

# THE MANY FACES OF (NEXT-TO) SOFT PHYSICS

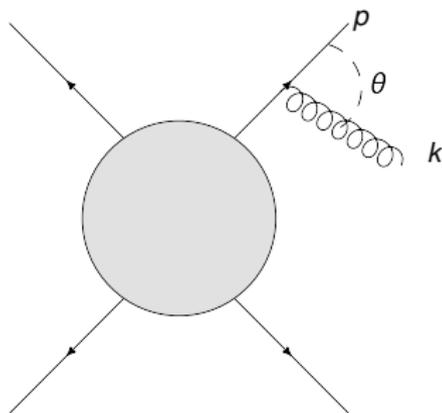
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Symmetries of the S-matrix and Infrared Physics

- Brief introduction to (next-to-) soft divergences.
- Applications in Collider Physics (mainly QCD).
- Applications in high energy scattering (mainly gravity).
- Outlook.

# INFRARED DIVERGENCES

- In scattering amplitudes, get singularities due to soft or collinear gauge bosons:



$$\frac{1}{p \cdot k} = \frac{1}{|\mathbf{p}||\mathbf{k}|\left(1 - \cos \theta\right)}.$$

- Formal divergences cancel upon combining real and virtual graphs ([Block](#), [Nordsieck](#)).
- Both soft and collinear radiation is *universal*.
- Physics: it has an infinite wavelength, so cannot resolve the underlying amplitude.

- Universality of soft / collinear radiation is expressed in *factorisation formulae*.
- Example: consider a tree-level amplitude  $\mathcal{A}_{n+1}(\{p_i\}, k)$  where momentum  $k$  becomes soft. We then get the *soft theorems*

$$\lim_{k^\mu \rightarrow 0} \mathcal{A}_{n+1}(\{p_i\}, k) = \mathcal{S}^{(0)}(\{p_i\}, k) \mathcal{A}_n(\{p_i\}),$$

where

$$\mathcal{S}_{\text{QED}}^{(0)} = \sum_{i=1}^n \frac{\epsilon_\mu(k) p_i^\mu}{p_i \cdot k}, \quad \mathcal{S}_{\text{grav.}}^{(0)} = \sum_{i=1}^n \frac{\epsilon_{\mu\nu}(k) p_i^\mu p_i^\nu}{p_i \cdot k}$$

(Yennie, Frautschi, Suura; Weinberg).

- All dependence on the soft momentum  $k$  is in the overall factor  $\mathcal{S}$ .

- It is also possible to write such formulae at one order higher in the  $k$  expansion (Cachazo, Strominger; Casali):

$$\mathcal{A}_{n+1}(\{p_i\}, k) = \left[ \mathcal{S}^{(0)} + \mathcal{S}^{(1)} \right] \mathcal{A}_n(\{p_i\}),$$

with

$$\mathcal{S}_{QED}^{(1)} = \sum_{i=1}^n \frac{\epsilon_\mu k_\rho J^{(i)\mu\rho}}{p_i \cdot k}, \quad \mathcal{S}_{grav.}^{(1)} = \sum_{i=1}^n \frac{\epsilon_\mu k_\rho J^{(i)\mu\rho}}{p_i \cdot k},$$

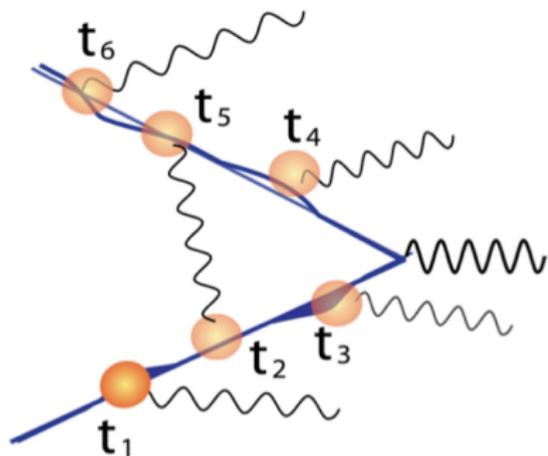
where  $J_{\mu\nu}^{(i)}$  is the total angular momentum of (hard) particle  $i$ .

