

Inclusive semileptonic decays using lattice QCD

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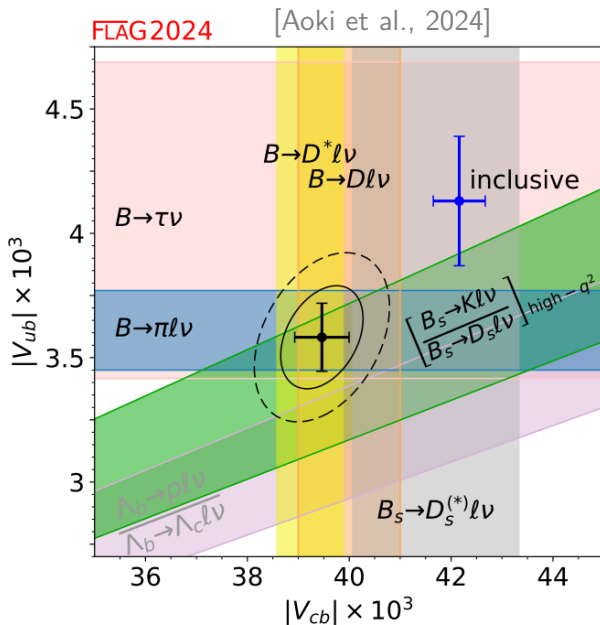
*In collaboration: Alessandro Barone, Shoji Hashimoto, Zhi Hu, Takashi Kaneko,
Ryan Kellermann*

Motivation

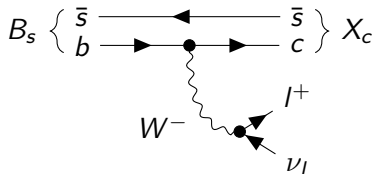
Different method of calculation.

- ▶ Exclusive $|V_{cb}| =$
Lattice QCD
- ▶ Inclusive $|V_{cb}| =$
Perturbative methods

Our aim is to calculate the **inclusive** $|V_{cb}|$ using Lattice QCD.



Focus on $B_s \rightarrow X_c \bar{\nu}_l l$



$$H_W = \frac{G_F}{\sqrt{2}} V_{cb} \underbrace{\bar{\nu}_l \gamma^\mu l}_J^\mu \underbrace{\bar{b}_L \gamma_\mu c_L}_{J_\mu}$$

$$\mathcal{M} = \frac{G_F}{\sqrt{2}} V_{cb} \langle l, \bar{\nu}_l | J_L^\mu | 0 \rangle \langle X_c | J_\mu | B_s \rangle$$

$$\frac{d\Gamma}{dq^2 dq_0 dE_l} = \frac{G_F^2 |V_{cb}|^2}{8\pi^3} L_{\mu\nu} W^{\mu\nu}$$

$$\text{where } q = p_{B_s} - p_{X_c}$$

$$\frac{d\Gamma}{dq^2 dq_0 dE_l} = \frac{G_F^2 |V_{cb}|^2}{8\pi^3} L_{\mu\nu} W^{\mu\nu}$$

Leptonic part fully kinematically determined,

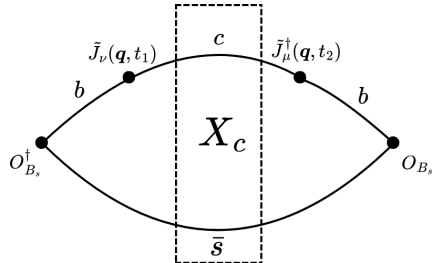
$$L^{\mu\nu} = p_\ell^\mu p_{\nu_\ell}^\nu + p_\ell^\nu p_{\nu_\ell}^\mu - g^{\mu\nu} (p_\ell \cdot p_{\nu_\ell}) - i\epsilon^{\mu\alpha\nu\beta} p_{\ell,\alpha} p_{\nu_\ell,\beta}$$

Hadronic tensor contains non-perturbative info,

$$W^{\mu\nu} \sim \sum_{X_c} \delta^{(4)}(p_{B_s} - q - p_{X_c}) \langle B_s | \tilde{J}^{\mu\dagger}(0) | X_c \rangle \langle X_c | \tilde{J}^\nu(0) | B_s \rangle$$

$$\begin{aligned} C_{\mu\nu}(\mathbf{q}, t) &= \int_0^\infty d\omega W_{\mu\nu}(\mathbf{q}, \omega) e^{-\omega t} \\ &\sim \langle B_s | \tilde{J}_\mu^\dagger(\mathbf{q}, 0) e^{-\hat{H}t} \tilde{J}_\nu(\mathbf{q}, 0) | B_s \rangle \end{aligned}$$

where $t = t_2 - t_1$, and $\omega = E_{X_c}$.



