



## DYNAMICAL MODELLING IN THE COSMIC WEB

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# MOTIVATION

- Halo assembly bias:
  - the dependence of the clustering of dark matter halos on physical parameters outside of halo mass. (e.g. halo age, shape, spin)
  - E.g. Sheth & Tormen04, Gao+05, Wechsler+06, Dalal+08
- Linked to the *large scale tidal environment* which is a direct effect of the anisotropy of the cosmic web!
  - (e.g. Hahn+09, Tojeiro+17, Kraljic+18)
- Found in various flavors of simulation but yet to converge in observations
  - Developing techniques to isolate assembly bias in observations is the aim here!

#### ANISOTROPY



*Red = material moving away* 

(Borzyszkowski+16)





**DEX (7/1/19)** 

### VELOCITY ANISOTROPY

PRELIMINARY

Stacking TNG300 haloes:  $10^{11.8}\text{--}10^{12.5}\,M_{\odot}$ 



DEX (7/1/19)

# DISCRETE JEANS

*CJAM : Jeans Anisotropic MGE modelling code written in C* Watkins+13, https://github.com/lauralwatkins/cjam Adaptation to include chemical properties: Zhu+16

- 1) Construct MGEs of tracer population + gravitational potential (or functional form)
- 2) Solve Jeans' equations taking *each* tracer position as a prior and compare model velocity to the observed (line-of-sight)
- 3) Use maximum likelihoods to find best fit free parameters + prediction for 3D velocities for all tracers!

# SUMMARY

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 Duckworth+18: No correlation between accretion rate seen by rotation offset between stars and gas with vicinity to filaments (or nodes!) for ~1000 MaNGA central galaxies

In progress:

- DisPerSE: Cosmic web identification in IllustrisTNG public catalogue to be released
- Discrete dynamical modelling applied to TNG stacked haloes
- Discrete dynamical modelling applied to large cluster in TNG
- Kinematic offsets as measures of merger rate