

Domain-wall-seeded bubble nucleation on the lattice

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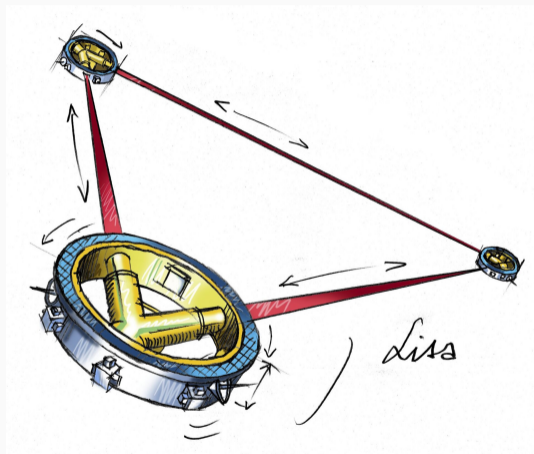
May 18, 2026

[arXiv:TBA]

With Simone Blasi, Andreas Ekstedt and Kari Rummukainen

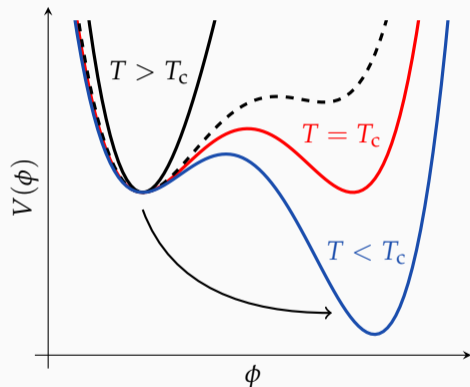
Cosmological phase transitions

- Many extensions of the standard model can have first order phase transitions
- These are expected to produce a model-dependent gravitational wave background
- Future GW observatories might be able to rule out models based on the detected background



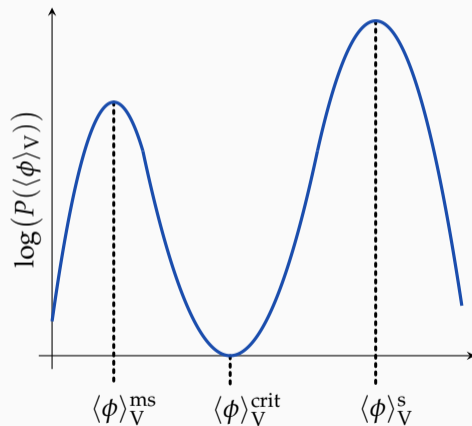
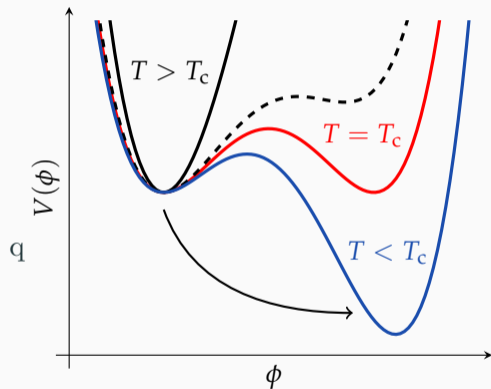
[European Space Agency]

First-order phase transitions



- First-order phase transitions happen through bubble nucleation from the metastable to the stable phase
- The nucleation rate tells us how often bubbles of the stable phase appear

First-order phase transitions



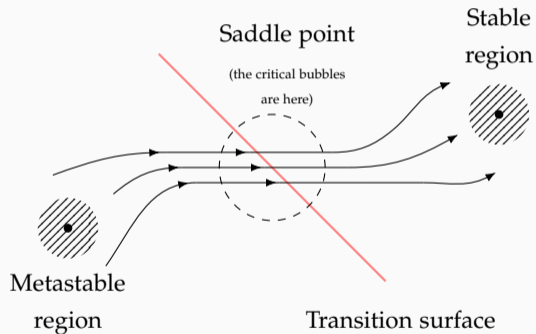
Computing the nucleation rate (perturbatively)

- For suppressed transitions the rate Γ is mostly determined by the probability of critical configurations
- A saddle-point approximation gives

$$\Gamma \approx A \exp\{-B\},$$

where B is the “bounce-action” and A contains dynamical information

- Both can be computed semi-analytically for spherical bubbles



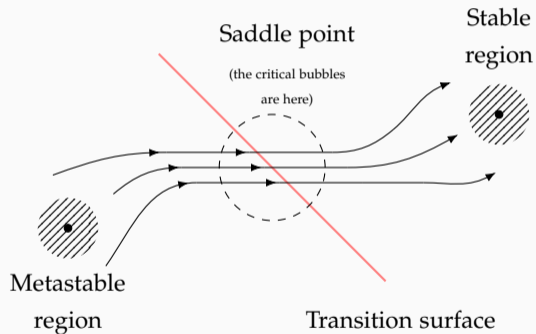
[Langer 1969; Coleman 1977; Callan et al. 1977]

Computing the nucleation rate (non-perturbatively)

- Two scenarios:

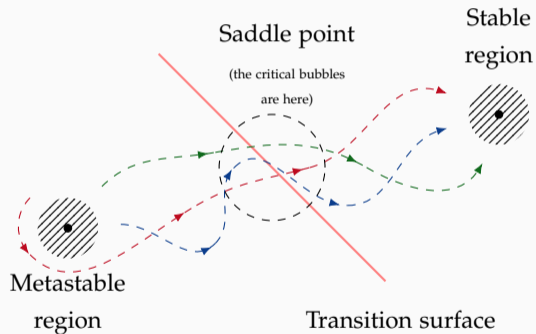
1. Small nucleation rate

⇒ sample critical configurations
using MCMC [Moore et al. 2001]



Computing the nucleation rate (non-perturbatively)

- Two scenarios:
 1. Small nucleation rate
 - ⇒ sample critical configurations using MCMC [Moore et al. 2001]
 2. Large nucleation rate
 - ⇒ real-time simulations of the metastable phase



Seeded nucleation

- Almost all observed nucleation is seeded by impurities
- Inhomogeneities such as topological defects could also enhance nucleation rates in the early universe
 - Cosmic strings [\[Steinhardt 1981\]](#)
 - Monopoles [\[Preskill et al. 1993\]](#)
 - Domain walls

⋮



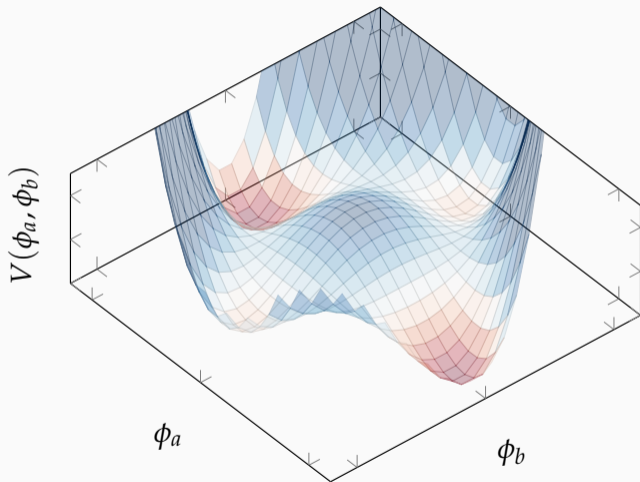
[Wikimedia Commons]

The setup: two scalar fields in 2 + 1 D

$$\mathcal{L} = \sum_{\alpha=a,b} \left[\frac{1}{2} (\partial_\mu \phi_\alpha)^2 - \frac{1}{2} m_\alpha^2 \phi_\alpha^2 \right] \\ + \frac{\lambda_a}{24} \phi_a^4 + \frac{\lambda_b}{24} \phi_b^4 + \frac{\mu}{4} \phi_a^2 \phi_b^2$$

$$m_a^2 = 0.09, \quad m_b^2 = 0.226, \\ \lambda_a = 0.006, \quad \lambda_b = 0.096, \\ \mu = 0.026 \text{ (0.028)}$$

(In units of some mass scale M^2)

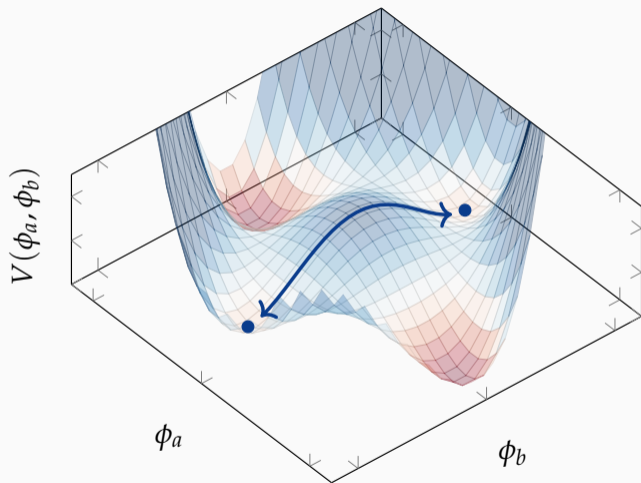


- Two-step transition:

$$(\phi_a, \phi_b): (0, 0)$$

$$\xrightarrow{\text{DWs form}} (0, \pm v_b)$$

$$\xrightarrow{\text{1st order}} (\pm v_a, 0)$$



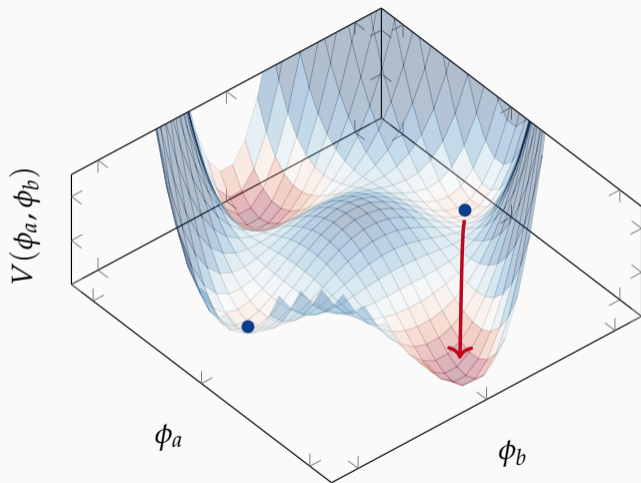
Domain-wall-seeded nucleation

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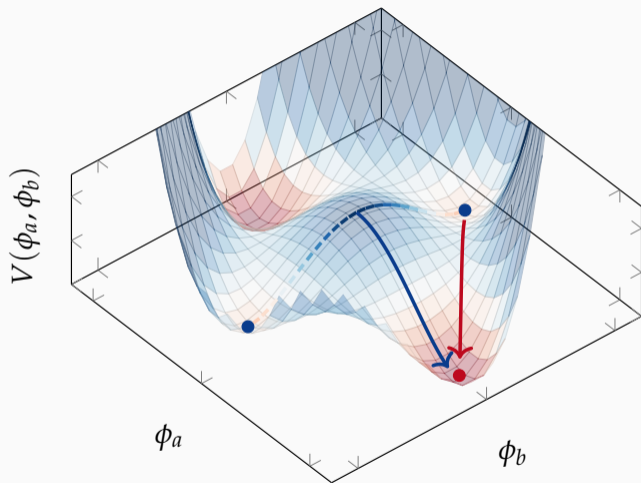
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- At the domain wall $\phi_b \approx 0$
and only gradient terms
prevent the transition



How to estimate the rate analytically?

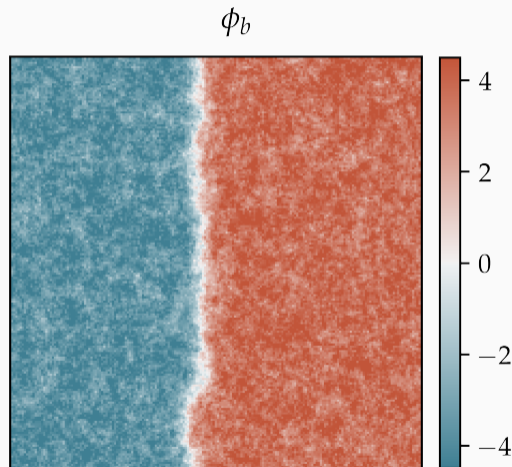
[Rajaraman 1982; Blasi et al. 2022]

- Decompose both fields perpendicular to the domain wall

$$\phi_b = DW(\mathbf{x}) + \sum_n \phi_{b,n}(y) \psi_n(\mathbf{x})$$

$$\phi_a = \sum_n \phi_{a,n}(y) \zeta_n(\mathbf{x}),$$

- Perform integrals numerically to obtain a 1D-theory with infinitely many fields



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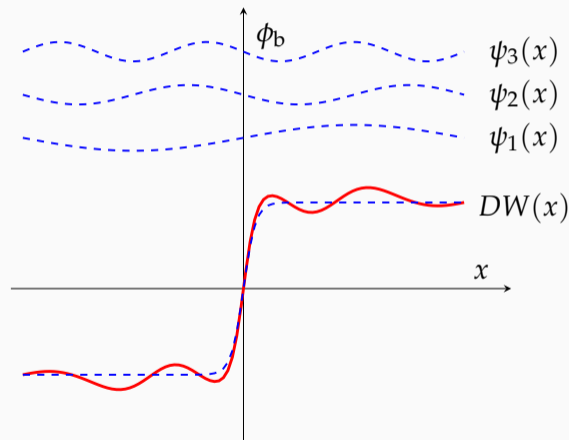
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$$\implies \mathcal{L}^{1D} = \sum_{i=0}^{\infty} \sum_{\alpha=a,b} \left[\frac{1}{2} (\partial_y \phi_{\alpha,i})^2 + \frac{1}{2} m_{\alpha,i}^2 \phi_{\alpha,i}^2 \right] + V(\phi_{a,j}, \phi_{b,k}),$$

where $m_{\alpha,0}^2 < m_{\alpha,1}^2 < \dots$

- Heavier fields can be integrated out to yield an EFT for $\phi_{a,0}$:

$$\mathcal{L}_{a,0}^{1D} = \frac{1}{2} (\partial_y \phi_{a,0})^2 + \frac{1}{2} m_{a,0}^2 \phi_{a,0}^2 + V(\phi_{a,0})$$

\implies a single field homogeneous nucleation problem in 1D

On the lattice:

- Antiperiodic boundary conditions perpendicular to the wall
- Stochastic Hamiltonian equations