[Hansen et al., 2017, Hashimoto, 2017, Hansen et al., 2019, Gambino and Hashimoto, 2020]

[Bailas et al., 2020, Gambino et al., 2022, Barone et al., 2023, De Santis et al., 2025b]

Integrate differential decay width

$$\frac{d\Gamma}{dq^2} \propto \int_{\omega_0}^{\omega_{\max}} d\omega k^{\mu\nu}(\mathbf{q}, \omega) W_{\mu\nu}(\mathbf{q}, \omega) = \bar{X}(\mathbf{q})$$

where ω is the energy of the final state hadron.

$$\bar{X}(\mathbf{q}) = \int_{\omega_0}^{\infty} d\omega W_{\mu\nu}(\mathbf{q}, \omega) k^{\mu\nu}(\mathbf{q}, \omega) \theta(\omega_{\max} - \omega)$$

$$\bar{X}(\mathbf{q}, \sigma) = \int_{\omega_0}^{\infty} d\omega W_{\mu\nu}(\mathbf{q}, \omega) k^{\mu\nu}(\mathbf{q}, \omega) \theta_{\sigma}(\omega_{\max} - \omega)$$

$$\bar{X}(\mathbf{q}, \sigma, N) \approx \sum_t^N c_{\mu\nu, t}(\mathbf{q}, \sigma) \underbrace{\int_{\omega_0}^{\infty} d\omega W^{\mu\nu}(\mathbf{q}, \omega) e^{-\omega t}}_{C^{\mu\nu}(\mathbf{q}, t)}$$

*Sum over
 $t = t_2 - t_1$ per
channel

Need to then take appropriate limits

$$\bar{X}(\mathbf{q}) = \lim_{\sigma \rightarrow 0} \lim_{N \rightarrow \infty} \bar{X}(\mathbf{q}, \sigma, N)$$

[Hansen et al., 2019], [Bailas et al., 2020], [Barone et al., 2023], [De Santis et al., 2025b].

Integrate differential decay width

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$$\bar{X}(\mathbf{q}, \sigma) = \int_{\omega_0}^{\infty} d\omega W_{\mu\nu}(\mathbf{q}, \omega) k^{\mu\nu}(\mathbf{q}, \omega) \theta_{\sigma}(\omega_{\max} - \omega)$$

$$\bar{X}(\mathbf{q}, \sigma, N) \approx \sum_t^N \tilde{c}_{\mu\nu, t}(\mathbf{q}, \sigma) \underbrace{\int_0^{\infty} d\omega W^{\mu\nu}(\mathbf{q}, \omega) T_t(e^{-\omega})}_{\sim \langle T_t(\mathbf{q}) \rangle_{\mu\nu} = \sum_j^t b_t^{(j)} c_{\mu\nu}(j) / c_{\mu\nu}(0)} \quad (1)$$

Spectral reconstruction using Chebyshevs

Need to then take appropriate limits

$$\bar{X}(\mathbf{q}) = \lim_{\sigma \rightarrow 0} \lim_{N \rightarrow \infty} \left(\lim_{V \rightarrow \infty} \right) \bar{X}(\mathbf{q}, \sigma, N) \quad (2)$$

Technical simulation details

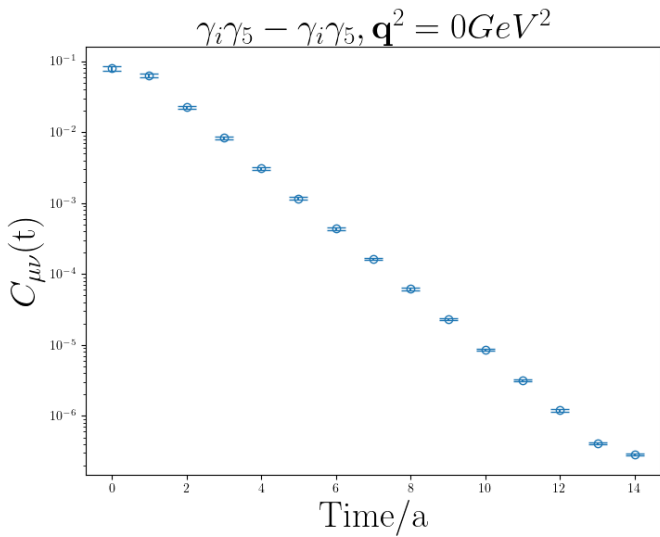
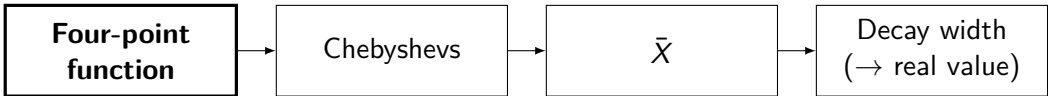
Simulations were carried out on the **DiRAC** Extreme Scaling service at the University of Edinburgh using the Grid [Boyle et al., <https://github.com/paboyle/Grid>] and Hadrons [Portelli et al., <https://github.com/aportelli/Hadrons>] software packages. We use RBC/UKQCD ensembles [Allton et al., 2008].