- Next-to-next-to-soft also possible for gravity.
- These and similar results have a surprisingly long history...

- Next-to-soft effects were first studied in gauge theory (QED) by [Low](#) (1958).
- He considered external scalars; generalised to fermions by [Burnett](#) and [Kroll](#) (1968).
- Both groups only considered massive particles (no collinear effects).
- Similar work in gravity by [Gross, Jackiw](#) (1968).
- [Del Duca](#) (1990) generalised the Low-Burnett-Kroll result to include collinear effects.

# PATH INTEGRAL APPROACH

- Next-to-soft effects for massive particles considered using worldline methods by [Laenen, Stavenga, White \(2008\)](#).



- Also works for gravity ([White, 2011](#)).

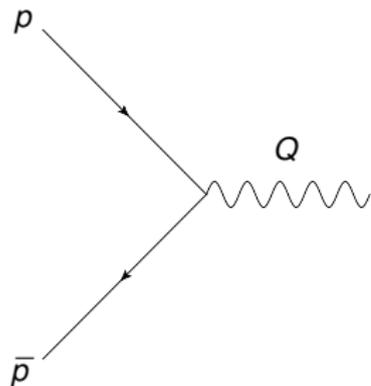
- Can replace propagators for external legs by quantum mechanics path integrals.
- Leading term in perturbative expansion is classical trajectory (soft limit).
- First-order wobbles give next-to-soft behaviour.

- The tree-level (next-to)-soft theorems can be obtained using Ward identities associated with asymptotic symmetries.
- This is the focus of much of this meeting!
- However, the history of next-to-soft physics suggests that there are many other applications of next-to-soft physics.
- Indeed, these have been reinvigorated by the recent work on next-to-soft theorems.
- The aim of this talk is to review some of these applications.

Key message: next-to-soft physics connects hep-th, hep-ph, hep-ex and gr-qc!

# COLLIDER PHYSICS

- A major application of (next-to) soft physics is to collider physics.
- We saw earlier that IR singularities cancel when real and virtual diagrams are combined.
- However, the cancellation can leave behind large contributions to perturbative quantities.
- Consider e.g. the production of a vector boson at a collider (“Drell-Yan production”):



- Let  $z = Q^2/s$  be the fraction of (squared) energy  $s$  carried by the vector boson.
- At LO,  $z = 1$ , and thus the cross-section is

$$\frac{d\sigma^{(0)}}{dz} \propto \delta(1 - z).$$

- At next-to-leading order (NLO), radiation can carry energy, so that

$$0 \leq z \leq 1.$$

- The NLO cross-section then turns out to be

$$\frac{d\sigma_{q\bar{q}}^{(1)}}{dz} \sim \frac{\alpha_s}{2\pi} \left[ 4(1+z^2) \left( \frac{\ln(1-z)}{1-z} \right)_+ - 2 \frac{1+z^2}{1-z} \ln(z) + \delta(1-z) \left( \frac{2\pi^2}{3} - 8 \right) \right].$$

- It contains highly divergent terms as  $z \rightarrow 1$ .
- Looks like perturbation theory is in trouble!
- Let's go one order higher and see what happens...

- At NNLO the problem is even worse! One has

$$\frac{d\sigma_{q\bar{q}}^{(2)}}{dz} \sim C_F^2 \left(\frac{\alpha_s}{2\pi}\right)^2 \left[ 128 \left(\frac{\ln^3(1-z)}{1-z}\right)_+ - 256 \left(\frac{\ln(1-z)}{1-z}\right)_+ + \dots \right],$$

where ... denotes terms suppressed by  $(1-z)$ .