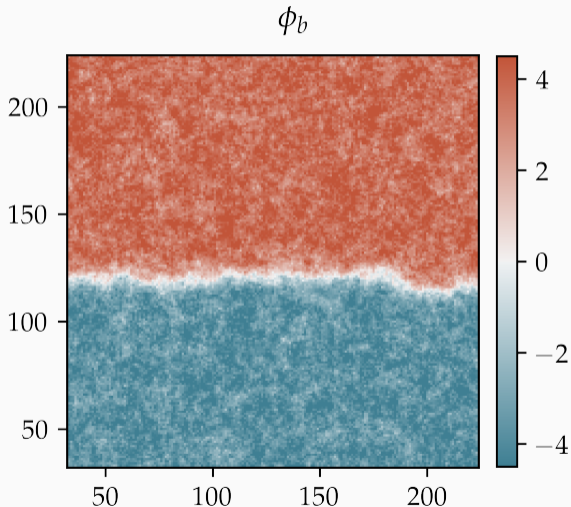
$$\frac{\partial \phi_\alpha}{\partial t} = \frac{\delta H}{\delta \pi_\alpha}, \quad \frac{\partial \pi_\alpha}{\partial t} = -\frac{\delta H}{\delta \phi_\alpha} - \gamma \frac{\delta H}{\delta \pi_\alpha} + \zeta$$

- Integrate numerically and take ζ into account by updating every $\Delta t \ll a$

$$\pi_\alpha(x, t) = (1 - \epsilon^2)^{1/2} \pi_\alpha(x, t) + \epsilon \eta_\alpha(x, t),$$

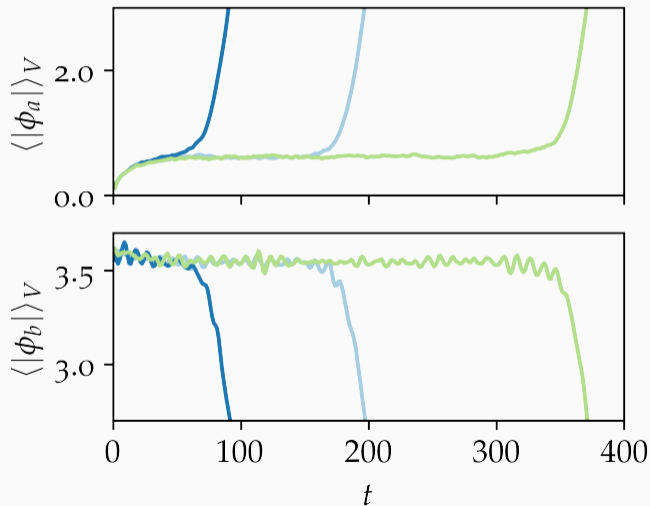
where $\epsilon^2 = 1 - \exp(-2\gamma\Delta t)$.

[Moore et al. 2001]