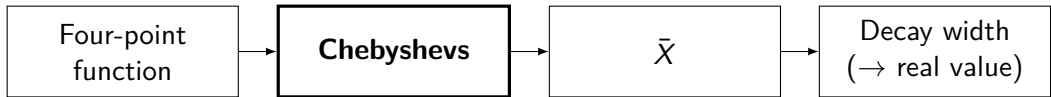
Parameters:

name	a/fm	$L^3 \times T$	M_π/MeV
C1M	0.11	$24^3 \times 64$	276
M1M	0.08	$32^3 \times 64$	286
F1M	0.07	$48^3 \times 96$	232

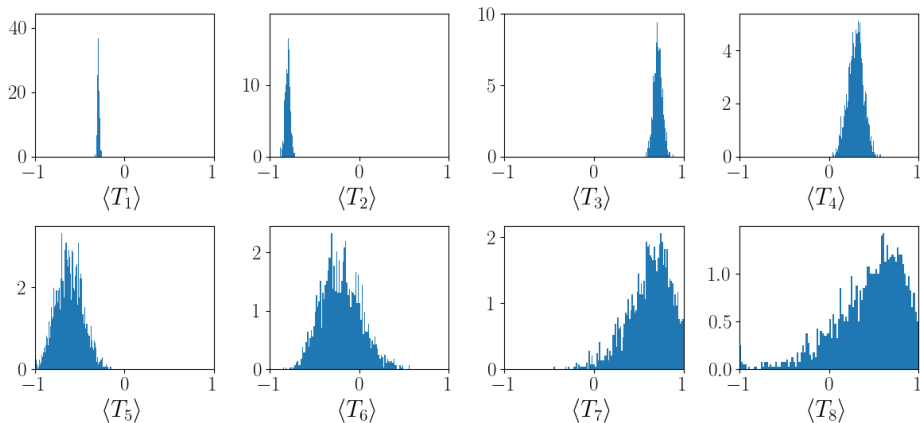
Actions:

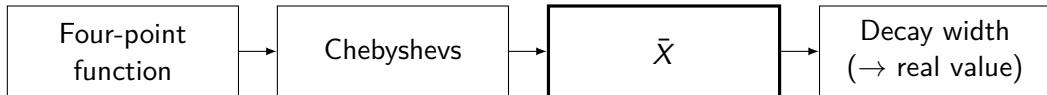
- ▶ b quark simulated at its physical mass (**RHQ action**)
[El-Khadra et al., 1997],[Christ et al., 2007] ,[Lin and Christ, 2007]
- ▶ s, c quarks simulated at near-to-physical mass (**DWF action**)
[Brower et al., 2017]



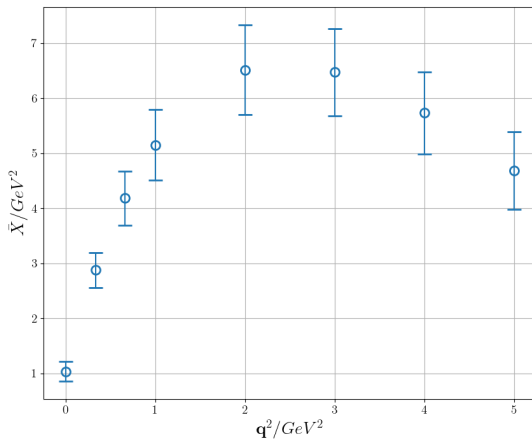


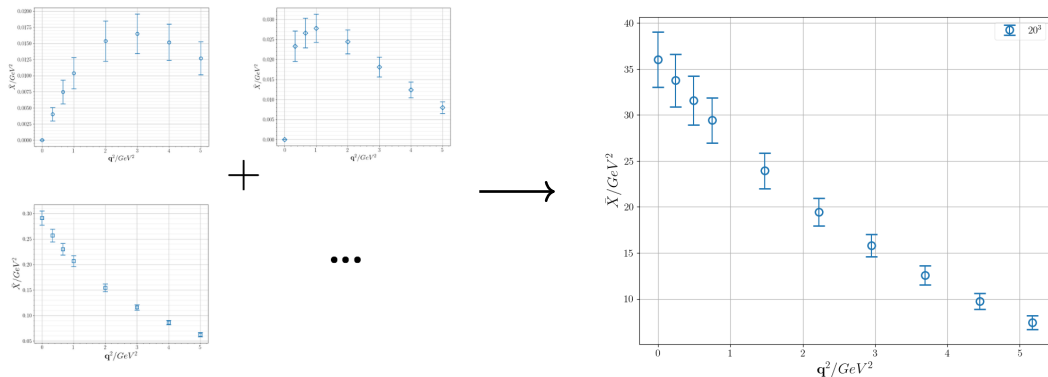
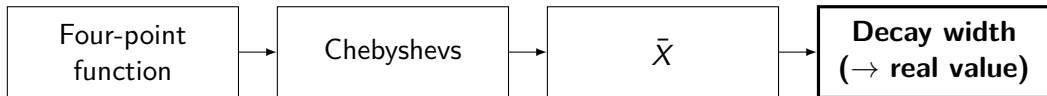
$$\langle T_t(\mathbf{q}) \rangle_{\mu\nu} = \sum_j^t b_t^{(j)} C_{\mu\nu}(j) / C_{\mu\nu}(0)$$

