





The 12 Days of Christmas:

PhD edition sing-along

Lucy Kotsiopoulou Supervisor: Dr. Andrzej Szelc

SING with me!

(Or else... Krampus will eat you :))



NASA/JPL-Caltech/Harvard-Smithsonian CfA

1: PhD

On the ... day of Christmas my true love gave to me:

1: PhD

My first day at the office in September 2023!



2: Experiments

On the ... day of Christmas my true love gave to me:

1: PhD

2: Experiments

I work on 2 experiments, SBND and DarkSide-20k!



3: Flavours

On the ... day of Christmas my true love gave to me:

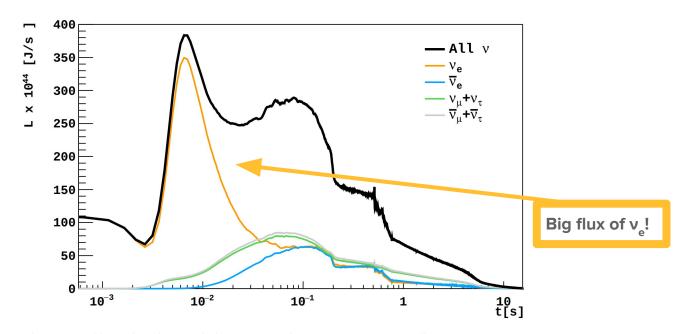
1: PhD

2: Experiments

3: Flavours

All 3 neutrinos come out a supernova explosion!

Neutrino flux of supernova over time:



Sensitivity of future liquid argon dark matter search experiments to core-collapse supernova neutrinos, **The DarkSide-20k Collaboration**, (2020).

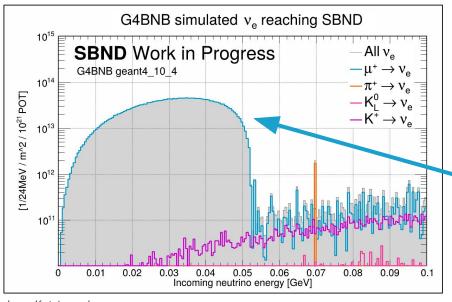
4: Fluxes

On the ... day of Christmas my true love gave to me:

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- 2: Experiments
- 3: Flavours
- 4: Fluxes

Man-made neutrino beam at Fermilab makes 4 v fluxes!

"Low energy" (<100MeV) neutrino fluxes reaching SBND:



Big flux of v_e from μ decaying at rest!

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Fermilab neutrino beam

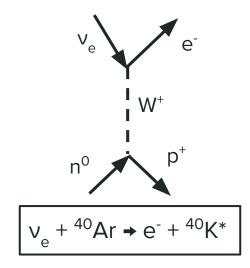


5: Particles

On the ... day of Christmas my true love gave to me:

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- 4: Fluxes
- 5: Particles

v_e on Argon cross section we want to measure! Observe the process involving 5 particles:



Charged current inelastic cross section at <100 MeV energies

6: Materials

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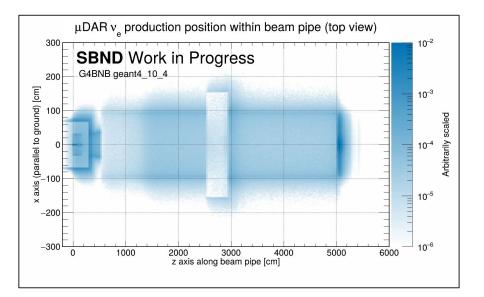
4: Fluxes

5: Particles

6: Materials

Muons can bump on to 6 things and decay at rest to v_e !

Top view on beam decay pipe:



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Concrete
Beryllium
Dolomite
Steel
Aluminium
Air

7: μs

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4: Fluxes

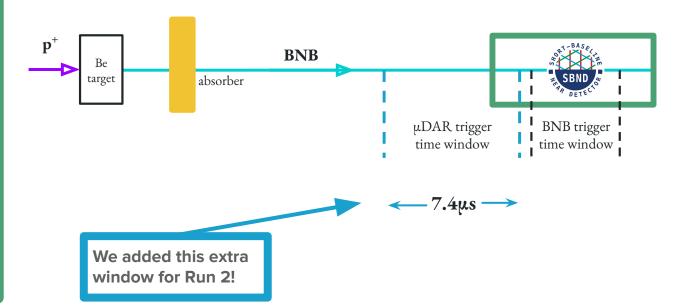
5: Particles

6: Materials

7: μs

We "open" trigger system for 7.4μs to catch these ν !