- Logs get higher at higher orders in perturbation theory...
- ... which indeed breaks down as  $z \rightarrow 1$ .
- Precisely the regime where the vector boson is produced near threshold, so that extra radiation is soft / collinear!
- The problem terms are echoes of IR singularities having been present.
- Thus, this problem affects many different scattering processes...

- For heavy particles produced near threshold, we can define a  $\xi$ , where  $\xi \rightarrow 0$  at threshold (e.g.  $\xi = (1 - z)$ ).
- Then the general structure of any such cross-section is:

$$\frac{d\sigma}{d\xi} = \sum_{n,m} \alpha^n \left[ c_{nm}^{(0)} \left( \frac{\ln^m \xi}{\xi} \right)_+ + c_{nm}^{(1)} \ln^m \xi + \dots \right].$$

- First set of terms correspond to (leading) threshold logs: pure soft and / or collinear.
- Second set of terms is next-to-leading power (NLP) threshold logs: next-to-soft and / or collinear.
- For  $\xi \rightarrow 0$ , we need to rethink perturbation theory.

- The solution to this problem is to somehow work out what the large logs are to all orders in  $\alpha_s$ .
- Then we can sum them up to get a function of  $\alpha_s$  that is better behaved than any fixed order perturbation expansion.
- Toy example: consider the function

$$e^{-\alpha_s x} = \sum_{n=0}^{\infty} \frac{\alpha_s^n (-x)^n}{n!}.$$

- Each term diverges as  $x \rightarrow \infty$ , but the all-order result is well-behaved.

- Many approaches exist for resumming leading threshold logs.
- There are many (hundreds?) of observables at e.g. the LHC for which this is relevant.
- Original diagrammatic approaches by e.g. [Sterman](#); [Catani, Trentadue](#)),
- Can also use Wilson lines ([Korchemsky, Marchesini](#)), or the renormalisation group ([Forte, Ridolfi](#)).
- A widely used approach is to treat soft and collinear gluons as separate fields in an effective theory: soft-collinear effective theory (SCET) ([Becher, Neubert](#); [Schwartz](#); [Stewart](#)).
- All approaches have the *factorisation* of soft / collinear physics at their heart.

- The general structure of an  $n$ -point amplitude is

$$\mathcal{A}_n = \mathcal{H}_n \times \mathcal{S} \times \frac{\prod_i J_i}{\prod_i \mathcal{J}_i}.$$

- This is the virtual generalisation of the soft theorem.
- Here  $\mathcal{H}_n$  is the *hard function*, and is IR finite.
- The *soft* and *jet functions*  $\mathcal{S}$  and  $J_i$  collect soft / collinear singularities respectively.
- The *eikonal jets*  $\mathcal{J}_i$  remove any double counting.
- The soft and jet functions have universal definitions in terms of Wilson line operators.

- The soft-collinear factorisation formula leads directly to resummation of threshold effects.
- Related ideas in other approaches (e.g. SCET).
- Summing successive towers of threshold logs requires calculating the soft and jet functions to a given order in perturbation theory.
- State of the art is two loops (Sterman, Aybat, Dixon, Kidonakis, Mitov, Sung, Becher, Neubert, Beneke, Falgari, Schwinn, Ferroglia, Pecjak, Yang).
- Progress towards three-loops and beyond (Gardi, Laenen, Stavenga, Smillie, White, Almelid, Duhr, Korchemsky, Henn, Huber, Grozin, Marquard, Correa, Maldacena, Sever).

- To date, much less has been known about NLP effects.
- Known for a while to be numerically significant e.g. in Higgs production ([Kramer Laenen, Spira](#); [Harlander, Kilgore](#); [Catani, de Florian, Grazzini, Nason](#)).
- This has been confirmed by recent N<sup>3</sup>LO Higgs results ([Anastasiou, Duhr, Dulat, Herzog, Mistlberger](#)).
- There are three good reasons to study NLP logs:
  - ① Resummation of them will improve precision.
  - ② Even without resummation, NLP logs may provide good approximate N<sup>n</sup>LO cross-sections.
  - ③ Can improve the stability of numerical codes.