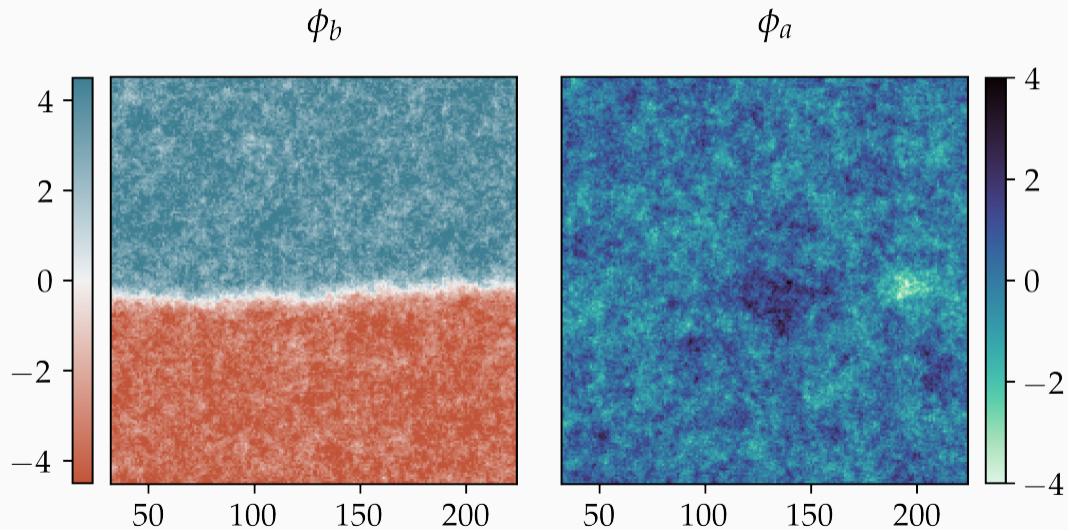


Simulation procedure

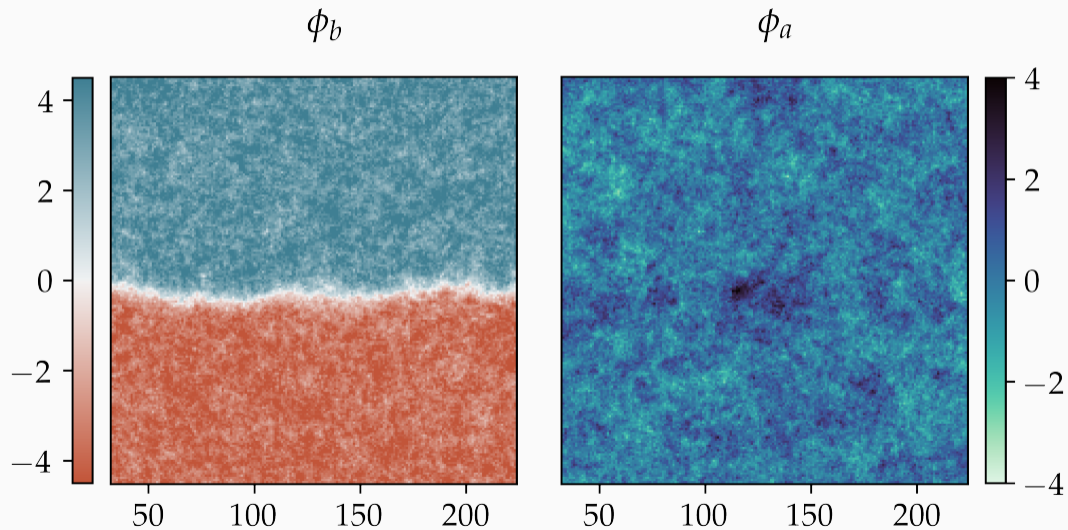
1. Initialize fields with anti-periodic boundary conditions to one direction
2. Evolve only ϕ_b to form the wall
3. Evolve both fields until $\langle |\phi_a| \rangle_V$ reaches some semi-arbitrary threshold value
4. Record t_i
5. See 1.