Trigger windows for SBND:



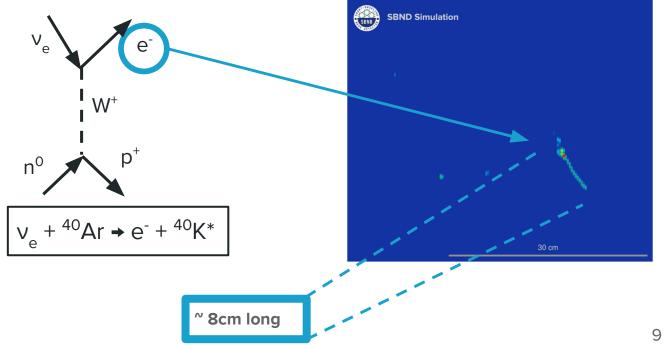
8: cm

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- 5: Particles
- 6: Materials
- 7: μs
- 8:

Final state e- of interaction is 8cm long in SBND

Neutrino flux vs. time of a Type II core-collapse supernova:

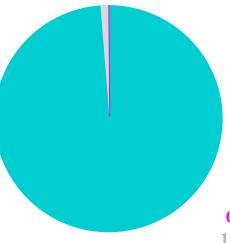


9:99%

On the ... day of Christmas my true love gave to me:

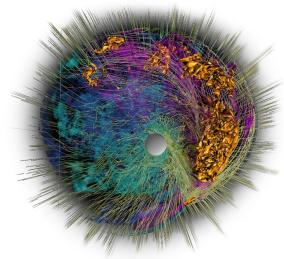
- 1: PhD
- 2: Experiments
- 3: Flavours
- 4: Fluxes
- 5: Particles
- 6: Materials
- 7: µs
- 8: cm
- 9:99%

Neutrinos carry away 99% of the entire supernova energy! (For a Type II core-collapse supernova)



0.01% EM radiation1% Kinetic energy of external star's layers

99% veutrinos



Neutrinos from a core collapse supernova, Blondin and Mezzacappa (2016).

10: kiloparsec

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7: µs

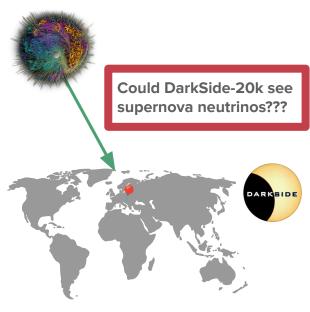
8: cm

9: 99%

10: kiloparsec

Now studying supernovae 10kpc away (" Earth to centre of galaxy)





Red lines are a 10kpc distance

11: Solar masses

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6: Materials

7: μs

8: cm

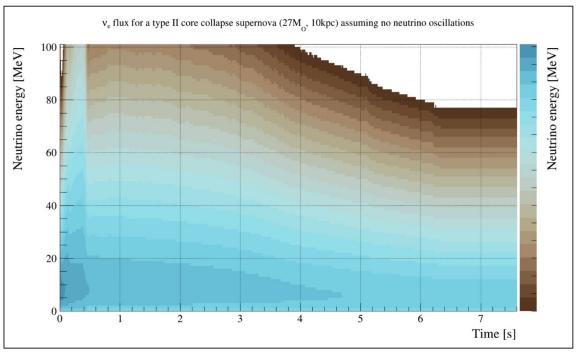
9:99%

10: kpc

11: Solar masses

Now Will study supernovae of 11M progenitor star mass!

Observe neutrino flux over time and energy of 27M_{__} (didn't have the other plot :()



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12: LTA months

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6: Materials

7: μs

8: cm

9:99%

10: kpc

11: Solar masses

12: LTA months

Had an amazing time on LTA in Fermilab (Chicago)!

Thank you PPE for sending me :D



Hosted a wine club at Fermilab





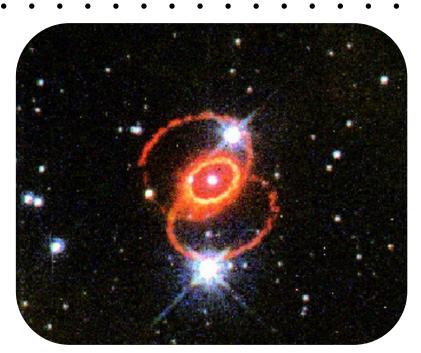
Made many new wonderful friends:)



Held a parrot!



Thanksgiving in Fermilab control room!



A photo of Supernova 1987A's remnants as seen in 1995 by the Hubble space telescope, GSFC/NOAO/NASA/ESA/STIS Instrument Definition Team (1995)

Thank you!

Backup slides