- Next-to-soft effects in particular scattering processes classified to all orders by (Almasy, Moch, Presti, Soar, Vermaseren, Vogt).
- Can also be classified using the *method of regions* (Beneke, Smirnov, Pak, Jantzen) (see e.g. Bonocore, Laenen, Magnea, Vernazza, White).
- None of the previous approaches is fully general - but strong hints of an underlying structure.
- Can we predict NLP logs in an arbitrary process?
- Can they be written in terms of universal functions (like LP effects)?
- Encouraging recent progress...

- It is well-known that LP effects can be described using *Soft-Collinear Effective Theory* SCET (Stewart, Schwartz, Bauer, Fleming; Becher, Neubert).
- The same language can be extended to NLP level.
- Originally explored in B physics (Beneke, Campanario, Mannel, Pecjak).
- Recent study for scattering amplitudes (Larkoski, Neill, Stewart).
- Phenomenology explored by Feige, Kolodrubetz, Moulst, Stewart, Rothen, Tackmann, Zhu; Boughezal, Liu, Petriello.
- Recent resummation of leading NLP log for some observables (Stewart et. al.).

- The soft-collinear factorisation formula can be generalised to next-to-leading power level (Bonocore, Laenen, Magnea, Melville, Vernazza, White).
- This provides a loop-level generalisation of the next-to-soft theorem.
- A new quantity appears at next-to-soft level: the *jet emission function*.
- Has been calculated at one-loop level for quarks.
- Non-trivial check: reproduces all NLP terms up to NNLO in Drell-Yan.
- Observable loop-level corrections to the tree-level next-to-soft theorem!

- Next-to-soft physics has a large number of applications in collider physics.
- Typically this involves summing up large terms in perturbative cross-sections...
- ... or finding approximate forms for fixed-order cross-sections.
- Such calculations improve the precision of theory predictions at the LHC.
- Current data demands this precision!

- Much of this conference focuses on relating (next-to) soft physics with asymptotic symmetries in gravity.
- However, (next)-to soft corrections have a different role to play in understanding the conceptual structure of quantum gravity...
- ...and may even have phenomenological consequences!
- More specifically, they are relevant to high energy scattering.
- Many papers from the 1990s onwards ([Amati](#), [Ciafaloni](#), [Veneziano](#), [Colferai](#), [Falcioni](#); 't Hooft; [Verlinde](#)<sup>2</sup>; [Jackiw](#), [Kabat](#), [Ortiz](#)).

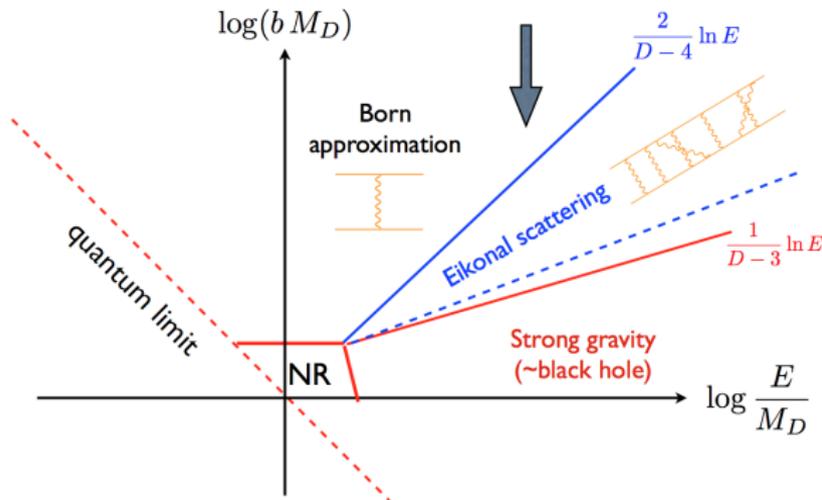
- More specifically, we will focus on  $2 \rightarrow 2$  scattering in the high energy or *Regge* limit

$$s \gg |t|,$$

where  $s$  is the squared centre of mass energy, and  $|t|$  the momentum transfer.