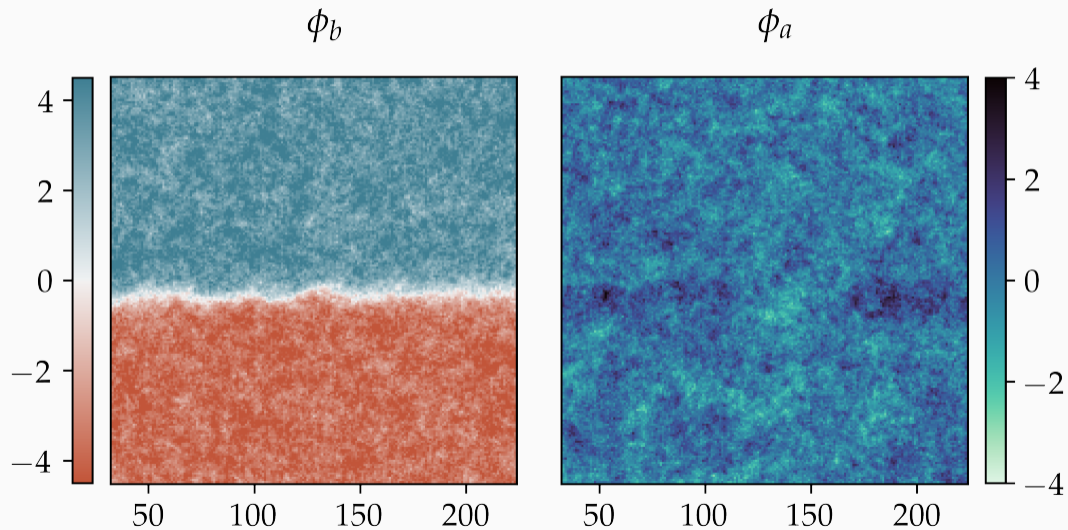
What it looks like



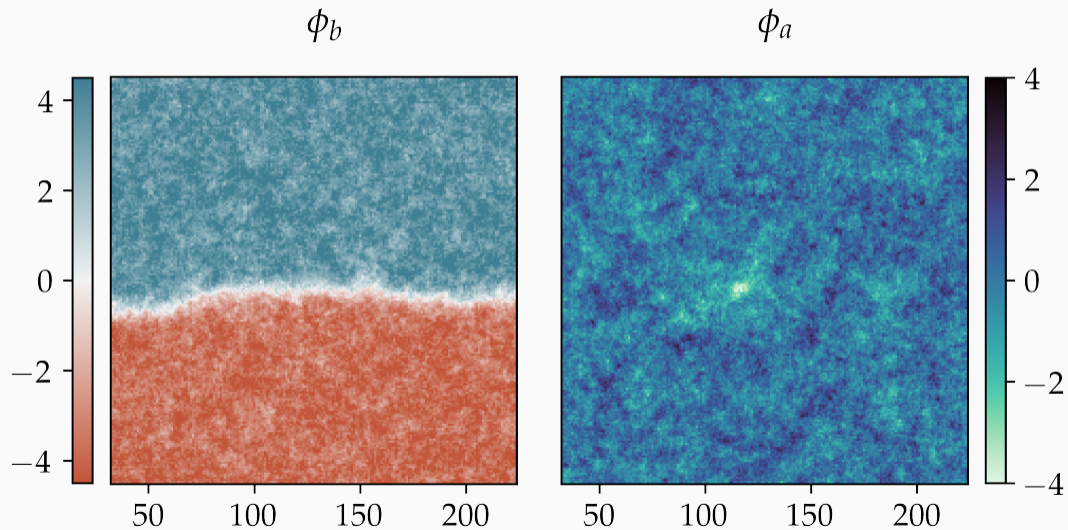
What it looks like



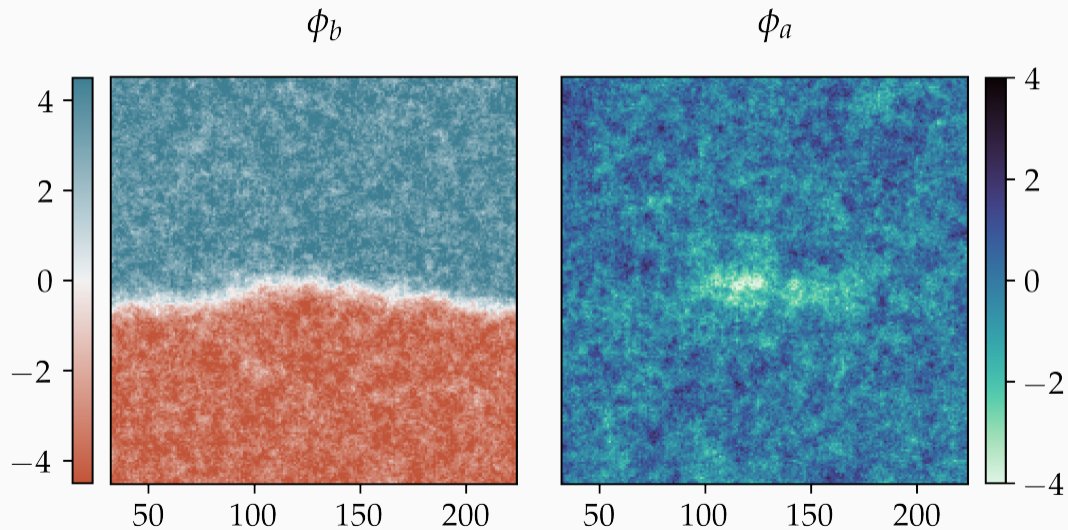
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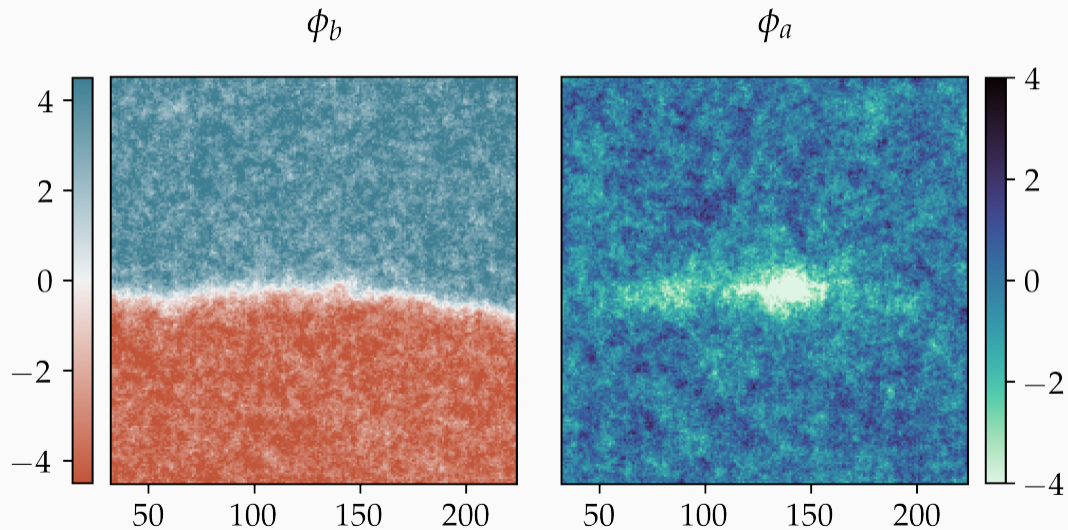
What it looks like