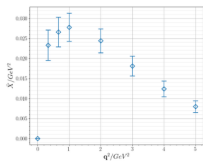
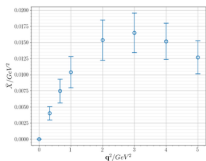
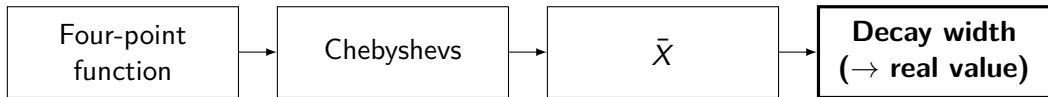




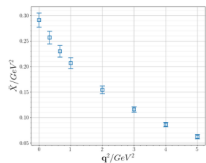
$$\bar{X}(\mathbf{q}^2, \sigma, N) = \sum_t^N \tilde{c}_{\mu\nu, t}(\mathbf{q}, \sigma) \langle T_t(\mathbf{q}) \rangle_{\mu\nu}$$





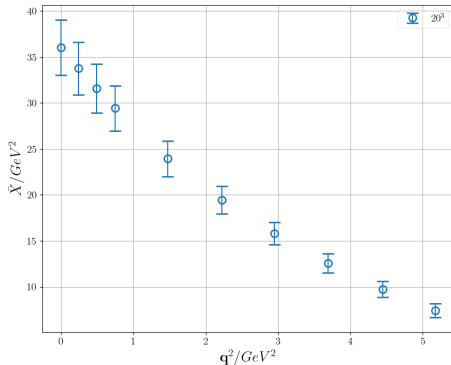


+



...

\rightarrow



Need to extrapolate to continuum

Extrapolations

1. Infinite volume extrapolation (study)
2. Continuum extrapolation (+ chiral/heavy quark etc...)
3. Smearing limit*

Smearing of $W_{\mu\nu}$

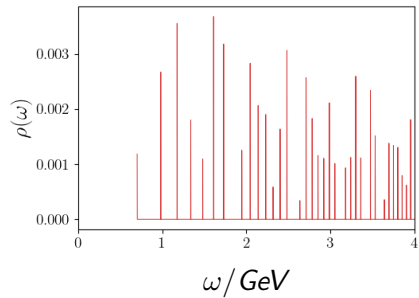
Introducing the Sigmoid is effectively convolving with smearing kernel.

$$\int_0^\infty d\omega W_{\mu\nu}(\mathbf{q}, \omega) k^{\mu\nu}(\mathbf{q}, \omega) \theta_\sigma(\omega_{\max} - \omega) \rightarrow \int_{-\infty}^\infty dy \theta(\omega_{\max} - y) (W^{\mu\nu} k_{\mu\nu} * \delta'_\sigma)(y)$$

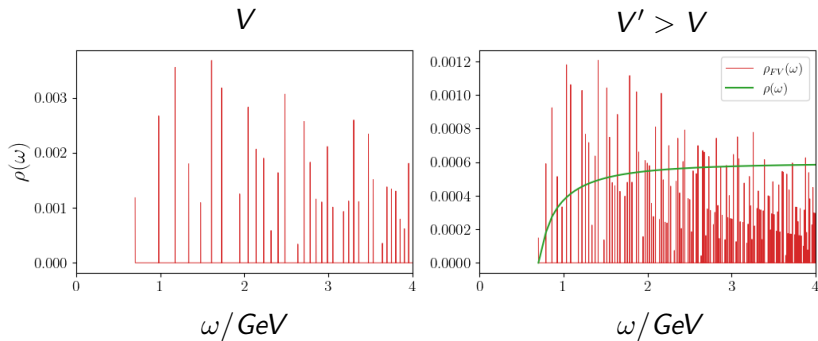
Model $W_{\mu\nu}(\omega)$ as two-body non-interacting spectral density, $\rho(\omega)$.

