bonnell, PhD
(December 2024)



Lavinia Dalla Vedova
(from September 2025)

Post-Docs



Alessandro Compagnucci
(February 2025)



Kevin Becker
(from July 2025)



LUNA collaborators

Zoltan Elekes
Piero Corvisiero
Jakub Skowronski
Axel Boeltzig
Gianluca Imbriani
Tom Davinson
Ragan Sidhu
Carlo Bruno
Lucia Barbieri
Duncan Robb
Joao Cruz
Gyuri Gyürky
Sandra Zavatarelli, ...

Notre Dame team

Khachatur Manukyan
local students & PhD...



ERC-NUCLEAR Kick-off Event - 15 May 2025			
09:30	09:45	welcome and introductions	Marialuisa (MLA)
09:45	10:00	ERC Overview	MLA
10:00	10:40	Scientific Goals	Michael + MLA
10:40	11:00	discussion	
11:00	11:30	coffee break	
		Experimental Programme	
11:30	12:00	10B(a,p) and 10B(a,d)	Jamie
12:00	12:30	10B(a,n)	Rhys
12:30	13:00	6,7Li(a,g)	Alessandro
13:00	14:00	Lunch & Photo	
		Computational Programme	
14:00	14:40	First Stars: Stellar Models and Nucleosynthesis	Marco
		Theoretical Programme	
14:50	15:30	Reactions with clusters	Guillaume
15:30	16:00	Coffee Break	
16:00	16:40	R-matrix developments	James
16:40	17:00	Closing Remarks and Next Steps	
19:30		Dinner at Nok's Kitchen by the Castle	