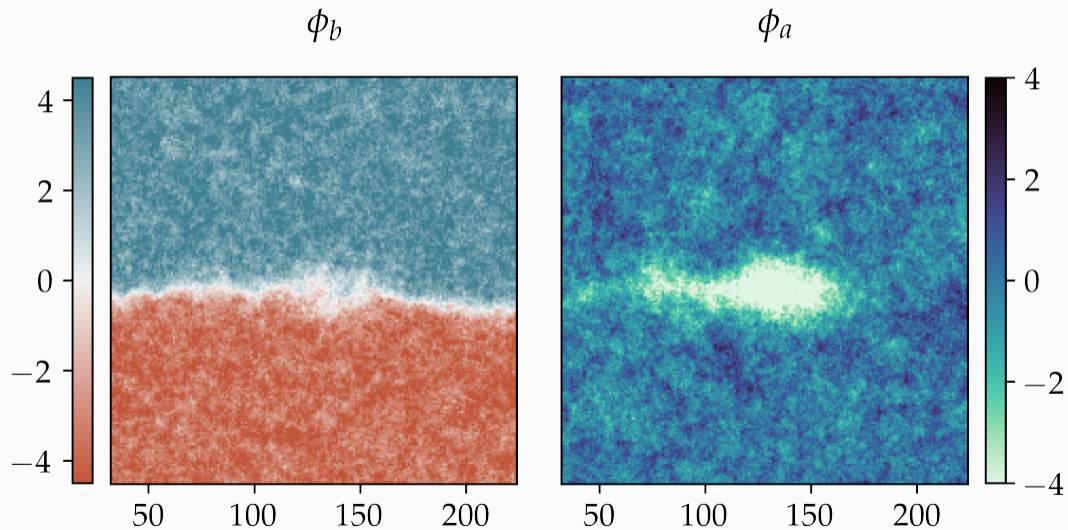
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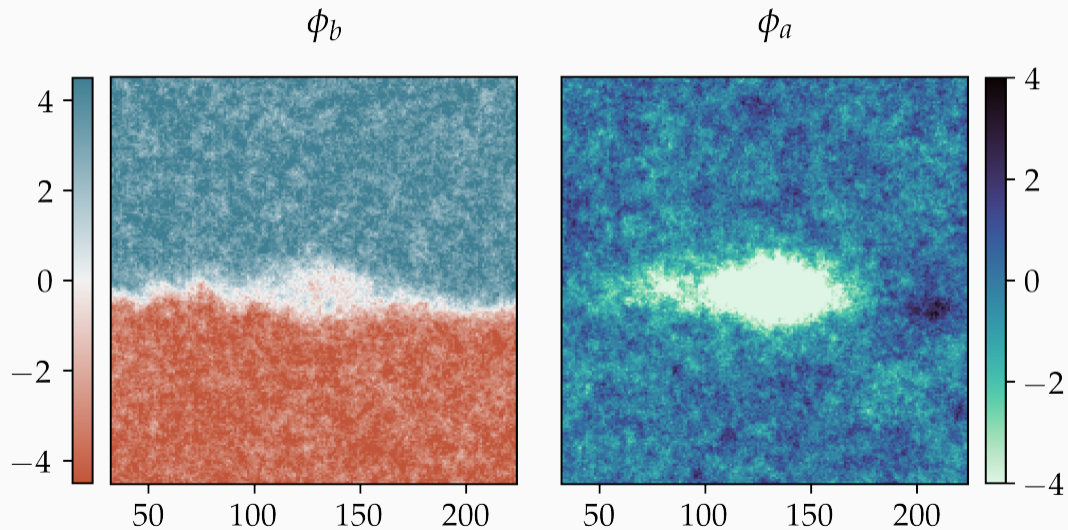
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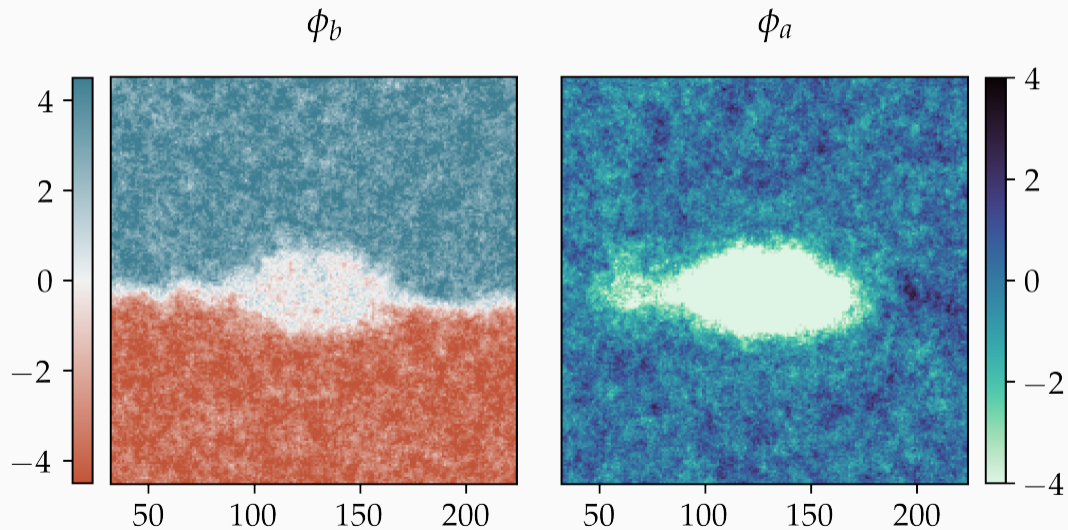
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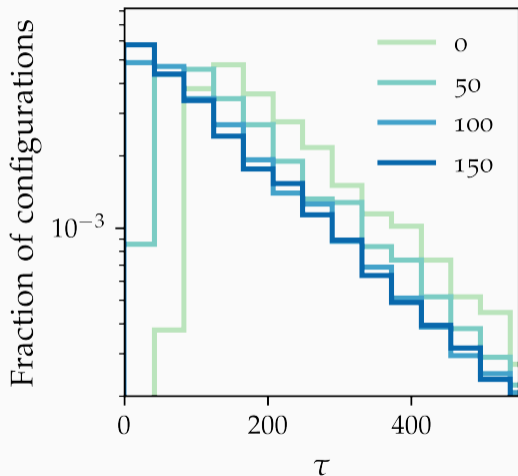


Analysing the nucleation times

- To control dependence on the initial conditions, construct from the full set of nucleation times $\{t_i\}$ the sets