$$\rho(\omega) = \pi \int \frac{d^3q}{(2\pi)^3} \frac{1}{4\epsilon_D \epsilon_K} \delta(\omega - \epsilon_K + \epsilon_D)$$
$$\rho_V(\omega) = \frac{\pi}{V} \sum_{\mathbf{q}} \frac{1}{4\sqrt{\mathbf{q}^2 + m_K^2} \sqrt{\mathbf{q}^2 + m_D^2}} \delta(\omega - \sqrt{\mathbf{q}^2 + m_K^2} - \sqrt{\mathbf{q}^2 + m_D^2})$$

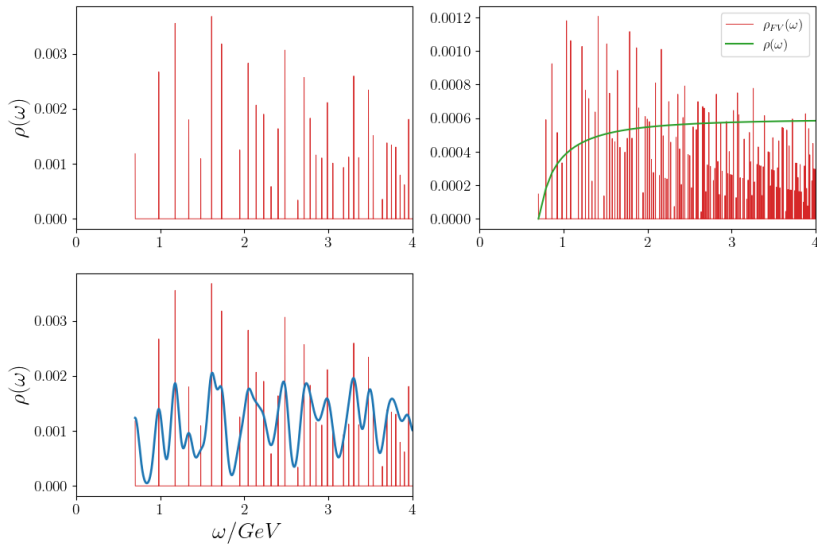
Smearing of $W_{\omega\omega}$



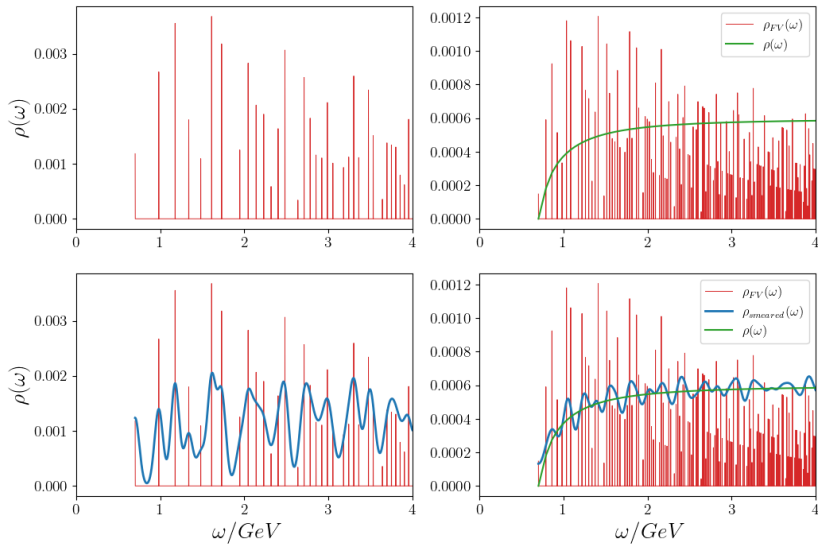
Smearing of $W_{\mu\nu}$



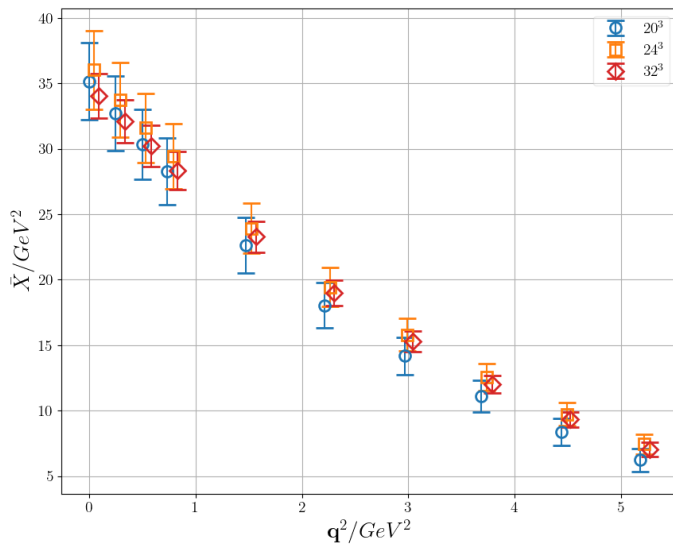
Smearing of W_{UV}



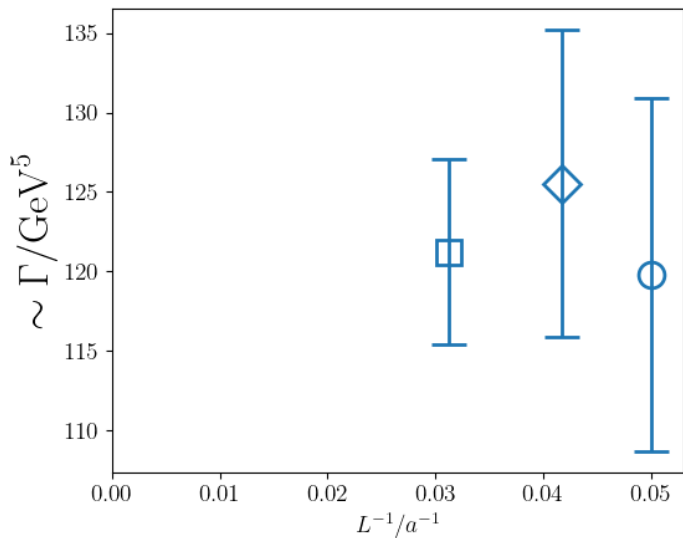
Smearing of W_{UV}



(1) Finite volume effects

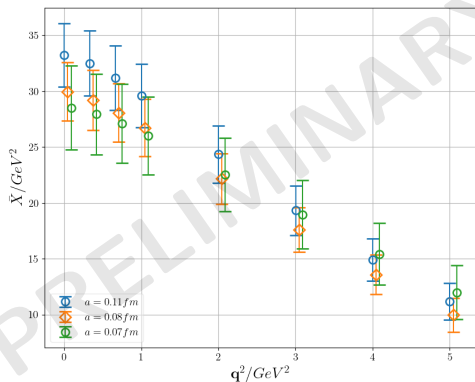


(1) Finite volume effects



(2) Continuum extrapolation - global fits

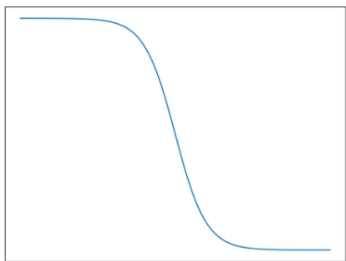