- Corresponds to scattering above the Planck scale in gravity.
- Naïvely, we might think that non-renormalisability is a problem.
- However, in this limit infinite numbers of *soft* gravitons are exchanged, and the results are well-behaved!

- Can consider different regions in impact parameter  $b$  (conjugate to  $|t|$ ), and energy  $E \sim \sqrt{s}$ :



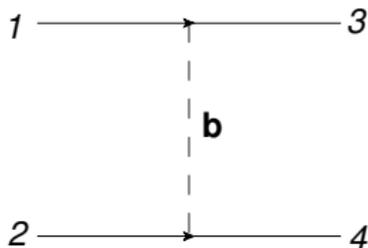
(see e.g. [Giddings, Schmidt-Sommerfeld, Andersen](#)).

- Next-to-soft corrections probe unknown parts of this diagram.

- Previous work on this topic focused on gravity only, including possible string theory corrections.
- Recent studies have used QCD methods to analyse gravity scattering ([Akhoury, Saotome, Sterman](#); [Melville, Naculich, Schnitzer, White](#)).
- Idea is to develop a common language, that makes the structure of both theories clear.
- Let us look first at QCD...

# WILSON LINES AND THE REGGE LIMIT

- The Regge limit can be described by two Wilson lines separated by a transverse distance (Korchemsky, Korchemskaya).
- See also Balitsky; Caron-Huot.



- Take particles of mass  $m$ , such that

$$s \gg -t \gg m^2.$$

- $\mathbf{b}$  is the (2-d) impact parameter (distance of closest approach).

- In the asymptotic high energy limit, the incoming / outgoing particles follow classical straight line trajectories i.e. they do not recoil.
- The only quantum behaviour they are allowed is to experience a phase change.
- However, gauge-covariance of the amplitude restricts this phase to have the form (for each particle)

$$\mathcal{P} \exp \left[ ig_s \int_{\mathcal{C}} dx^\mu \mathbf{A}_\mu(x) \right],$$

where  $\mathcal{C}$  is the spacetime contour of the particle.

- This is a Wilson line!

- The momentum space amplitude is then given by

$$\tilde{\mathcal{A}} = \int d^2 \mathbf{b} e^{-i \mathbf{b} \cdot \mathbf{q}} \langle 0 | \mathcal{W}(p_1, 0) \mathcal{W}(p_2, z) | 0 \rangle,$$

where

$$\mathcal{W}(p, z) = \mathcal{P} \exp \left[ i g_s p^\mu \int_{-\infty}^{\infty} ds \mathbf{A}_\mu(sp + z) \right].$$

- The momentum  $\mathbf{q}$  is conjugate to the impact parameter, and satisfies  $t \simeq -\mathbf{q}^2$ .
- Can now calculate the position space amplitude at one-loop, using dimensional regularisation.

- The answer is (in  $d = 4 - 2\epsilon$  dimensions)

$$\mathcal{A}^{(1)} = \frac{g_s^2 \Gamma(1 - \epsilon) (\mu^2 \mathbf{b}^2)^\epsilon}{4\pi^{2-\epsilon} 2\epsilon} \left[ i\pi \mathbf{T}_s^2 + \mathbf{T}_t^2 \log \left( \frac{s}{-t} \right) + \frac{1}{2} \left( \log \left( -\frac{t}{m^2} \right) - i\pi \right) \sum_{i=1}^4 C_i \right] + \mathcal{O}(\epsilon^0),$$

where

$$\mathbf{T}_s^2 = (\mathbf{T}_1 + \mathbf{T}_2)^2, \quad \mathbf{T}_t^2 = (\mathbf{T}_1 + \mathbf{T}_3)^2$$

are quadratic colour operators for pure  $s$ - and  $t$ -channel exchanges;  $C_i$  the quadratic Casimir of particle  $i$ .