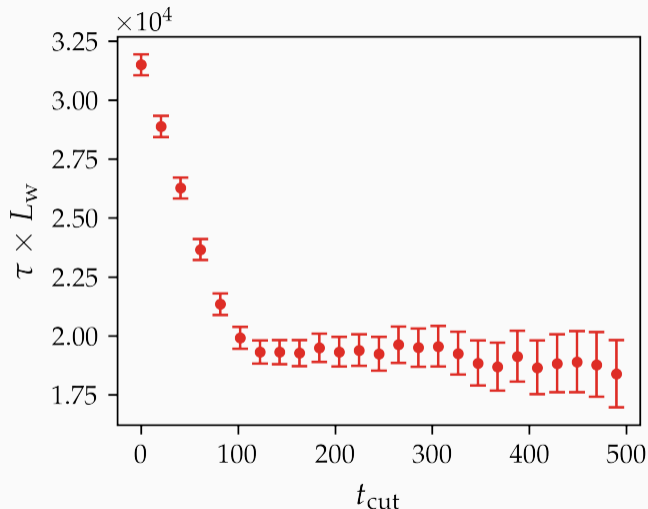
$$\{t_i - t_{\text{cut}} \mid t_i > t_{\text{cut}}\}$$

- Compute the mean nucleation time τ for larger and larger values of t_{cut}
- Extract the result from behaviour at large t_{cut}

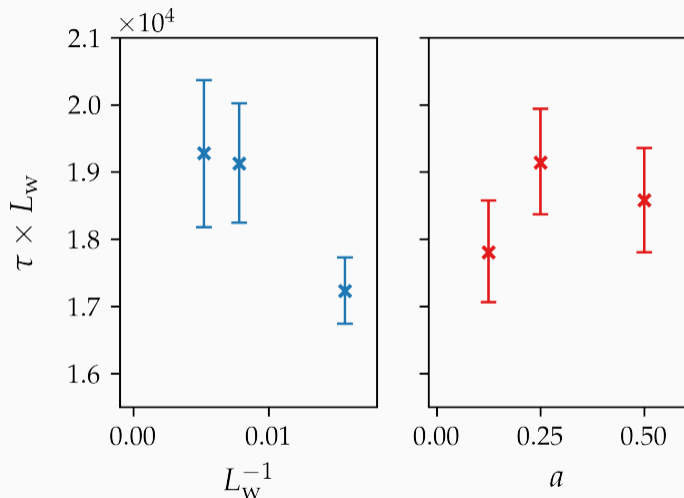


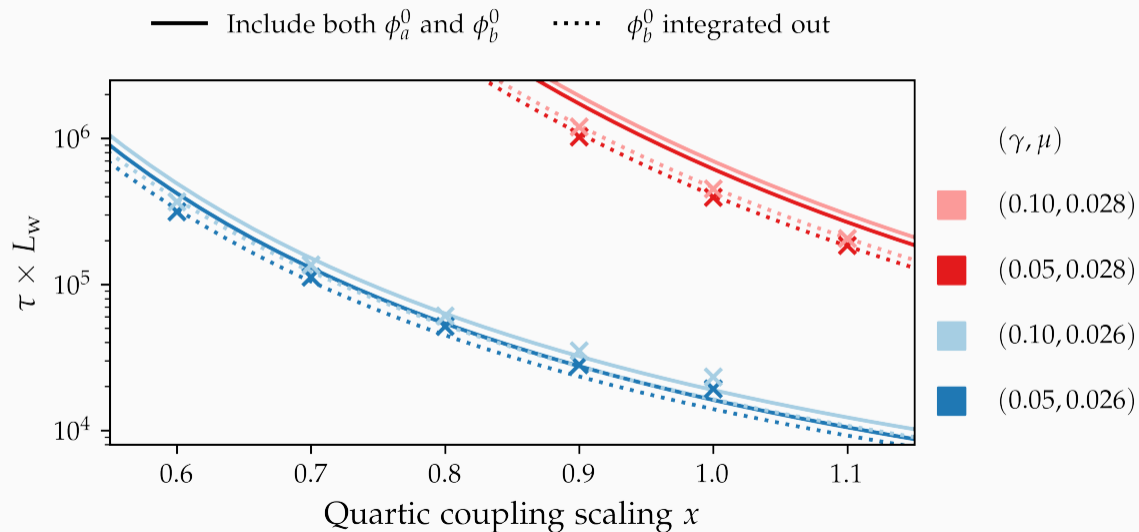
Analysing the nucleation times

- τ stabilises as $t_{\text{cut}} \rightarrow \infty$
- Works well if the thermalisation time is not much longer than the average nucleation time τ
- Estimate statistical errors by resampling the initial set of nucleation times $\{t_i\}$



- Moderate dependency on the lattice spacing
- Small dependency on the length of the domain wall
- No discernible dependency on the value of $\langle |\phi_a| \rangle$ used as the nucleation threshold





What's next?

DW-seeded nucleation in the
Singlet Extended SM:

- Mapping rules to a dimensionally reduced 3D theory exist [Niemi et al. 2024]
 - ⇒ Easy to simulate on a lattice
- Too slow to use pure real-time approach
 - ⇒ Sample bubbles using MCMC