$$\text{fit} = \left(\sum_{i=0}^{N_p} c_i (\mathbf{q}^2)^i \right) \left(1 + c_{M_c} \left(\frac{M_{D_s}^{\text{phy}} - M_{D_s}}{\Lambda_c} \right) \right) \left(1 + c_a (a\Lambda_c)^2 \right) \left(1 + c_{a\mathbf{q}} (a\mathbf{q})^2 \right)$$




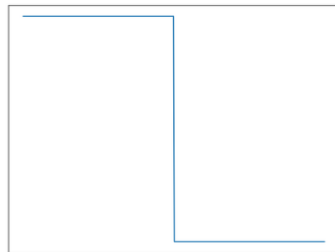
- ▶ Finest ensemble data still in production.
- ▶ Symanzik improvement not yet applied:

$$\mathcal{O}(a) \rightarrow \mathcal{O}(a^2)$$

Smearing limit



$$\sigma \rightarrow 0$$




(3) Error of \bar{X} with σ/N

[Kellermann et al., 2025, Barone et al., 2026]

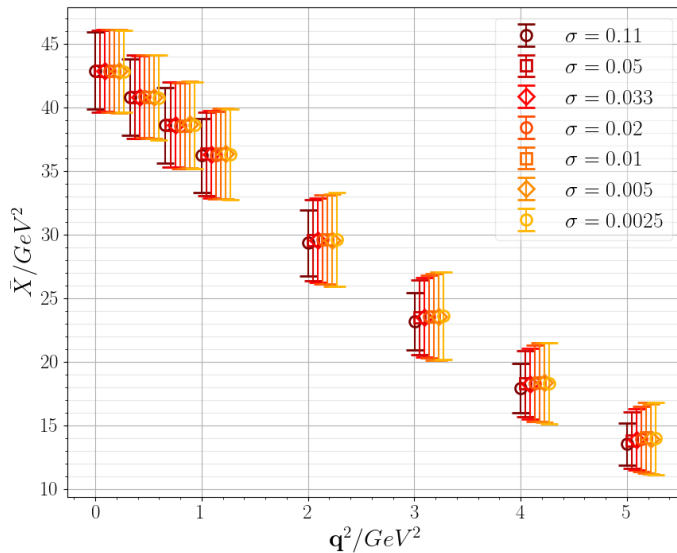
$$k^{\mu\nu}(\mathbf{q}, \omega)\theta(\omega_{\max} - \omega) \rightarrow k^{\mu\nu}(\mathbf{q}, \omega)\theta_{\sigma}(\omega_{\max} - \omega) \rightarrow \sum_{t=0}^N \tilde{c}_{\mu\nu, t}(\mathbf{q}, \sigma) T_t(e^{-\omega})$$

- ▶ Order of expansion = time separation of currents
- ▶ Maximum order is limited by the data.