- From the known properties of Wilson lines, we can immediately exponentiate this!

- One then has

$$\mathcal{A} = \exp \left\{ K \left[ i\pi \mathbf{T}_s^2 + \mathbf{T}_t^2 \log \left( \frac{s}{-t} \right) \right] + \dots \right\}, \quad K = \frac{g_s^2 \Gamma(1 - \epsilon)}{4\pi^{2-\epsilon}} \frac{(\mu^2 \mathbf{b}^2)^\epsilon}{2\epsilon}$$

- There are two terms with non-trivial colour dependence:

(I) A  $t$ -channel term:  $\propto \mathbf{T}_t^2 \log(\frac{s}{-t})$ .

(II) A pure *eikonal phase*:  $\propto i\pi \mathbf{T}_s^2$ .

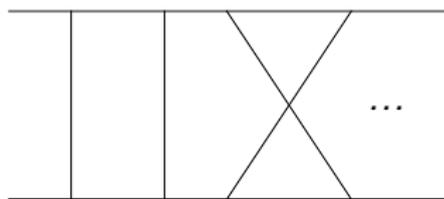
- The former is responsible for *Reggeisation* of  $t$ -channel exchanges:

$$-\frac{i\eta_{\mu\nu}}{q^2} > -\frac{i\eta_{\mu\nu}}{q^2} \left( \frac{s}{-t} \right)^\alpha$$

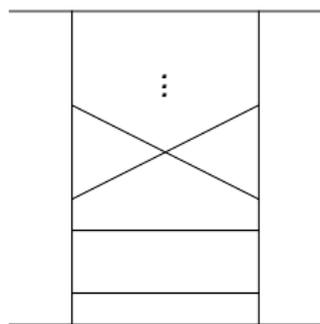
- The latter describes a spectrum of bound states (e.g. positronium).

# EIKONAL PHASE AND REGGE TRAJECTORY

- The eikonal phase comes from horizontal (crossed) ladder diagrams, whereas the Regge trajectory comes from vertical ladders.



(a)



(b)

- In QCD, the vertical ladders dominate.
- It is known that horizontal ladders dominate in gravity: the eikonal phase is enhanced by a factor  $s/(-t)$  w.r.t. the Reggeisation term.
- The Wilson line approach gives an elegant view on this.

- First, we need to find appropriate Wilson lines for gravity.
- Here, we mean specifically the operator describing soft graviton emission.
- The relevant quantity has appeared in various places (Brandhuber, Heslop, Spence, Travaglini; Naculich, Schnitzer; White):

$$\exp \left[ \frac{i\kappa}{2} \int_{\mathcal{C}} ds \dot{x}^{\mu} \dot{x}^{\nu} h_{\mu\nu}(x) \right].$$

- For straight line contours  $x^{\mu} = x_0^{\mu} + p^{\mu} s$ , this becomes

$$\exp \left[ \frac{i\kappa}{2} p^{\mu} p^{\nu} \int_{\mathcal{C}} ds h_{\mu\nu}(x) \right].$$

- Closely related to its QCD counterpart!

- The Wilson line approach for the QCD Regge limit (Korchemsky, Korchemskaya) can be ported directly to gravity.
- The momentum space gravity amplitude is given by

$$\tilde{\mathcal{M}} = \int d^2\mathbf{b} e^{-i\mathbf{b}\cdot\mathbf{q}} \langle 0 | \mathcal{W}_g(p_1, 0) \mathcal{W}_g(p_2, z) | 0 \rangle,$$

where

$$\mathcal{W}_g(p, z) = \exp \left[ \frac{i\kappa}{2} p^\mu p^\nu \int_{-\infty}^{\infty} ds h_{\mu\nu}(sp + z) \right].$$

- Exponentiation of the one-loop calculation can be carried out as before.

- One finds

$$\mathcal{M} = \exp \left\{ -K_g (\mu^2 \mathbf{b}^2)^\epsilon \left[ i\pi s + t \log \left( \frac{s}{-t} \right) \right] + \mathcal{O}(\epsilon^0) \right\},$$

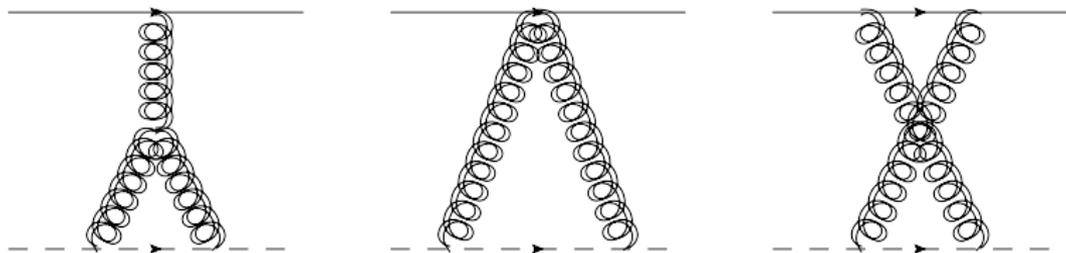
$$K_g = \left( \frac{\kappa}{2} \right)^2 \frac{\Gamma(1 - \epsilon)}{8\pi^{2-\epsilon}}.$$

- The eikonal phase wins as  $\frac{s}{-t} \rightarrow \infty$ , in contrast to QCD.
- However, the structure of the result is basically the same, and can be obtained by the procedure

$$g_s \rightarrow \frac{\kappa}{2}; \quad \mathbf{T}_{s,t}^2 \rightarrow s, t; \quad C_i \rightarrow 0.$$

- This is the BCJ double copy! (see also [Akhoury, Saotome; Sabio Vera, Campillo, Vazquez-Mozo, Johansson](#)).

- Diagrammatic study of Regge limit by [Akhoury, Saotome, Sterman](#).



- Considered a light particle scattering on a black hole.
- Next-to-soft corrections lead to a modified eikonal phase:

$$\chi \rightarrow \chi_E + \chi_{NE},$$

where  $\chi_{NE} \propto R_s$  (Schwarzschild radius of black hole).

- Correction corresponds to deflection angle of light particle (see also [D'Appollonio, Di Vecchia, Russo, Veneziano](#); [Bjerrum-Bohr, Donoghue, Holstein, Plante, Vanhove](#)).

- Can also extend the Wilson line approach to next-to-soft level (Lanen, Stavenga, White).
- Has been applied to the Regge limit in both QCD and gravity (Luna, Melville, Naculich, White).
- General case of two massive particles.
- In QCD, get a power-suppressed correction to the Regge trajectory of the gluon.
- In gravity, the correction to the NE phase corresponds to two simultaneous deflection angles for the colliding particles (as conjectured by Andersen, Schmidt-Sommerfeld, Giddings).
- Previous results of Akhoury, Saotome, Sterman emerge as a special case.

- Next-to-soft corrections are relevant to transplanckian scattering in gravity.
- More generally, similar methods can be applied to understand radiation from scattering black holes.
- Full solutions for colliding shockwaves / black holes are not always known.
- The next-to-soft calculation allows us to build them up perturbatively i.e. order-by-order in the deflection angle.
- Methods exist for relating QCD and gravity results.

- (Next-to)-soft physics has a large number of applications, in different areas of physics.
- For hep-ph, hep-ex: increased precision for collider observables.
- For hep-th, gr-qc: transplanckian scattering in gravity, radiation in black hole scattering.
- Common languages for QCD and gravity (e.g. Wilson lines) make underlying structures / common behaviour clearer.

- Can we resum next-to-leading power (NLP) threshold logs?
- Other applications in precision physics?
- Do next-to-soft methods help in calculating radiation from scattering black holes?
- What are gravitational Wilson lines useful for?
- What does anything in this talk have to do with BMS symmetry?

# THANKS FOR LISTENING!