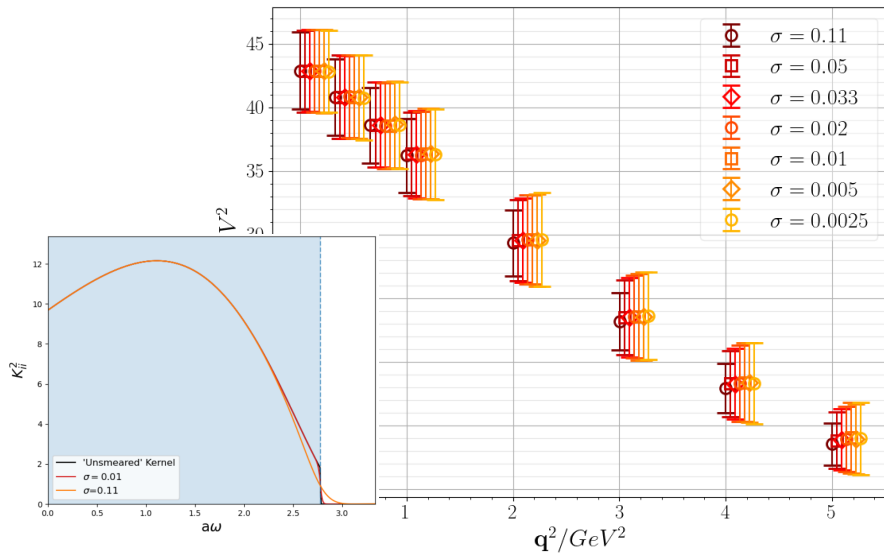
$$\bar{X} = \underbrace{\sum_t^{t_{\max}} \tilde{c}_{\mu\nu, t}(\sigma) \langle T_t \rangle_{\mu\nu}}_{\text{Signal}} + \sum_{t_{\max}+1}^N \tilde{c}_{\mu\nu, t}(\sigma) \cdot \text{Noise}$$

- ▶ Noise is bounded as $|\langle T_t \rangle_{\mu\nu}| \leq 1$.
- ▶ Random sample values for $\langle T_{k > t_{\max}} \rangle$ from $\{-1, 1\}$.
- ▶ Vary σ to see dependence of $\delta\bar{X}$.

(3) Smearing limit



(3) Smearing limit



Summary




- ▶ Explored this alternate calculation of inclusive decays (results for $D_{(s)}$ - [De Santis et al., 2025a, De Santis et al., 2025b, Kellermann et al., 2025, Kellermann et al., 2026])

Future Work





- ▶ Full continuum limit with systematics - Symanzik improvement, data production.
- ▶ Full phenomenological prediction with quantified errors (very close).

THANK YOU!




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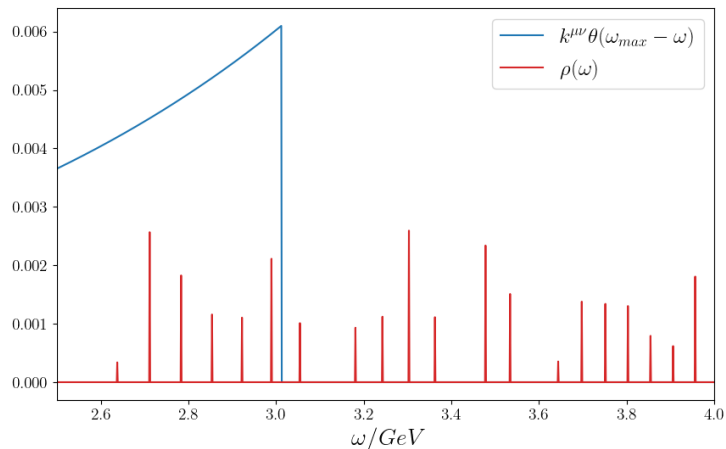
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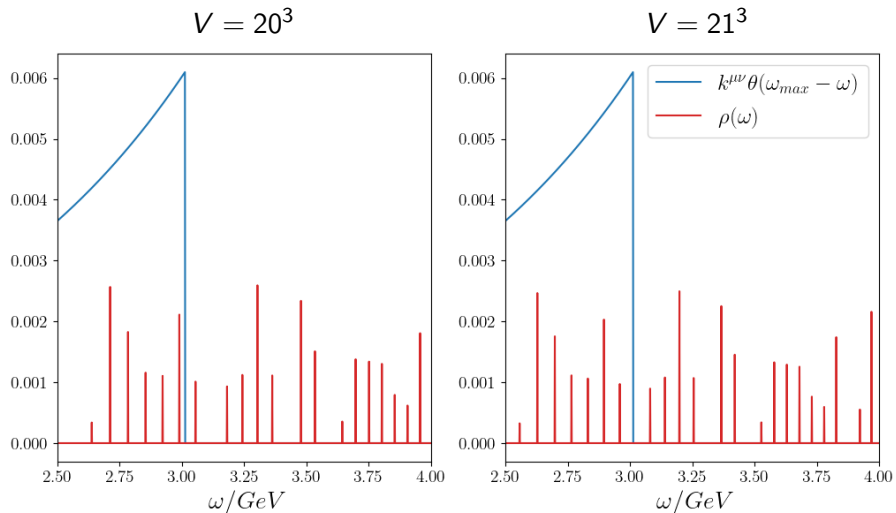
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Nonperturbatively determined relativistic heavy quark action.
Physical Review DParticles, Fields, Gravitation, and Cosmology, 76(7):074506.

How does \bar{X} behave?

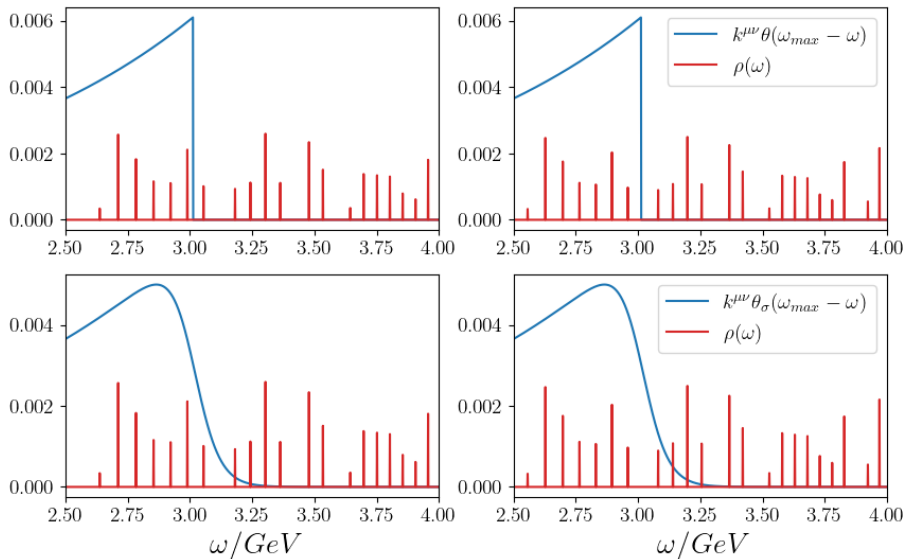
$$\bar{X}(\mathbf{q}) = \int_{\omega_0}^{\infty} d\omega W_{\mu\nu}(\mathbf{q}, \omega) k^{\mu\nu}(\mathbf{q}, \omega) \theta(\omega_{\max} - \omega)$$



How does \bar{X} behave?



(2) $V \rightarrow \infty$ and smearing



Relation between $C_{\mu\nu}$ and $W_{\mu\nu}$?

$$\frac{\langle O_{B_s}^S(t_{\text{snk}}) \tilde{J}_\mu^\dagger(\mathbf{q}, t_2) \tilde{J}_\nu(\mathbf{q}, t_1) O_{B_s}^{S\dagger}(t_{\text{src}}) \rangle}{\langle O_{B_s}^S(t_{\text{snk}}) O_{B_s}^{S\dagger}(t_2) \rangle \langle O_{B_s}^S(t_2) O_{B_s}^{S\dagger}(t_{\text{src}}) \rangle} = \frac{A_{LL}}{2M_{B_s}} \langle B_s | \tilde{J}_\mu^\dagger(\mathbf{q}, t_2) \tilde{J}_\nu(\mathbf{q}, t_1) | B_s \rangle$$

where we have

$$C_{\mu\nu}(\mathbf{q}, t) = \frac{1}{2M_{B_s}} \langle B_s | \tilde{J}_\mu^\dagger(\mathbf{q}, 0) e^{-\hat{H}t} \tilde{J}_\nu(\mathbf{q}, 0) | B_s \rangle$$

This object then relates to $W_{\mu\nu}$ via the Laplace transform.

$$\begin{aligned} C_{\mu\nu}(\mathbf{q}, t) &= \int_0^\infty d\omega \frac{1}{2M_{B_s}} \langle B_s | \tilde{J}_\mu^\dagger(\mathbf{q}, 0) \delta(\hat{H} - \omega) \tilde{J}_\nu(\mathbf{q}, 0) | B_s \rangle e^{-\omega t} \\ &= \int_0^\infty d\omega W_{\mu\nu}(\mathbf{q}, \omega) e^{-\omega t} \end{aligned}$$